

## AN UPPER LIMIT ON THE TAU NEUTRINO MASS

Richard D. Ehrlich  
CLEO collaboration  
Cornell University  
Ithaca NY 14853  
USA



We use the distribution of total charged energy in three-track  $\tau$  decay to place an upper limit on the mass of the  $\tau$  neutrino. Including both statistical and systematic uncertainties, the 95% confidence limit is  $m_\nu < 85 \text{ MeV}/c^2$ .

## I. Introduction

Neutrino masses have assumed increasing prominence in recent years. Indeed, several of this workshop's major topics ; e.g., dark matter, the solar neutrino deficit, and double  $\beta$ -decay are deeply connected with the possibility of non-zero mass. Because the  $\tau$  lepton at 1784 MeV/c<sup>2</sup> is the most massive charged lepton, one might expect its neutral partner,  $\nu_\tau$ , to be the most massive neutrino. In fact, in the class of theories (see-saw mechanism) where the neutrino mass scales as the *square* of the charged-lepton mass<sup>1)</sup>, an electron-neutrino mass of 20 eV would imply a  $\nu_\tau$  mass of 125 MeV/c<sup>2</sup>.

Several earlier experiments have presented limits on  $m_{\nu}^{2-5}$ ; here I report results based on a large sample of  $\tau^+\tau^-$  events collected at the Cornell Electron Storage Ring (CESR), using the CLEO detector.<sup>6)</sup> The data sample, acquired during 1982-1984, has been used in CLEO's recent determination of the Michel parameter<sup>7)</sup> in  $\tau$  decay. Both that work and an earlier one<sup>8)</sup> use the characteristic 1-vs.-3  $\tau\tau$  topology (see fig. 1) and discuss details of the event selection criteria. Because of space limitations, these will only be sketched herein.

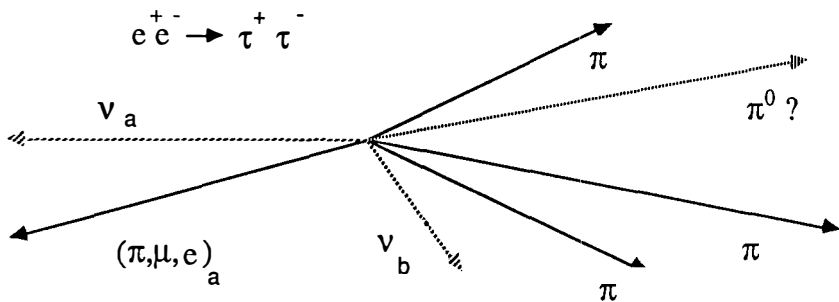


Figure 1 . 1-vs.-3 topology for  $\tau\tau$  events .

## II. Characteristics of the sample

An integrated luminosity of 128.3 pb<sup>-1</sup>, accumulated over a range of center-of-mass energy 10.34 < W < 11.18 GeV, yielded 9135  $\tau\tau$  events of the correct topology which passed the above-mentioned criteria, which are designed to reject QED and hadronic backgrounds. No attempt was made to reject events with one or more photons/ $\pi^0$ 's. This minimized the dependence upon CLEO's shower detector, which has c.a. 50% solid angle coverage and only modest resolution. Our primary experimental tools were the 1.0 Tesla solenoid and the CLEO drift chamber, which with >90% solid angle and  $\Delta P/P = 1.2\% \cdot P(\text{GeV}/c)$ , provided high efficiency tracking

and good resolution in the total energy and invariant mass of the 3-track decay. A higher-purity subsample with an identified lepton ( $e$  or  $\mu$ ) as the single track was available to allow crosschecks and internal determination of the background fraction and 3-track invariant mass distribution (See Fig.2)

### III. Method of analysis and Monte Carlo simulation

The "neutral energy"  $E_n = E_{bm} - E_{3\pi}$  is computed for each 3-track decay, where  $E_{bm}$  is the beam energy and  $E_{3\pi}$  is the sum of the 3-track energies, assuming the tracks to be pions. The  $E_n$  spectrum is nearly independent of the particular beam energy. It does depend on the masses of the missing neutrino and of the hadronic system. The low  $E_n$  endpoint (the edge of phase space) varies directly with  $m_\nu$ ; for somewhat larger  $E_n$  the spectral shape depends on a higher power of  $m_\nu$ . For very large  $E_n$  the spectrum is featureless and more affected by missing pions than by  $m_\nu$ .

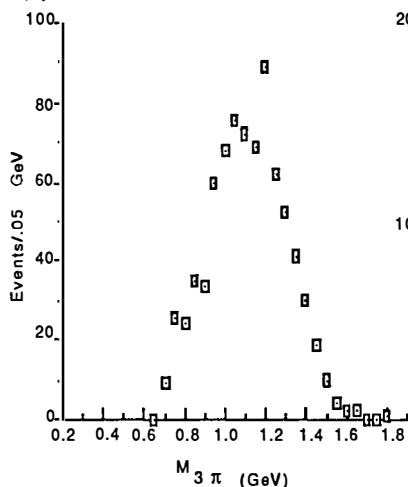


Figure 2. Invariant mass distribution for lepton-tagged events with all particles treated as pions.

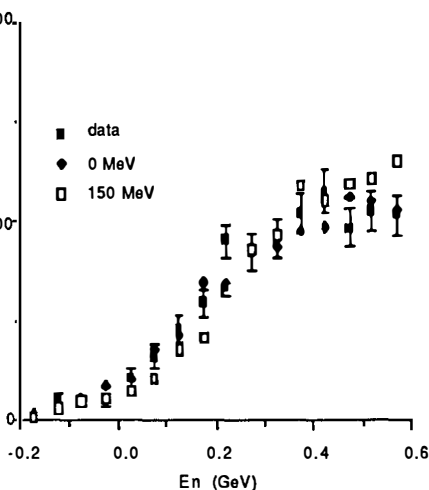


Figure 3.  $E_n$  spectra for data and for two values of  $m_\nu$ : 0 (diamonds) and 150 (closed squares).

The  $E_n$  spectrum from the data is compared to Monte Carlo simulations in which  $m_\nu$  is assigned the values 0, 25, 50, 100, and 150  $\text{MeV}/c^2$ . These simulations include realistic branching fractions into  $\rho$ 's,  $A_1$ ; etc., initial state radiation down to 5 MeV, and a 22% hadronic background, determined from the relative R values and trigger efficiencies. Various distributions (decay angle, invariant mass) are checked for conformity to the data. The events are weighted so as to yield

3-track invariant mass spectra which are constrained to agree with that derived from the lepton-tagged subsample (Fig. 2). In the comparisons, the "theoretical" distributions are normalized to the total data in the fitted region. Figure 3 shows the two cases of  $m_\nu = 0, 150 \text{ MeV}/c^2$  compared to the data, for the range  $E_n < 0.6 \text{ GeV}/c^2$ .

#### IV. Results

The data are fitted for four ranges of  $E_n$ . For each case  $\chi^2$  is plotted versus  $m_\nu$  and the minimum is found from a parabolic fit. The statistical error in  $m_\nu$  is found from the curvature of the parabola in the usual way. Averaging these results gives  $m_\nu = 31 \pm 25 \text{ MeV}/c^2$  (statistical). Dominant systematic errors are indicated as follows: 1) variation with fitted region (over and above that expected from statistics),  $\pm 15 \text{ MeV}/c^2$ , 2) absolute momentum scale,  $\pm 8 \text{ MeV}/c^2$ , 3)  $\pi^0$  fraction,  $\pm 7 \text{ MeV}/c^2$ , 4) errors in hadronic mass distribution,  $\pm 5 \text{ MeV}/c^2$ . Combining errors in quadrature yields  $m_\nu = (31 \pm 25_{\text{stat}} \pm 20_{\text{sys}}) \text{ MeV}/c^2$ . Following the Particle Data Group procedures, we extract a 95% confidence limit  $m_\nu < 85 \text{ MeV}/c^2$ . This value is of comparable sensitivity and in agreement with other recent measurements of  $m_\nu$ .

#### V. References

1. See, e.g., M. Gell-Man, P. Ramond, and R. Slansky in Supergravity, P. van Nieuwenhuizen and D. Z. Freedman eds. (North Holland Publishing Amsterdam, 1979)
2. W. Bacino et al., Phys. Rev Lett. 42, 749 (1979)
3. C. Matteuzzi et al., Phys. Rev. Lett. 52, 1869 (1984), Phys. Rev. D32, 800 (1985); P. Burchat et al., Phys. Rev. Lett. 54, 2489 (1985)
4. S. Abachi et al., Phys. Rev. Lett. 56, 1039 (1986)
5. H. Albrecht et al., Phys. Lett. 163B, 404 (1985)
6. D. Andrews et al., Nucl. Instr. and Meth. 211, 47 (1983)
7. S. Behrends et al., Phys. Rev. D32, 2468 (1985)
8. R. Giles et al., Phys. Rev. Lett. 50, 877 (1983)