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Special Issue

Strongly Interacting Matter at Extreme Conditions—the Role of Symmetry Energy in Nuclear Physics and Astrophysics

Edited by

Prof. Dr. Odilon Lourenço, Dr. Helena Sofia Pais and Prof. Dr. Mariana Dutra



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Approaching the Conformal Limit of Quark Matter with Different Chemical Potentials

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Abstract: We study in detail the influence of different chemical potentials (baryon, electric charge, strange, and neutrino) on how and how fast a free gas of quarks in the zero-temperature limit reaches the conformal limit. We discuss the influence of non-zero masses, the inclusion of leptons, and different constraints, such as charge neutrality, zero-net strangeness, and fixed lepton fraction. We also investigate for the first time how the symmetry energy of the system under some of these conditions approaches the conformal limit. We find that the inclusion of all quark masses (even the light ones) can produce different results depending on the chemical potential values or constraints assumed. A positive or negative deviation of 10% from the pressure of free massless quarks with the same chemical potential was found to take place as low as $\mu_B = 77$ to as high as 48,897 MeV. This illustrates the fact that the “free” or conformal limit is not a unique description. Finally, we briefly discuss what kind of corrections are expected from perturbative QCD as one goes away from the conformal limit.

Keywords: conformal limit; quark matter; chemical potential; symmetry energy



Citation: Brown, C.; Dexheimer, V.; Jacobsen, R.B.; Farias, R.L.S. Approaching the Conformal Limit of Quark Matter with Different Chemical Potentials. *Symmetry* **2024**, *16*, 852. <https://doi.org/10.3390/sym16070852>

Academic Editor: Shusu Shi

Received: 2 May 2024

Revised: 21 June 2024

Accepted: 1 July 2024

Published: 5 July 2024



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1. Introduction and Formalism

In the zero temperature limit, baryons start to overlap at a few times the saturation density, and through some mechanism that is not yet understood, quarks become effectively deconfined [1]. In this work, we discuss dense matter in terms of baryon chemical potential μ_B instead of baryon (number) density n_B , as the former (together with other chemical potentials, such as electric charge μ_Q or strange μ_S) is the fixed or independent quantity in the grand canonical ensemble. The correspondence between n_B and μ_B is model-dependent, but at finite temperature, the μ_B at which deconfinement takes place is expected to be even lower (see, e.g., [2]), which highlights the importance of studying quark matter. We are particularly interested in understanding the conformal limit, the asymptotically high μ_B at which matter can be described by a free (non-interacting) gas of massless quarks. For this reason, in the present work, we focus on modeling quark matter only and, for the time being, restrict ourselves to the zero-temperature limit.

To describe the quarks, we make use of a free Fermi gas under different assumptions. To start, we describe them simply as a massless gas, then introduce different non-zero but constant quark masses, and vary the baryon, electric charge, and strange chemical potentials independently. We further link the chemical potentials by imposing charge neutrality and/or zero net strangeness. We also discuss the role played by leptons, including beta equilibrium and the role played by neutrinos (with chemical potential μ_ν). We investigate large μ_B and different μ_Q and μ_ν , as these are important for astrophysical scenarios, such as neutron stars and neutron-star mergers. On the other hand, we investigate the effects of μ_S , which is important for discussions related to relativistic heavy-ion collisions and the early universe [3].

We also discuss the symmetry energy of quark matter for some of the constraints we study and investigate how it changes as we approach the conformal limit. The so-called symmetry energy (which is really the *asymmetry energy*) is one of the most important features of nuclear physics in general since it is related to the ratio between the different components of the nuclear systems [4]. Several works have addressed the symmetry energy of quark matter [5–8]. This physical quantity is defined as the difference in energy per baryon E/N_B (or energy density per baryon density ε/n_B) between completely isospin-asymmetric matter $\delta = 1$ and isospin-symmetric matter $\delta = 0$:

$$E_{\text{sym}} = \frac{E_{\delta=1}}{N_B} - \frac{E_{\delta=0}}{N_B} = \frac{\varepsilon_{\delta=1}}{n_B} - \frac{\varepsilon_{\delta=0}}{n_B}, \quad (1)$$

where δ is originally defined for matter with neutrons and protons in terms of densities n_i as

$$\delta = \frac{n_n - n_p}{n_n + n_p}. \quad (2)$$

In this case and also when one is considering up and down quarks, δ can also be written as

$$\delta = -2Y_I = 1 - 2Y_Q, \quad (3)$$

for non-strange matter using the Gell–Mann–Nishijima formula [9] with fractions Y_I and Y_Q summing over i = baryons or quarks and defined in terms of particle isospin I_i and electric charge Q_i , respectively:

$$Y_I = \frac{\sum_i I_i n_i}{\sum_i n_i}, \quad Y_Q = \frac{\sum_i Q_i n_i}{\sum_i n_i}, \quad (4)$$

with baryon (number) density $n_B = \sum n_i$, where quark densities n_i are divided by 3.

However, it is important to note that, as discussed in Ref. [10] and Appendix A of Ref. [11], in the presence of hyperons (or in our case, strange quarks), Equation (3) does not apply. For this reason, we constrain ourselves to the discussion of symmetry energy for the 2-flavor case (with up and down quarks).

When leptons are included, we assume beta equilibrium, in which case electrons and muons have chemical potential $\mu_e = \mu_\mu = -\mu_Q$. In the special case that (electron and muon) neutrinos are trapped, μ_ν is determined by fixing the lepton fraction

$$Y_l = \frac{\sum_{lep} n_{lep}}{\sum_i n_i}, \quad (5)$$

usually held equal to the canonical value 0.4, to simulate conditions created in supernova explosions [12].

Finally, we briefly discuss the effects of interactions in the case that they are weak enough to be discussed perturbatively, i.e., using perturbative Quantum Chromodynamics (pQCD)). At large temperatures and/or quark chemical potentials, the strong coupling becomes small enough to allow an infinite number of terms to be approximated by a finite number of terms to describe interactions [13]. At zero temperature, QCD needs perturbative expansions and normalization group techniques. This problem was first addressed in the late 1970s [14–16], showing that already at the second order in α_s , log-singularities and non-trivial effects due to the renormalization scale appear. In the particular case of beta equilibrium, first-order corrections cancel out, leading to the very simple and popular description for neutron stars in terms of free quarks plus a constant bag correction [17]. This description is consistent with our work, up to a bag constant (usually fitted to phenomenology). But, even in the massless beta equilibrium case, it has been shown that the effect of interactions is not negligible in the density or chemical potential regime relevant for astrophysics [18].

At zero temperature, pQCD corrections have been calculated up to next-to-next-to-next-to-leading order (N^3LO) [19,20] (see Equation (42) of [21], where our free calculations correspond to the term of order zero in the coupling α_s). With non-zero quark masses, pQCD corrections have been calculated up to next-to-next-to-leading order (N^2LO) [22–25]. In this case, the QCD running coupling constant implies running non-zero masses at finite chemical potentials. See the recently published lecture notes of pQCD in the context of astrophysics for more details [21]. In this work, we do not study interactions but simply discuss different kinds of conformal limits (for free quarks) and quantify how different they are from each other. Nevertheless, even in this simplest scenario, as a first attempt to clarify, e.g., the influence of various chemical potentials on the conformal limit, our approach is complementary to the study of interactions.

2. Results

We describe in detail the free Fermi gas formalism we use in this work (for quarks and leptons) in Appendix A. We begin our discussion by ignoring the contribution of leptons to the thermodynamical quantities (later, we include different possibilities and discuss them). In the figures that follow, pressure P and baryon density n_B are normalized by respective values of a free gas with the same number of quark flavors included, but with quark masses $m_i = 0$ and $\mu_Q = \mu_S = 0$. Simple analytical equations for the pressure of all the massless cases discussed in this work are derived in Appendix B. We start our discussion considering only one chemical potential and then expand our discussion to two and three chemical potentials.

2.1. One Chemical Potential μ_B

We start by comparing the quark mass effect on n_B versus μ_B in the left upper panel of Figure 1. Because in this case μ_Q and μ_S are zero, all quarks present the same chemical potential $\mu_i = \mu_u = \mu_d = \mu_s = \frac{1}{3}\mu_B$. Because of our normalization (thermodynamical quantities divided by the massless case with the respective number of flavors), all massless cases have a constant value of one. Nevertheless, this does not mean that they are the same (if not normalized). To discuss the effect of quark masses, we start with one flavor with up and down quark masses $m = 2.3$ and $m = 4.8$, respectively, and then, we look at the two-flavor case with these masses for both light quarks. After that, we look at three flavors and use first only non-zero mass for the strange quark $m = 95$ MeV and then the masses for the three quarks. The quark masses we use correspond to the Particle Data Group (PDG [26], within the error bars of updated values [27]). From here on, we refer to these masses as “realistic”.

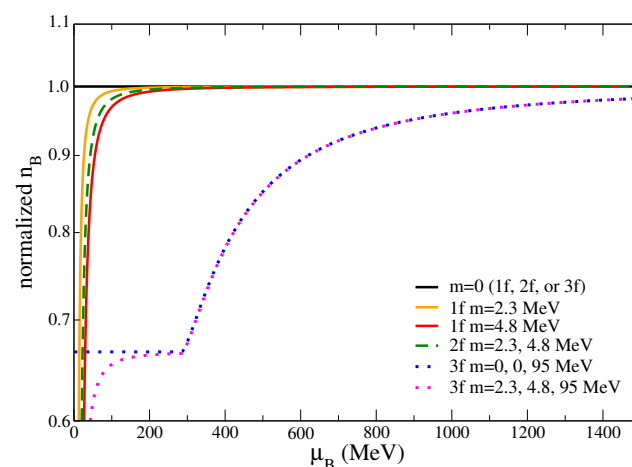


Figure 1. Cont.

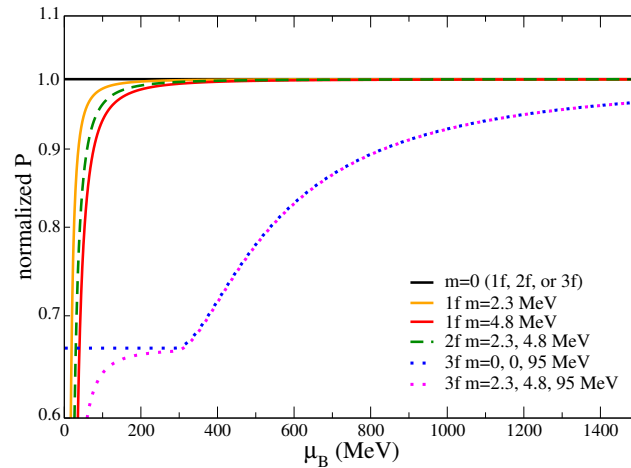


Figure 1. Baryon density (upper panel) and pressure (lower panel) of quarks with different numbers of flavors and different masses normalized by the respective massless cases.

We find that the introduction of realistic quark masses decreases the density for low μ_B , with the s-quark mass affecting the density until μ_B is larger (up to 621 MeV) than the two light quarks (up to 55 MeV). To calculate these thresholds, we use the criterium of a deviation of 10% from the black line with a value of one throughout this paper. For P versus μ_B , shown in the lower panel of Figure 1, the lines are very similar in shape to the ones in the upper panel of the figure. The introduction of realistic quark masses decreases again P for low μ_B , with the s-quark mass affecting the pressure until μ_B is larger (up to 834 MeV) than the two light quarks (up to 77 MeV).

2.2. Two Chemical Potentials μ_B and μ_Q

Now, we abandon the unphysical one-flavor case and continue with two- and three-flavor cases. The two-flavor case has recently become more relevant for dense matter because it has been shown that the core of neutron stars can harbor three-, as well as two-flavor quark matter [28]. For this case, we add another (electric charge) chemical potential, breaking some of the degeneracy in the quark chemical potentials: $\mu_{\text{up}} = \frac{1}{3}\mu_B + \frac{2}{3}\mu_Q$, $\mu_{\text{down}} = \mu_{\text{strange}} = \frac{1}{3}\mu_B - \frac{1}{3}\mu_Q$. Once more, we normalize thermodynamical quantities, dividing by the respective values of the same quantity for a free gas with the same number of quark flavors included, but with $m_i = 0$, in addition to $\mu_Q = 0$. Following this procedure, we aim to determine how the conformal limit and its deviation depend on μ_Q .

When μ_Q is determined by charge neutrality, the results, even for the massless case, depend on the number of flavors. In this case, only the three-flavor case is coincidentally equal to the $\mu_Q = 0$ case (see the explanation following Equations (A45) to (A48) in Appendix B). For two flavors, this is not the case, and the pressure is lower than in the $\mu_Q = 0$ case, establishing a new lower conformal limit (see upper panel of Figure 2). Expressions for the pressure for each particular chemical potential case (always keeping $m_i = 0$ for simplicity) can be found in Appendix B. Compare, e.g., Equations (A33) and (A42). When adding quark masses, μ_Q determined by charge neutrality lowers the pressure (in comparison with the respective massless case and the massless case with $\mu_Q = 0$) such that it goes to the respective conformal limit at larger μ_B . Using again the criteria of 10% deviations from the respective conformal limit, the s-quark mass affects pressure until $\mu_B = 839$ MeV and the two light quark masses until $\mu_B = 118$ MeV.

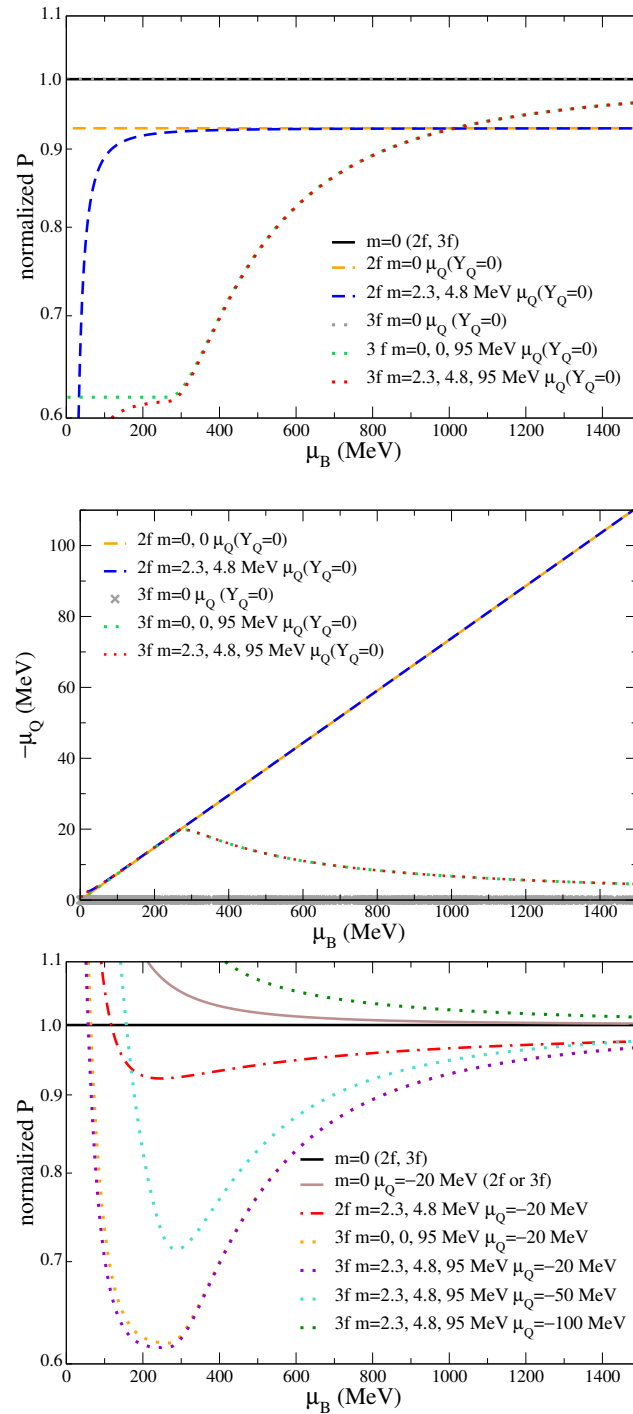


Figure 2. Pressure (**upper panel**) and electric charge chemical potential (**middle panel**) of quarks with two chemical potentials normalized by the respective massless case with one chemical potential, μ_B . The electric charge chemical potential is determined by charge neutrality. For massless three-flavor quarks, the cases with and without μ_Q coincide. (**Lower panel**): the pressure of quarks with two chemical potentials, being μ_Q fixed to different values, normalized by the respective massless case with one chemical potential, μ_B .

Nevertheless, one issue about this approach should be noted: we are comparing very small values of μ_Q with very large values of μ_B . See the middle panel of Figure 2 for a comparison. This is particularly the case for three flavors of quarks, and (except for extremely low μ_B) this behavior is independent of the quark masses. For small values of μ_B , both for two and three flavors, the dependence of μ_Q and μ_B can be predicted

in fair agreement with Equation (A41). For this reason, next, we add a fixed electric charge chemical potential to study how it affects the conformal limit, which translates into an increase in pressure (see, e.g., the different lines for three-flavor quark matter with realistic masses in the lower panel of Figure 2), especially at low values of μ_B . For massless quarks and $\mu_Q = -20$ MeV, the pressure is always above the conformal limit for $\mu_Q = 0$, independent of the number of flavors. Once the quark masses are finite, the pressure decreases, especially in the three-flavor case (but also for the two-flavor case). For larger absolute values of μ_Q , the pressure becomes larger, even going above the conformal case (with and without μ_Q). For example, for the three-flavor case with realistic quark masses and $\mu_Q = -50$ MeV, the pressure deviates by 10% (of the $\mu_Q = 0$ conformal limit) at $\mu_B = 698$ MeV and for $\mu_Q = -100$ MeV at $\mu_B = 415$ MeV (the latter one from above). Finally, there is one important remark regarding the behavior of the normalized pressure: in the lower panel of Figure 2, it is shown that this physical quantity decreases for small values of μ_B ; however, this behavior does not mean that the pressure itself (not normalized) is not a monotonically increasing function of μ_B . Here, we must remember that our normalization is carried out by dividing the thermodynamical quantities (such as pressure) by the massless case with the respective number of flavors and the free Fermi pressure of this system of massless quarks used for normalization scales as μ_B^4 ; therefore, in those ranges of μ_B where P for massive quarks increases at a lower rate than μ_B^4 , the normalized pressure decreases without implying any thermodynamical inconsistency.

2.3. Three Chemical Potentials μ_B , μ_Q , and μ_S or μ_ν

Going further, we can add another (strange) chemical potential and constrain it, e.g., to strangeness neutrality. The issue is that at zero temperature, strangeness neutrality means that there are no strange quarks, and the three-flavor case reduces to the two-flavor case. For this reason, we fix μ_S instead of specific values. μ_S breaks the degeneracy in the remaining quark chemical potentials: $\mu_{\text{up}} = \frac{1}{3}\mu_B + \frac{2}{3}\mu_Q$, $\mu_{\text{down}} = \frac{1}{3}\mu_B - \frac{1}{3}\mu_Q$, $\mu_{\text{strange}} = \frac{1}{3}\mu_B - \frac{1}{3}\mu_Q + \mu_S$. Once more, we normalize thermodynamical quantities by dividing them by the respective values of the same quantity for a free gas with the same number of quark flavors included, but with $m_i = 0$, in addition to $\mu_Q = 0$.

Fixing μ_S increases the pressure, similar to fixing μ_Q . Compare, for example, the massless three-flavor case in the upper panel in Figure 3 and lower panel in Figure 2 and note that the pressure for a given μ_B is now much higher. When quark masses are added, the similarity disappears because μ_S only affects the strange quarks, which do not appear for low values of μ_B unless the μ_S value is larger than the strange quark mass, which corresponds to our case of $\mu_S = 100$ MeV. For $\mu_S = 50$ and $\mu_S = 100$ MeV, the 10% deviation from the conformal limit takes place at $\mu_B = 1743$ and $\mu_B = 4227$ MeV, respectively (both from above).

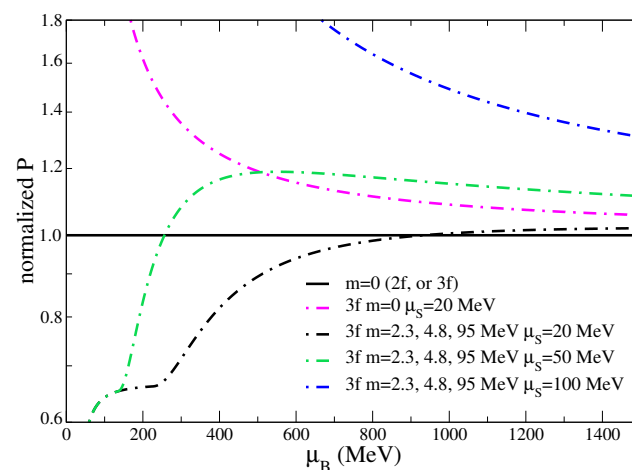


Figure 3. Cont.

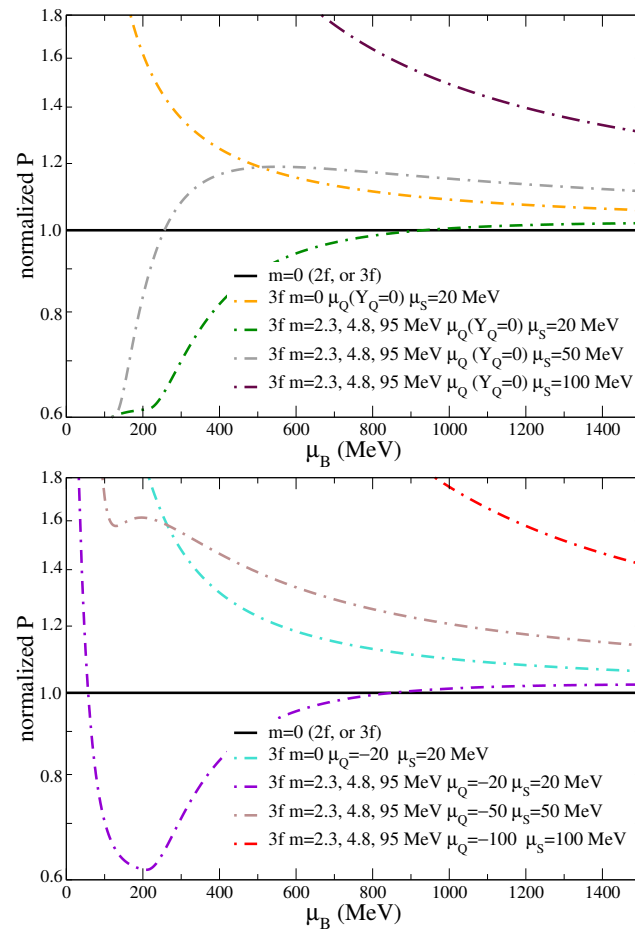


Figure 3. Pressure of quarks with two or three chemical potentials, including the strange chemical potential, normalized by the respective massless case (with one chemical potential, μ_B). The electric charge chemical potential is either zero (**upper panel**), determined by charge neutrality (**middle panel**), or fixed (**lower panel**). For massless three-flavor quarks, the cases with charge neutrality and without μ_Q coincide.

Now, we consider the additional case in which $\mu_Q \neq 0$, determined to reproduce charge neutrality (middle panel of Figure 3). For massless three-flavor quarks, the cases with charge neutrality and without μ_Q coincide. When masses are introduced, the curves are still very similar (to the upper panel for the $\mu_Q = 0$ case), except at very small μ_B , where the quark masses are comparable to both μ_B and μ_Q . For $\mu_S = 50$ and $\mu_S = 100$ MeV, the 10% deviation from the conformal limit takes place at $\mu_B = 1743$ and $\mu_B = 4227$ MeV, respectively (both from above). When a fixed value of μ_Q is used, it increases the pressure further, specifically at low μ_B (see lower panel of Figure 3). For $\mu_Q = \mu_S = 50$ and $\mu_Q = \mu_S = 100$ MeV, the 10% deviation from the conformal limit takes place at $\mu_B = 2070$ and $\mu_B = 4723$ MeV, respectively (both from above).

Next, we investigate the effects of having much larger values of μ_Q and μ_S , comparable to μ_B , for three flavors of quarks in the upper panel of Figure 4. As expected, the changes due to the additional chemical potentials take place at much lower μ_B (notice the different scale in the y-axis of the figure), and practically all the curves are above the one chemical potential (μ_B) conformal limit. An exception is the case with large (negative) μ_Q (and $\mu_S = 0$) because, according to Equations (A1) and (A21), quarks can only exist after a given $\mu_B = 381$ MeV, at which the momentum k_i and P become finite (see Equation (A44) for the massless case). In this case, the pressure differs from the one chemical potential conformal limit by more than 10% until $\mu_B = 10\,583$ MeV. In the case of large μ_S , quarks can exist at any μ_B , and the pressure differs from the one chemical potential conformal limit by more than 10% until $\mu_B = 44\,237$ MeV. When we combine large μ_S and (the absolute value of)

μ_Q , the pressure differs from the one chemical potential conformal limit by more than 10% until $\mu_B = 48,897$ MeV. In this case, the curve in the upper panel of Figure 4 begins only at $\mu_B = 1000$ MeV. This can be understood once more from Equations (A1) and (A21). The same effect can also be seen (although more subtle) in the bottom panel of Figure 2, where the fixed μ_Q cases start at $\mu_B = -\mu_Q$.

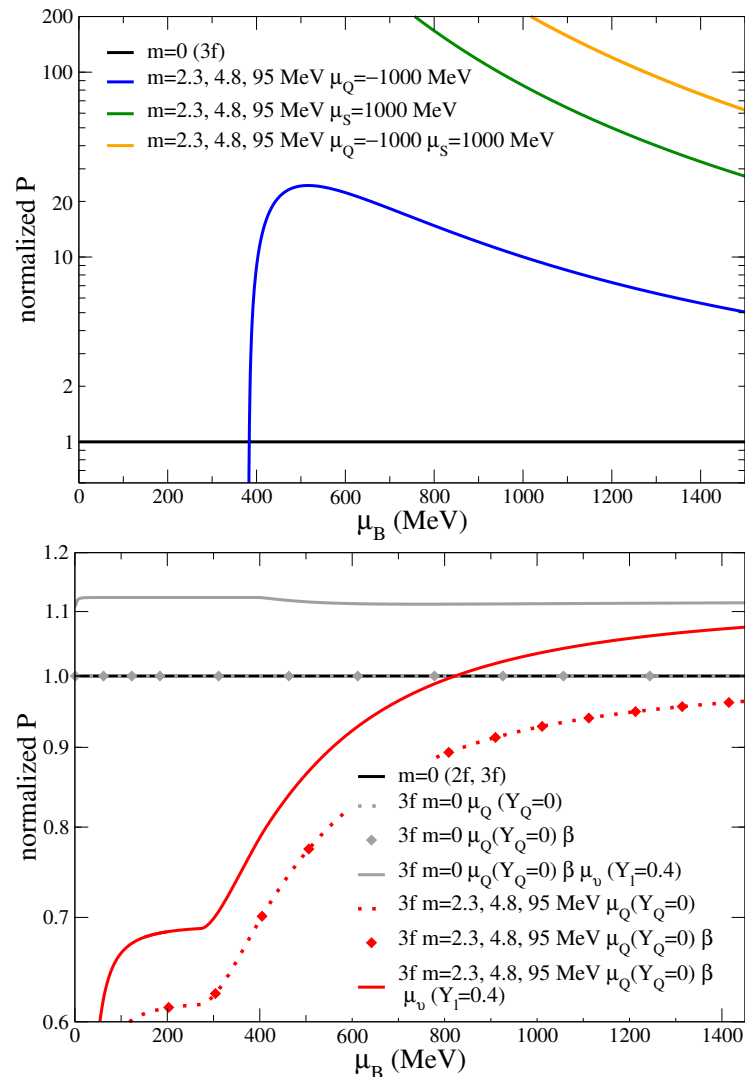


Figure 4. (Upper panel): Pressure of quarks with two or three large chemical potentials, normalized by the respective massless case with one chemical potential, μ_B . **(Lower panel):** The pressure of quarks and leptons with two or three chemical potentials, normalized by the respective massless case with one chemical potential, μ_B . For beta equilibrium with leptons, μ_Q is determined by charge neutrality. When neutrinos are present, their chemical potential μ_ν is determined by fixing the lepton fraction, Y_l .

Finally, we investigate changes due to the inclusion of a free gas of leptons (electrons and muons) in beta equilibrium (and participating in the fulfillment of charge neutrality). As can be seen in the lower panel of Figure 4, the inclusion of leptons (which appear in very small numbers or not at all) does not change the pressure. The picture changes, though, when the lepton fraction is fixed. In this case, which also includes neutrinos, the pressure is considerably higher, not because of the neutrinos themselves but because the larger number of negative leptons forces the appearance of a large number of up quarks, changing considerably the quark composition of the system and the stiffness of the equation of state [29]. This same stiffening occurs with nucleons in equilibrium with a fixed fraction of

leptons in the context of protoneutron stars and supernova explosions (see, e.g., [30,31]). The gray full line shows a kink for $\mu_B \sim 400$ MeV, when the muons appear. Note that the difference in massless versus massive quarks is still very pronounced when Y_I is fixed.

2.4. Symmetry Energy

As already discussed, we calculate the symmetry energy only for the two-flavor case, for which it was originally defined. The symmetry energy can be defined for strange matter, but the problem in this case is that it becomes ambiguous, as the two sides that appear in Equation (3) (in terms of Y_I or Y_Q) become different because of the Gell–Mann–Nishijima formula. Instead of choosing one particular definition for strange matter, we prefer not to use it. For more details regarding the treatment of the symmetry energy for strange matter, see, e.g., [5].

To perform the calculation, we fix n_B in this case (instead of μ_B as we have been doing) because the symmetry energy is defined for a given n_B but limit the x-axis to approximately the corresponding range from the previous figures. The upper panel of Figure 5 shows that the curves are a monotonically increasing function of density and that the light quark masses do not affect the results. Indeed, the effect of the mass is expected to be negligible in the symmetry energy of a free quark gas since the quark masses are taken as very small at any density, while physically, they should increase as density decreases. At very high density, much above the range shown in Figure 5, when the interactions are so weak that they can be neglected, the independence observed on the quark mass means that it is correct to consider the conformal (or massless) limit for the high density limit of QCD because massless quarks or quarks with physical masses of the order of the MeV are basically equivalent. This feature is reinforced by the clear overlapping of the equations of state $P(\epsilon)$ for the corresponding two-flavor cases here analyzed, as shown in the lower panel of Figure 5. In this plot, the range of μ_B is the same as in the previous figures (running from 0 to approximately 1400 MeV). Notice that the independence on the light quark masses applies to every thermodynamical quantity that is not normalized by the respective conformal limit (and does not include derivatives). Numerically, we define $\delta = 0$ as the two-flavor $\mu_Q = 0$ case (corresponding to the two-flavor lines in Figure 1) and $\delta = 1$ as the two-flavor $Y_Q = 0$ case (with $\mu_Q \neq 0$ corresponding to the two-flavor lines in the top and middle panels of Figure 2).

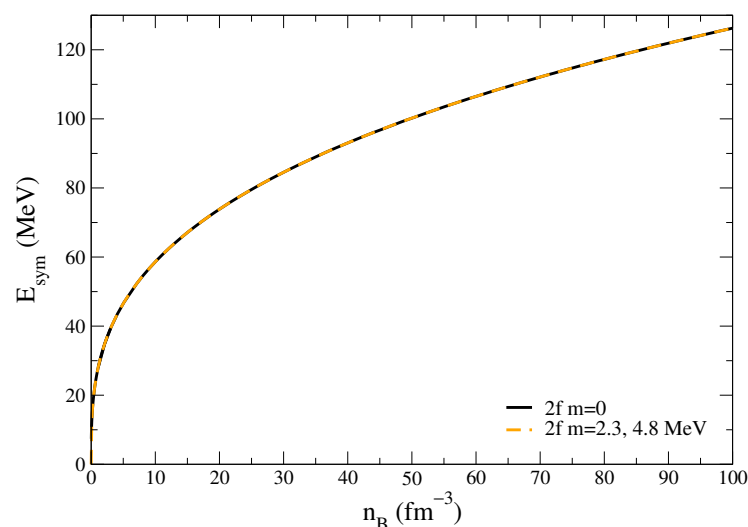


Figure 5. Cont.

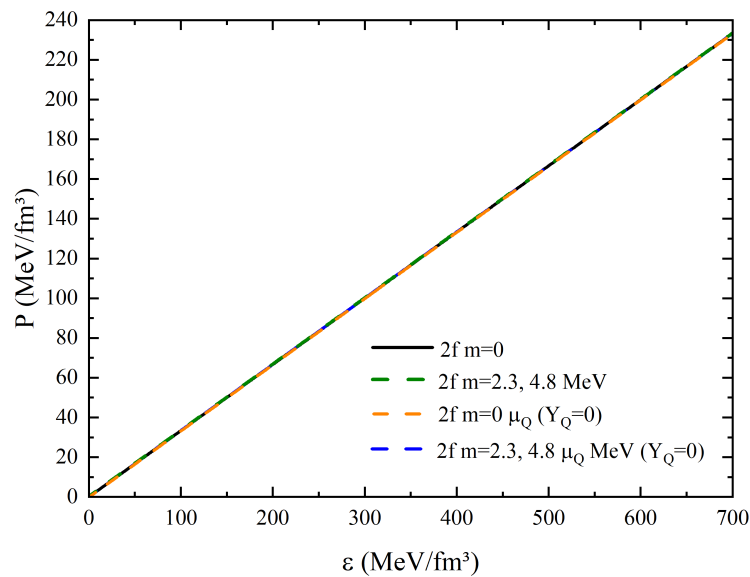


Figure 5. (Upper panel) Symmetry energy of two flavors of quarks as a function of baryon density for different masses. The two curves overlap. **(Lower panel)** Equation of state (pressure vs. energy density) for the two extreme cases subtracted in the panel above, isospin-symmetric matter $\delta = 0$ and charge neutral matter $\delta = 1$, shown for realistic or zero quark masses. In this scale, the four curves overlap.

3. Discussion and Conclusions

Perturbative corrections to a free gas of quarks due to interactions always bring down the pressure to lower values. Although these corrections have been calculated to higher orders for massless and massive (strange) quarks, they cannot directly be carried out to low baryon chemical potentials μ_B (or, interchangeably, low baryon densities n_B in the zero-temperature limit). This stems from the fact that matter in this regime is not perturbative, and at some point, a phase transition to hadronic matter takes place. As a result, for the relevant regime of densities inside neutron stars, $\mu_B \leq 1500$ MeV, pQCD predicts that the pressure is lower than 80% of the free gas value (see, for example, Figure 1 of Ref. [24]), but with a very large band going all the way to $P = 0$. It is important to clarify that this band is not an error bar but the envelope considering the different scenarios of strange quark mass proposed in [23] and a variation of the renormalization parameter $\bar{\Lambda}$ (see [22,32,33] for more details).

Note that pQCD figures are usually shown normalized by the same case, i.e., same massless quark and chemical potential conditions, only turning off the interactions (in Ref. [24], the effect of masses are highlighted, but the chemical potential conditions are kept the same). Their intention is to show the changes in the equation of state due to the interactions. Our approach is different and, in our view, complementary: we are only looking at free quarks, but we are normalizing our figures to different quark masses and, most importantly, chemical potential values or conditions. Our intention is to measure how much these masses and chemical potentials affect the equation of state.

To do so, in this work, we investigate the equation of state of a free gas of quarks, focusing on how the conformal limit is reached when different chemical potentials are varied and different constraints (e.g., for laboratory vs. astrophysics) are considered. This is achieved by using combinations of one, two, or three chemical potentials out of the four we consider, each related to a possible conserved quantity: baryon number B (μ_B), electric charge (μ_Q), strangeness (μ_S), and lepton number (μ_ν). We also derive expressions for

massless quarks under different conditions and use the proportionality between the results to illustrate our discussion.

We study the effects of using different quark masses (including PDG values), numbers of flavors, and different ways to fix the various chemical potentials considered. The latter procedure implies enforcing charge neutrality and, when leptons are included, beta equilibrium. When leptons (electrons, muons, and their respective neutrinos) are present, the pressure is not altered. An exception is the case in which the lepton fraction is fixed. For different cases, we quantify the deviation from the one-chemical-potential (massless) conformal limit by verifying at which μ_B the pressure deviates by more than 10%. This value varies from $\mu_B = 77$ to 48,897 MeV. Depending on the values of chemical potentials, e.g., μ_Q , even the light quark masses can become relevant at large μ_B . This shows that one must be careful about making statements concerning comparisons with “the” conformal limit.

Finally, we show that the behavior of the symmetry energy is monotonically increasing and does not depend on the light quark masses.

Author Contributions: Conceptualization, V.D.; methodology, V.D.; software, C.B. and V.D.; validation, R.B.J. and R.L.S.F.; formal analysis, V.D. and R.B.J.; investigation, C.B., V.D., R.B.J. and R.L.S.F.; data curation, V.D. and R.B.J.; writing—original draft preparation, C.B. and V.D.; writing—review and editing, R.B.J. and R.L.S.F.; visualization, C.B. and V.D.; supervision, V.D. and R.L.S.F.; project administration, V.D.; funding acquisition, V.D. and R.L.S.F. All authors have read and agreed to the published version of the manuscript.

Funding: This research was funded by the National Science Foundation, grants number PHY1748621, MUSES OAC2103680, and NP3M PHY-2116686 (V.D.). This work was partially supported by Conselho Nacional de Desenvolvimento Científico e Tecnológico (CNPq), grant number 312032/2023-4 (R.L.S.F.) and is also part of the project Instituto Nacional de Ciência e Tecnologia-Física Nuclear e Aplicações (INCT-FNA), grant number 464898/2014-5 (R.L.S.F.). It also received support from the Kent State University SURE program (C.B.).

Data Availability Statement: The raw data supporting the conclusions of this article will be made available by the authors on request.

Conflicts of Interest: The authors declare no conflicts of interest.

Abbreviations

The following abbreviations are used in this manuscript: perturbative Quantum Chromodynamics (pQCD), Particle Data Group (PDG).

Appendix A. General Expressions

For each particle i , we can write its chemical potential as the combination of the independent chemical potentials of the system (each associated with a conserved quantity) weighted by the respective particle quantum number. In our case, of conserved baryon number, electric charge, and strangeness, we have

$$\mu_i = Q_B \mu_B + Q_i \mu_Q + Q_{S_i} \mu_S, \quad (\text{A1})$$

where the baryon number for quarks is $Q_B = 1/3$ and Q_i and Q_{S_i} are the electric charge and strangeness of each quark, respectively. Here, μ_B , μ_Q , and μ_S are the baryon, electric charge, and strange chemical potentials of the system, respectively. In our formalism, the isospin chemical potential $\mu_I = \mu_Q$ [10].

Alternatively, Equation (A1) can be derived from the principles of thermodynamics. One may start considering the additive property of the internal energy U for an n -component system:

$$U(\lambda S, \lambda V, \lambda N_1, \dots, \lambda N_n) = \lambda U(S, V, N_1, \dots, N_n), \quad (\text{A2})$$

where λ is arbitrary and S , V , and N_i denote the entropy, the volume, and the number of particles of a given component i , respectively. Let us differentiate this “extensivity condition” with respect to λ :

$$\begin{aligned} \frac{\partial U(\lambda S, \dots)}{\partial(\lambda S)} \frac{\partial(\lambda S)}{\partial \lambda} + \frac{\partial U(\lambda V, \dots)}{\partial(\lambda V)} \frac{\partial(\lambda V)}{\partial \lambda} + \frac{\partial U(\lambda N_1, \dots)}{\partial(\lambda N_1)} \frac{\partial(\lambda N_1)}{\partial \lambda} + \dots \\ + \frac{\partial U(\lambda N_n, \dots)}{\partial(\lambda N_n)} \frac{\partial(\lambda N_n)}{\partial \lambda} = U(S, V, N_1, \dots, N_n). \end{aligned} \quad (\text{A3})$$

Setting $\lambda = 1$ in the above equation, we obtain

$$\frac{\partial U}{\partial S} S + \frac{\partial U}{\partial V} V + \frac{\partial U}{\partial N_1} N_1 + \dots + \frac{\partial U}{\partial N_n} N_n = U. \quad (\text{A4})$$

Now, using the definition of the intensive parameters T (temperature), P (pressure), and μ_i (chemical potential), we arrive at the Euler equation:

$$U = TS - PV + \mu_1 N_1 + \dots + \mu_n N_n. \quad (\text{A5})$$

Inserting the above equations into the expression for the Gibbs energy $G = U + PV - TS$, we have

$$G = TS - PV + \mu_1 N_1 + \dots + \mu_n N_n + PV - TS \implies G = \sum_{i=1}^n \mu_i N_i, \quad (\text{A6})$$

which is known as the Gibbs–Duhem relation. In the condition of chemical equilibrium, the Gibbs energy must be minimized with respect to one of the quantities N_j . For constant temperature and pressure, this condition reads as

$$\sum_{i=1}^n \left(\frac{\partial G}{\partial N_i} \right)_{T, P, N_{i \neq j}} \frac{dN_i}{dN_j} = 0. \quad (\text{A7})$$

Now, let η_j stand for the coefficient that gives the proportion of the component j with respect to the other components of the system. If the component j suffers a variation $dN_j = \bar{\eta}_j$, all the other components must also have a variation given by $dN_i = (\bar{\eta}_j / \eta_j) \eta_i$ in order to keep the balance between the components implied by the condition of chemical equilibrium. Therefore, $dN_i / dN_j = \eta_i / \eta_j$. Additionally, according to Equation (A6),

$$\mu_i = \left(\frac{\partial G}{\partial N_i} \right)_{T, P, N_{i \neq j}}. \quad (\text{A8})$$

As a result, Equation (A7) can be written as

$$\sum_{i=1}^n \eta_i \mu_i = 0. \quad (\text{A9})$$

Considering that baryon number Q_B , electric charge Q , and strangeness Q_S are conserved quantities, the three conservation laws can be respectively written as

$$\sum_{i=1}^n \eta_i Q_{B_i} = 0, \quad \sum_{i=1}^n \eta_i Q_i = 0 \quad \text{and} \quad \sum_{i=1}^n \eta_i Q_{S_i} = 0. \quad (\text{A10})$$

As long as we have n variables and three equations, it is possible to write three of the η_i as functions of the other $n - 3$, as follows:

$$\eta_1 Q_{B_1} + \eta_2 Q_{B_2} + \eta_3 Q_{B_3} = - \sum_{i \neq 1, 2, 3}^n \eta_i Q_{B_i}, \quad (\text{A11})$$

$$\eta_1 Q_1 + \eta_2 Q_2 + \eta_3 Q_3 = - \sum_{i \neq 1,2,3}^n \eta_i Q_i, \quad (\text{A12})$$

$$\eta_1 Q_{S_1} + \eta_2 Q_{S_2} + \eta_3 Q_{S_3} = - \sum_{i \neq 1,2,3}^n \eta_i Q_{S_i}. \quad (\text{A13})$$

Clearly, the three independent components chosen to construct the above relations are completely arbitrary. As a consequence, we may consider a certain species 1, such that $Q_{B_1} = 1$, $Q_1 = 0$, and $Q_{S_1} = 0$; a certain species 2, such that $Q_{B_2} = 0$, $Q_2 = 1$, and $Q_{S_2} = 0$; and a certain species 3, such that $Q_{B_3} = 0$, $Q_3 = 0$, and $Q_{S_3} = 1$. In this case, the above equations are simplified to

$$\eta_1 = - \sum_{i \neq 1,2,3}^n \eta_i Q_{B_i}, \quad \eta_2 = - \sum_{i \neq 1,2,3}^n \eta_i Q_i \quad \text{and} \quad \eta_3 = - \sum_{i \neq 1,2,3}^n \eta_i Q_{S_i}. \quad (\text{A14})$$

Plugging Equation (A14) into Equation (A9), we find

$$\sum_{i \neq 1,2,3}^n \eta_i \mu_i = \sum_{i \neq 1,2,3}^n \eta_i (Q_{B_i} \mu_1) + \sum_{i \neq 1,2,3}^n \eta_i (Q_i \mu_2) + \sum_{i \neq 1,2,3}^n \eta_i (Q_{S_i} \mu_3). \quad (\text{A15})$$

Defining $\mu_1 = \mu_B$ (the baryon chemical potential), $\mu_2 = \mu_Q$ (the electric charge chemical potential), and $\mu_3 = \mu_S$ (the strange chemical potential), the above equation can be rewritten as

$$\sum_{i \neq 1,2,3}^n \eta_i \mu_i = \sum_{i \neq 1,2,3}^n \eta_i (Q_{B_i} \mu_B) + \sum_{i \neq 1,2,3}^n \eta_i (Q_i \mu_Q) + \sum_{i \neq 1,2,3}^n \eta_i (Q_{S_i} \mu_S). \quad (\text{A16})$$

Finally, since all the factors η_i are independent, the above equation only holds if the coefficients are equal, i.e.,

$$\mu_i = Q_{B_i} \mu_B + Q_i \mu_Q + Q_{S_i} \mu_S, \quad (\text{A17})$$

which precisely corresponds to Equation (A1) if we consider that all quarks have an identical baryon number Q_B , such that $Q_{B_i} = Q_B = 1/3$.

The general expressions for energy density and pressure of a relativistic free Fermi gas of particles i can be derived from the Dirac Lagrangian density extracting the diagonal components of the energy-momentum tensor (assuming an ideal fluid). The (number) density is simply the integral of the distribution function. Using the natural system of units, they are

$$n_i = \frac{g_i}{2\pi^2} \int_0^\infty dk_i k_i^2 (f_{i+} - f_{i-}), \quad (\text{A18})$$

$$\varepsilon_i = \frac{g_i}{2\pi^2} \int_0^\infty dk_i E_i k_i^2 (f_{i+} + f_{i-}), \quad (\text{A19})$$

$$P_i = \frac{1}{3} \frac{g_i}{2\pi^2} \int_0^\infty dk_i \frac{k_i^4}{E_i} (f_{i+} + f_{i-}), \quad (\text{A20})$$

where $g_i = 6$ is the spin and color degeneracy factor; k_i is the momentum;

$$E_i = \sqrt{k_i^2 + m_i^2} \geq 0 \quad (\text{A21})$$

is the energy of the state; m_i the mass; f_{\pm} is the distribution function of particles and antiparticles $f_{i\pm} = (e^{(E_i \mp \mu_i)/T} + 1)^{-1}$, with μ_i being the particle chemical potential; and T the temperature.

In the $T = 0$ limit, antiparticles provide no contribution; $f_- = 0$ and $f_+ = 1$ up to the Fermi momentum; $k_i = k_{F_i}$, $E_i = \mu_i$; and the integrals for the above thermodynamic quantities are evaluated analytically:

$$n_i = \frac{g_i}{6\pi^2} k_{F_i}^3, \quad (\text{A22})$$

$$\varepsilon_i = \frac{g_i}{2\pi^2} \left[\left(\frac{1}{8} m_i^2 k_{F_i} + \frac{1}{4} k_{F_i}^3 \right) \sqrt{m_i^2 + k_{F_i}^2} - \frac{1}{8} m_i^4 \ln \frac{k_{F_i} + \sqrt{m_i^2 + k_{F_i}^2}}{m_i} \right], \quad (\text{A23})$$

$$P_i = \frac{1}{3} \frac{g_i}{2\pi^2} \left[\left(\frac{1}{4} k_{F_i}^3 - \frac{3}{8} m_i^2 k_{F_i} \right) \sqrt{m_i^2 + k_{F_i}^2} + \frac{3}{8} m_i^4 \ln \frac{k_{F_i} + \sqrt{m_i^2 + k_{F_i}^2}}{m_i} \right]. \quad (\text{A24})$$

Appendix B. Massless Quarks

For the massless particle case, the expressions above further reduce to

$$n_i = \frac{g_i}{6\pi^2} k_{F_i}^3 = \frac{g_i}{6\pi^2} \mu_i^3, \quad (\text{A25})$$

$$\varepsilon_i = \frac{g_i}{8\pi^2} k_{F_i}^4 = \frac{g_i}{8\pi^2} \mu_i^4, \quad (\text{A26})$$

$$P_i = \frac{1}{3} \frac{g_i}{8\pi^2} k_{F_i}^4 = \frac{1}{3} \frac{g_i}{8\pi^2} \mu_i^4, \quad (\text{A27})$$

reproducing $\varepsilon_i = 3P_i$.

Note that in the case of massless free quarks, we can also write $\mu_i = k_i$. Therefore, we can write the chemical potential for each quark flavor using Equation (A1):

$$\mu_u = \frac{1}{3} \mu_B + \frac{2}{3} \mu_Q = k_u, \quad (\text{A28})$$

$$\mu_d = \frac{1}{3} \mu_B - \frac{1}{3} \mu_Q = k_d, \quad (\text{A29})$$

$$\mu_s = \frac{1}{3} \mu_B - \frac{1}{3} \mu_Q + \mu_S = k_s. \quad (\text{A30})$$

We use the convention that both the strangeness and μ_S are positive. Alternatively, one can use both as negative without changing the results. Equations (A28) and (A29) are equal if $\mu_Q = 0$. Equations (A28)–(A30) are equal if $\mu_Q = 0$ and $\mu_S = 0$. The density and pressure of each quark flavor can be written further as

$$n_i = \frac{\mu_i^3}{\pi^2} = \frac{k_i^3}{\pi^2}, \quad (\text{A31})$$

$$P_i = \frac{\mu_i^4}{4\pi^2} = \frac{k_i^4}{4\pi^2}. \quad (\text{A32})$$

Next, we discuss the pressure for specific conditions concerning the number of flavors and chemical potential constraints (not including leptons):

- Two-flavor, $\mu_Q = 0$:

$$P = P_u + P_d = 2P_u = \frac{2\mu_u^4}{4\pi^2} = \frac{\mu_B^4}{162\pi^2} = \frac{\mu_B^4}{1598.88}. \quad (\text{A33})$$

- Three-flavor, $\mu_Q = 0$, $\mu_S = 0$:

$$P = P_u + P_d + P_s = 3P_u = \frac{3\mu_u^4}{4\pi^2} = \frac{\mu_B^4}{108\pi^2} = \frac{\mu_B^4}{1065.92}. \quad (\text{A34})$$

- Two-flavor, μ_Q fixed:

$$\begin{aligned}
P = P_u + P_d &= \frac{1}{4\pi^2} (\mu_u^4 + \mu_d^4) = \frac{1}{4\pi^2} \left[\left(\frac{1}{3}\mu_B + \frac{2}{3}\mu_Q \right)^4 + \left(\frac{1}{3}\mu_B - \frac{1}{3}\mu_Q \right)^4 \right] \\
&= \frac{1}{4\pi^2} \left(\frac{\mu_B^4}{81} + \frac{4\mu_B^3}{27} \frac{2\mu_Q}{3} + \frac{6\mu_B^2}{9} \frac{4\mu_Q^2}{9} + \frac{4\mu_B}{3} \frac{8\mu_Q^3}{27} + \frac{16\mu_Q^4}{81} \right. \\
&\quad \left. + \frac{\mu_B^4}{81} - \frac{4\mu_B^3}{27} \frac{\mu_Q}{3} + \frac{6\mu_B^2}{9} \frac{\mu_Q^2}{9} - \frac{4\mu_B}{3} \frac{\mu_Q^3}{27} + \frac{\mu_Q^4}{81} \right) \\
&= \frac{1}{324\pi^2} [2\mu_B^4 + 4\mu_B^3\mu_Q + 30\mu_B^2\mu_Q^2 + 28\mu_B\mu_Q^3 + 17\mu_Q^4] . \quad (\text{A35})
\end{aligned}$$

- Two-flavor, μ_Q from charge neutrality:

Starting from $\sum_i Q_i n_i = 0$,

$$\frac{2}{3}n_u - \frac{1}{3}n_d = 0, \quad (\text{A36})$$

$$\frac{2}{3} \frac{\mu_u^3}{\pi^2} - \frac{1}{3} \frac{\mu_d^3}{\pi^2} = 0, \quad (\text{A37})$$

$$2\mu_u^3 = \mu_d^3, \quad (\text{A38})$$

$$2 \left(\frac{1}{3}\mu_B + \frac{2}{3}\mu_Q \right)^3 = \left(\frac{1}{3}\mu_B - \frac{1}{3}\mu_Q \right)^3, \quad (\text{A39})$$

$$2^{\frac{1}{3}} \frac{1}{3}\mu_B - \frac{1}{3}\mu_B = -2^{\frac{1}{3}} \frac{2}{3}\mu_Q - \frac{1}{3}\mu_Q, \quad (\text{A40})$$

$$\mu_Q = \frac{-\left(2^{\frac{1}{3}} - 1\right)\mu_B}{2^{\frac{4}{3}} + 1} = -0.0738 \mu_B. \quad (\text{A41})$$

We can then use Equations (A38) and (A41) to calculate the pressure:

$$\begin{aligned}
P &= P_u + P_d = \frac{1}{4\pi^2} (\mu_u^4 + \mu_d^4) = \frac{1}{4\pi^2} (\mu_u^4 + 2^{\frac{4}{3}}\mu_u^4) \\
&= \frac{1}{4\pi^2} (1 + 2^{\frac{4}{3}}) \mu_u^4 = \frac{1}{4\pi^2} (1 + 2^{\frac{4}{3}}) \left(\frac{1}{3}\mu_B + \frac{2}{3}\mu_Q \right)^4 \\
&= \frac{1}{4\pi^2} (1 + 2^{\frac{4}{3}}) \left[\frac{1}{3}\mu_B - \frac{2}{3} \left(\frac{2^{\frac{1}{3}} - 1}{2^{\frac{4}{3}} + 1} \mu_B \right) \right]^4 \\
&= \frac{1}{4\pi^2} (1 + 2^{\frac{4}{3}}) \left[\frac{2^{\frac{4}{3}} + 1 - 2^{\frac{4}{3}} + 2}{3(2^{\frac{4}{3}} + 1)} \right]^4 \mu_B^4 \\
&= \frac{1}{4\pi^2 (2^{\frac{4}{3}} + 1)^3} \mu_B^4 = \frac{\mu_B^4}{1721.59}. \quad (\text{A42})
\end{aligned}$$

- Three-flavor, μ_Q fixed, $\mu_S = 0$:

$$P = P_u + P_d + P_s = \frac{1}{4\pi^2} (\mu_u^4 + \mu_d^4 + \mu_s^4) = \frac{1}{4\pi^2} (\mu_u^4 + 2\mu_d^4) \quad (\text{A43})$$

because $\mu_d = \mu_s$ are equal, resulting in

$$\begin{aligned}
 P &= \frac{1}{4\pi^2} \left[\left(\frac{1}{3}\mu_B + \frac{2}{3}\mu_Q \right)^4 + 2 \left(\frac{1}{3}\mu_B - \frac{1}{3}\mu_Q \right)^4 \right] \\
 &= \frac{1}{4\pi^2} \left(\frac{\mu_B^4}{81} + \frac{4\mu_B^3}{27} \frac{2\mu_Q}{3} + \frac{6\mu_B^2}{9} \frac{4\mu_Q^2}{9} + \frac{4\mu_B}{3} \frac{8\mu_Q^3}{27} + \frac{16\mu_Q^4}{81} \right. \\
 &\quad \left. + 2 \frac{\mu_B^4}{81} - 2 \frac{4\mu_B^3}{27} \frac{\mu_Q}{3} + 2 \frac{6\mu_B^2}{9} \frac{\mu_Q^2}{9} - 2 \frac{4\mu_B}{3} \frac{\mu_Q^3}{27} + 2 \frac{\mu_Q^4}{81} \right) \\
 &= \frac{1}{324\pi^2} (3\mu_B^4 + 36\mu_B^2\mu_Q^2 + 24\mu_B\mu_Q^3 + 18\mu_Q^4). \quad (\text{A44})
 \end{aligned}$$

- Three-flavor, μ_Q from charge neutrality, $\mu_S = 0$:

Starting again from $\sum_i Q_i n_i = 0$,

$$\frac{2}{3}n_u - \frac{1}{3}n_d - \frac{1}{3}n_s = 0, \quad (\text{A45})$$

$$\frac{2}{3} \left(\frac{\mu_u^3}{\pi^2} \right) - \frac{1}{3} \left(\frac{\mu_d^3}{\pi^2} \right) \frac{1}{3} - \left(\frac{\mu_s^3}{\pi^2} \right) = 0, \quad (\text{A46})$$

$$2\mu_u^3 - \mu_d^3 - \mu_s^3 = 0, \quad (\text{A47})$$

but since in this case $\mu_d = \mu_s$, we have

$$\mu_u^3 = \mu_d^3, \quad (\text{A48})$$

which implies (from Equations (A28) and (A29)) $\mu_Q = 0$ and reproduces the three-flavor case with $\mu_Q = 0$, $\mu_S = 0$.

- Three-flavor, zero net strangeness:

Starting from $\sum Q_{S_i} n_i = 0$, at $T = 0$, this implies $n_s = 0$, no matter if $\mu_Q = 0$ or $\mu_Q \neq 0$. As a consequence, this case reproduces the respective two-flavor case.

- Three-flavor, μ_Q fixed, μ_S fixed:

$$\begin{aligned}
 P &= P_u + P_d + P_s = \frac{1}{4\pi^2} (\mu_u^4 + \mu_d^4 + \mu_s^4) \\
 &= \frac{1}{4\pi^2} \left[\left(\frac{1}{3}\mu_B + \frac{2}{3}\mu_Q \right)^4 + \left(\frac{1}{3}\mu_B - \frac{1}{3}\mu_Q \right)^4 + \left(\frac{1}{3}\mu_B - \frac{1}{3}\mu_Q + \mu_S \right)^4 \right]. \quad (\text{A49})
 \end{aligned}$$

Using the result from Equation (A44),

$$\begin{aligned}
 P &= \frac{1}{324\pi^2} (3\mu_B^4 + 36\mu_B^2\mu_Q^2 + 24\mu_B\mu_Q^3 + 18\mu_Q^4) \\
 &\quad + \frac{1}{4\pi^2} \left(\mu_S^4 - \frac{4}{27}\mu_Q^3\mu_S - \frac{4}{3}\mu_Q\mu_S^3 + \frac{4}{3}\mu_B\mu_S^3 + \frac{4}{27}\mu_B^3\mu_S \right. \\
 &\quad \left. + \frac{6}{9}\mu_Q^2\mu_S^2 + \frac{6}{9}\mu_B^2\mu_S^2 - \frac{12}{27}\mu_B^2\mu_Q\mu_S + \frac{12}{27}\mu_B\mu_Q^2\mu_S - \frac{12}{9}\mu_B\mu_Q\mu_S^2 \right). \quad (\text{A50})
 \end{aligned}$$

- Three-flavor $\mu_Q = 0$, μ_S fixed:

Using Equation (A50) with $\mu_Q = 0$,

$$P = \frac{1}{\pi^2} \left(\frac{1}{108} \mu_B^4 + \frac{\mu_S^4}{4} + \frac{1}{3} \mu_B \mu_S^3 + \frac{1}{27} \mu_B^3 \mu_S + \frac{1}{6} \mu_B^2 \mu_S^2 \right). \quad (\text{A51})$$

- Three-flavor, μ_Q from charge neutrality, μ_S fixed:

Starting from $\sum Q_i n_i = 0$,

$$\frac{2}{3} n_u - \frac{1}{3} n_d - \frac{1}{3} n_s = 0, \quad (\text{A52})$$

$$2\mu_u^3 - \mu_d^3 - \mu_s^3 = 0, \quad (\text{A53})$$

$$2 \left(\frac{1}{3} \mu_B + \frac{2}{3} \mu_Q \right)^3 - \left(\frac{1}{3} \mu_B - \frac{1}{3} \mu_Q \right)^3 - \left(\frac{1}{3} \mu_B - \frac{1}{3} \mu_Q + \mu_S \right)^3 = 0, \quad (\text{A54})$$

$$\begin{aligned} & \frac{2\mu_B^3}{27} + \frac{12\mu_B^2\mu_Q}{27} + \frac{24\mu_B\mu_Q^2}{27} + \frac{16\mu_Q^3}{27} - \frac{2\mu_B^3}{27} + \frac{6\mu_B^2\mu_Q}{27} - \frac{6\mu_B\mu_Q^2}{27} \\ & + \frac{2\mu_Q^3}{27} - \mu_S^3 - \frac{3\mu_B^2\mu_S}{9} + \frac{6\mu_B\mu_Q\mu_S}{9} - \frac{3\mu_Q^2\mu_S}{9} - \frac{3\mu_B\mu_S^2}{3} + \frac{3}{3}\mu_Q\mu_S^2 = 0, \end{aligned} \quad (\text{A55})$$

$$\frac{2\mu_B^2\mu_Q}{3} + \frac{2\mu_B\mu_Q^2}{3} + \frac{2\mu_Q^3}{3} - \mu_S^3 - \frac{\mu_B^2\mu_S}{3} + \frac{2\mu_B\mu_Q\mu_S}{3} - \frac{\mu_Q^2\mu_S}{3} - \mu_B\mu_S^2 + \mu_Q\mu_S^2 = 0. \quad (\text{A56})$$

In the above expression, we still need to isolate μ_Q and replace in Equation (A50).

References

1. Baym, G.; Hatsuda, T.; Kojo, T.; Powell, P.D.; Song, Y.; Takatsuka, T. From hadrons to quarks in neutron stars: A review. *Rept. Prog. Phys.* **2018**, *81*, 056902. [\[CrossRef\]](#) [\[PubMed\]](#)
2. Alford, M.G.; Schmitt, A.; Rajagopal, K.; Schäfer, T. Color superconductivity in dense quark matter. *Rev. Mod. Phys.* **2008**, *80*, 1455–1515. [\[CrossRef\]](#)
3. Letessier, J.; Tounsi, A.; Heinz, U.W.; Sollfrank, J.; Rafelski, J. Strangeness conservation in hot nuclear fireballs. *Phys. Rev. D* **1995**, *51*, 3408–3435. [\[CrossRef\]](#) [\[PubMed\]](#)
4. Baldo, M.; Burgio, G.F. The nuclear symmetry energy. *Prog. Part. Nucl. Phys.* **2016**, *91*, 203–258. [\[CrossRef\]](#)
5. Chu, P.C.; Chen, L.W. Quark matter symmetry energy and quark stars. *Astrophys. J.* **2014**, *780*, 135. [\[CrossRef\]](#)
6. Chen, L.W. Symmetry Energy in Nucleon and Quark Matter. *Nucl. Phys. Rev.* **2017**, *34*, 20–28. [\[CrossRef\]](#)
7. Wu, X.; Ohnishi, A.; Shen, H. Effects of quark-matter symmetry energy on hadron-quark coexistence in neutron-star matter. *Phys. Rev. C* **2018**, *98*, 065801. [\[CrossRef\]](#)
8. Thakur, S.; Dhiman, S.K. Temperature Effects in Quark Matter Symmetry Free Energy. *DAE Symp. Nucl. Phys.* **2017**, *62*, 696–697.
9. Nakano, T.; Nishijima, K. Charge Independence for V-particles. *Prog. Theor. Phys.* **1953**, *10*, 581–582. [\[CrossRef\]](#)
10. Aryal, K.; Constantinou, C.; Farias, R.L.S.; Dexheimer, V. High-Energy Phase Diagrams with Charge and Isospin Axes under Heavy-Ion Collision and Stellar Conditions. *Phys. Rev. D* **2020**, *102*, 076016. [\[CrossRef\]](#)
11. Yao, N.; Sorensen, A.; Dexheimer, V.; Noronha-Hostler, J. Structure in the speed of sound: From neutron stars to heavy-ion collisions. *Phys. Rev. C* **2023**, *109*, 065803. [\[CrossRef\]](#)
12. Burrows, A.; Lattimer, J.M. The birth of neutron stars. *Astrophys. J.* **1986**, *307*, 178–196. [\[CrossRef\]](#)
13. Politzer, H.D. Reliable Perturbative Results for Strong Interactions? *Phys. Rev. Lett.* **1973**, *30*, 1346–1349. [\[CrossRef\]](#)
14. Freedman, B.A.; McLerran, L.D. Fermions and Gauge Vector Mesons at Finite Temperature and Density. 1. Formal Techniques. *Phys. Rev. D* **1977**, *16*, 1130. [\[CrossRef\]](#)
15. Freedman, B.A.; McLerran, L.D. Fermions and Gauge Vector Mesons at Finite Temperature and Density. 2. The Ground State Energy of a Relativistic electron Gas. *Phys. Rev. D* **1977**, *16*, 1147. [\[CrossRef\]](#)
16. Freedman, B.A.; McLerran, L.D. Fermions and Gauge Vector Mesons at Finite Temperature and Density. 3. The Ground State Energy of a Relativistic Quark Gas. *Phys. Rev. D* **1977**, *16*, 1169. [\[CrossRef\]](#)
17. Alcock, C.; Farhi, E.; Olinto, A. Strange stars. *Astrophys. J.* **1986**, *310*, 261–272. [\[CrossRef\]](#)
18. Fraga, E.S.; Pisarski, R.D.; Schaffner-Bielich, J. Small, dense quark stars from perturbative QCD. *Phys. Rev. D* **2001**, *63*, 121702. [\[CrossRef\]](#)
19. Gorda, T.; Kurkela, A.; Paatelainen, R.; Säppi, S.; Vuorinen, A. Cold quark matter at N3LO: Soft contributions. *Phys. Rev. D* **2021**, *104*, 074015. [\[CrossRef\]](#)
20. Gorda, T.; Kurkela, A.; Paatelainen, R.; Säppi, S.; Vuorinen, A. Soft Interactions in Cold Quark Matter. *Phys. Rev. Lett.* **2021**, *127*, 162003. [\[CrossRef\]](#)

21. Vuorinen, A. Particle-theory Input for Neutron-star Physics. *Acta Phys. Polon. B* **2024**, *55*, 4-A4. [[CrossRef](#)]
22. Fraga, E.S.; Romatschke, P. The Role of quark mass in cold and dense perturbative QCD. *Phys. Rev. D* **2005**, *71*, 105014. [[CrossRef](#)]
23. Kurkela, A.; Romatschke, P.; Vuorinen, A. Cold Quark Matter. *Phys. Rev. D* **2010**, *81*, 105021. [[CrossRef](#)]
24. Graf, T.; Schaffner-Bielich, J.; Fraga, E.S. The impact of quark masses on pQCD thermodynamics. *Eur. Phys. J. A* **2016**, *52*, 208. [[CrossRef](#)]
25. Gorda, T.; Säppi, S. Cool quark matter with perturbative quark masses. *Phys. Rev. D* **2022**, *105*, 114005. [[CrossRef](#)]
26. Olive, K.A. Review of Particle Physics. *Chin. Phys. C* **2014**, *38*, 090001. [[CrossRef](#)]
27. Particle Data Group; Workman, R.L.; Burkert, V.D.; Crede, V.; Klempt, E.; Thoma, U.; Tiator, L.; Agashe, K.; Aielli, G.; Allanach, B.C. Review of Particle Physics. *PTEP* **2022**, *2022*, 083C01. [[CrossRef](#)]
28. Holdom, B.; Ren, J.; Zhang, C. Quark matter may not be strange. *Phys. Rev. Lett.* **2018**, *120*, 222001. [[CrossRef](#)]
29. Jiménez, J.C. Interacting Quark Matter Effects on the Structure of Compact Stars. Ph.D. Thesis, Rio de Janeiro Federal University, Rio de Janeiro, Brazil, 2020. [[CrossRef](#)]
30. Roark, J.; Dexheimer, V. Deconfinement phase transition in proto-neutron-star matter. *Phys. Rev. C* **2018**, *98*, 055805. [[CrossRef](#)]
31. Gudmundsson, E.H.; Buchler, J.R. On the consequence of neutrino trapping in gravitational collapse. *Astrophys. J.* **1980**, *238*, 717–721. [[CrossRef](#)]
32. Kurkela, A.; Vuorinen, A. Cool quark matter. *Phys. Rev. Lett.* **2016**, *117*, 042501. [[CrossRef](#)] [[PubMed](#)]
33. Gorda, T.; Kurkela, A.; Romatschke, P.; Säppi, S.; Vuorinen, A. Next-to-Next-to-Next-to-Leading Order Pressure of Cold Quark Matter: Leading Logarithm. *Phys. Rev. Lett.* **2018**, *121*, 202701. [[CrossRef](#)] [[PubMed](#)]

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