# **NEXT GENERATION Nb**3**Sn CAVITIES FOR LINEAR ACCELERATORS**<sup>∗</sup>

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# *Abstract*

Niobium-3 Tin  $(Nb_3Sn)$  is a very promising alternative material for SRF accelerator cavities. The material can achieve higher quality factors, higher temperature operation and potentially higher accelerating gradients ( $E_{acc} \approx$ 96 MV/m) compared to conventional niobium. This material is formed by vaporizing Sn in a high temperature vacuum furnace and letting the Sn absorb into a Nb substrate to form a 2-3  $\mu$ m Nb<sub>3</sub>Sn layer. Current Nb<sub>3</sub>Sn cavities produced at Cornell achieve  $Q \approx 2 \cdot 10^{10}$  at 4.2 K and  $E_{acc} = 17$  MV/m. Here we present a summary of the current performance of Nb3Sn cavities at Cornell and recent progress in improving *E*acc.

# **INTRODUCTION**

 $Nb<sub>3</sub>Sn$  possesses a high critical temperature (18 K [1]) and superheating field (425 mT [2]) giving it the potential for higher quality factors than Nb, 4.2 K operation, and for accelerating gradients of  $\approx 96 \,\mathrm{MV/m}$ . Cornell has a leading program in developing  $Nb<sub>3</sub>Sn$  accelerator cavities [3–7]. Here we present the current status of the material and review recent progress. In particular we discuss the cause of the low (relative to the maximum possible for this material) quench fields and progress towards raising them.

# **Nb3Sn GROWTH PROCESS**

Nb<sub>3</sub>Sn accelerator cavities cannot be pressed out of a  $Nb<sub>3</sub>Sn$  sheet as can conventional Nb cavities due to the difficulty of making Nb<sub>3</sub>Sn and its brittle nature. Instead, current  $Nb<sub>3</sub>Sn$  cavity programs use Sn vapor deposition to grow Nb3Sn on a Nb substrate that is already in the shape of a cavity [8]. In this process an Nb cavity is placed in a high temperature vacuum furnace with a crucible full of Sn and bowl of SnCl<sub>2</sub>. After a 24 – 48 hr degas step at 180 °C, the temperature is ramped to  $530\,^{\circ}\text{C}$ , and held for 5 hrs while the nucleation agent, SnCl<sub>2</sub> vaporizes and decomposes on the Nb leaving behind Sn. The temperature is then raised to  $\approx 1150^{\circ}$ C for coating where there is an appreciable Sn vapor pressure. Sn vapor lands on the surface of the cavity where it forms  $Nb<sub>3</sub>Sn$ . To further increase the Sn vapor pressure the Sn crucible is heated to an even higher temperature. After 1.5 hrs at this temperature the crucible heating is turned off and the surface is allowed to anneal for 1 hr. This process forms a 2−3 µm Nb3Sn layer. Figure 1 shows the temperature profile for a Nb<sub>3</sub>Sn coating run.

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Coat Cavity Anneal Ramp 1200 1000  $\widehat{\omega}$ Temperature 800 600 Nucleation  $200$  $00$  $01$  $02$  $0<sub>3</sub>$  $04$ 05 06  $07$ 08 09  $10$  $11$  $12$ Time (Hours)

Figure 1: Temperature during Nb<sub>3</sub>Sn coating showing both the cavity and the tin source temperatures.

# **CURRENT PERFORMANCE**

Current performance offers transformational quality factors of  $2 \cdot 10^{10}$  at 4.2 K and 1.3 GHz (see Fig. 2). This is more than a 20 times improvement in Q over Nb at 4.2 K, allowing 4.2 K operation. In CW operation our  $Nb<sub>3</sub>Sn$  cavities currently quench around 17 MV/m, capable of meeting accelerator gradient requirement of some current CW LINACs. Pulsed operation (see Fig. 3) sees greater accelerating gradients of up to 26 MV/m [9]. These are usable accelerating gradients, but far below the theoretical maximum of 96  $MV/m$ , giving  $Nb<sub>3</sub>Sn$  a lot of room for improvement.



Figure 2: Q versus E performance of a  $Nb<sub>3</sub>Sn$  cavity versus a Nb cavity at 4.2 K and 1.3 GHz.

### **THERMOMETRY OF DEFECT**

Thermometry (temperature maps) reveal that quench are being caused by heating at some localized defect [10]. The key to raising the quench fields is removing the defect. This requires a firm understanding of what the defect is. To that end Cornell used time resolved thermometry of the quench

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Figure 3: Pulsed quench field versus temperature squared with CW quench field at 4.2 K also shown. At 4.2 K the pulsed quench field reaches 26 MV/m.

site [11]. This revealed an interesting feature of the defect (see Fig. 4): at first when raising the electromagnetic field strength in the cavity the temperature increases with the field squared, as expected from Ohmic heating, but near the quench field there appear sudden jumps in temperature. These jumps appear to be quantized (multiples of the smallest jump). When the field is lowered again there are jumps back down but they occur at a lower field.



Figure 4: A plot showing the temperature rise of the defect (on the outside of the cavity) while the RF power is cycled and the maximum field is close to, but below, the quench field. Note the sudden jumps in temperature and the hysteresis between the rise and fall of the RF power.

The quantization, hysteresis, and some additional observations discussed in [4, 11] lead us to believe this is magnetic flux entry at a defect. A potential was proposed by Ahmad Sheikhzada and Alex Gurevich [12]. In this model a bad grain boundary could act like a Josephson junction and at critical fields permit quantized flux entry. It replicates many of the observed phenomena and we will be conducting further studies into the thermal properties of the defect to determine if it matches the theory.

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### **SURFACE DEFECTS**

In addition to thermal studies we have been using microscopy to look for surface defects that degrade performance or could cause quench. Two discovered defects are Sn depleted regions [10] and thin film regions [13].

Sn depleted  $Nb<sub>3</sub>Sn$  has the correct crystal structure but with a lower atomic percentage of Sn (see Fig. 5). Sn depletion can heavily lower the critical temperature and possibly the superheating field (maximum possible quench field). These Sn depleted regions occur both on the surface and in the layer and seem to have different formation mechanisms. The surface Sn depleted regions seem to form when  $Nb<sub>3</sub>Sn$ is annealed for too long. After all the free surface Sn is gone the layer will continue growing by stealing Sn from already formed Nb<sub>3</sub>Sn, creating Sn depleted sites on the surface. So far, no surface sufficiently Sn depleted regions on the surface have been observed that could explain the current quench fields, but this may be because we can only look at a small fraction of the cavity surface and may not have seen particularly bad regions in the small sample. Sn depleted regions within the layer tend to have worse Sn depletion, but so far none have been observed near enough to the surface to impact performance. However, some may be closer to the surface and any chemical treatments to remove surface defects may expose this defect, necessitating removal/prevention. In the next section we will discuss a proposed method to prevent this defect.



Figure 5: An EDS cross-section of a  $Nb<sub>3</sub>Sn$  layer showing a tin depleted region within the  $Nb<sub>3</sub>Sn$  layer.

In addition to Sn depletion, thin regions have been observed in the  $Nb<sub>3</sub>Sn$  layer. These regions are a single crystal grain only a few hundred nm thick and large compared to the standard grain size. Figure 6 shows an EDS map of an  $Nb<sub>3</sub>Sn$  surface where the thin regions are shown in red. This may not be thick enough to fully shield the RF field from the possibly highly lossy Nb-Nb<sub>3</sub>Sn interface, decreasing Q. Great progress has been made on preventing this type of defect. It was found that growing a thick oxide on the surface (anodization) of the Nb substrate before coating could completely suppress this defect [13]. Studies into why anodization is effective are still on going.

# **LAYER GROWTH STUDIES AND GROWTH DYNAMICS**

Once a defect is identified it must be removed or we must find a way to prevent the defect from forming to be-

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Figure 6: An EDS map of a  $Nb<sub>3</sub>$ Sn surface. The regions in red are where the Nb<sub>3</sub>Sn layer is thin and unable to shield the Nb from the RF field.

gin with. To prevent defects from forming we need a good understanding of how the  $Nb<sub>3</sub>Sn$  grows. To accomplish this we have been conducting layer growth studies where the coating process is stopped at different points in the process (primarily while ramping up to the coating temperature of 1150 $\degree$ C) [13]. These samples are then analyzed with microscopy.

False colored SEM scans are shown in Fig. 7 for various stopping temperatures and for substrates that were both anodized and with only the native oxide layer. After nucleation we can already see the difference between the two preparations, with the anodized surface having large droplets of Sn (red). At 800 ◦C we can see chunks of a Sn rich Nb-Sn compound of the surface (purple) surrounded by a thin layer of Nb3Sn (green). These large Nb-Sn compounds undergo a phase transition to  $Nb<sub>3</sub>Sn$  at 875 °C where purple is thick Nb<sub>3</sub>Sn and green is thinner Nb<sub>3</sub>Sn. Finally by 1120 °C we see more uniform  $Nb<sub>3</sub>Sn$ , but the native oxide sample has formed a thin region highlighted in green.

The layer growth studies inform theoretical models based on Density Functional Theory that are being developed in collaboration with the Center for Bright Beams (CBB). Recently these models suggested a way to modify the growth process to prevent Sn depleted regions from forming in the layer [14]. In this model the Sn depleted regions form during a Sn-starved phase of the process when all the free Sn from nucleation is used before Sn the crucible gets hot enough to supply more (during the temperature ramp to the coating temperature). These sites can then get buried beneath enough  $Nb<sub>3</sub>Sn$  that they can't get filled in once more Sn is available (Sn diffusion though Nb<sub>3</sub>Sn is slow though bulk but fast through grain boundaries). A change was suggested of adding more SnCl<sub>2</sub> to prevent this from happening and is in the process of being tested.

# **CONCLUSION**

Nb3Sn is a promising material for future LINACs due to the high Q, 4.2 K operation, and high accelerating gradient limit. It currently offers high Q at 4.2 K and usable gradients. Great progress has been made in identifying and removing defects that limit  $E_{acc}$ . Many defect types have been identified such as thin film regions and Sn depleted



Figure 7: False colored SEM images of layer growth samples of Nb3Sn terminated at various temperatures in the growth process (see Fig. 1). Two different preparations are shown: one with a 70 nm oxide (anodized) and one with the native oxide layer. The images have been false colored to highlight regions with different morphology. Not false colored SEM images published in [13].

regions. Collaboration in the Center for Bright Beams with microscopists and theorists, combined with layer growth samples has led to an increased understanding of our growth process and has already led to suggestions on how to prevent Sn depleted regions. Flux entry at grain boundaries has been identified as a key quench mechanism. Further work will be conducted to understand this defect and mitigate its effects on performance.

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