



Importance of confinement in instanton induced potential for bottomonium spectroscopy

Bhoomika Pandya^{1,a}, Manan Shah^{2,b}, P. C. Vinodkumar^{1,c}

¹ Department of Physics, Sardar Patel University, Vallabh Vidyanagar, Gujarat 388120, India

² P. D. Patel Institute of Applied Sciences, CHARUSAT, Changa, India

Received: 21 October 2020 / Accepted: 22 January 2021 / Published online: 3 February 2021

© The Author(s) 2021

Abstract Mass spectra of bottomonium states are computed using the Instanton Induced potential obtained from Instanton Liquid Model for QCD vacuum and incorporating a stronger confinement term. Spin dependent interactions through confined one gluon exchange potential are incorporated to remove the mass degeneracy. The mass spectra of the $b\bar{b}$ states up to $4S$ states are found to be in good agreement with the values reported by PDG(2020). Mixing of nearby isospin states are also studied. We found the state $\Upsilon(10,860)$ as an admixture of 5^3S_1 and 6^3D_1 Upsilon states with mixing angle $\theta = 39.98^\circ$ and the mixed state di-leptonic decay width is found to be 0.25 keV as against the width of $0.31 \pm 0.07 \text{ keV}$ reported by PDG. Further the state $\Upsilon(11,020)$ is also found to be the admixture of 6^3S_1 and 5^3D_1 Upsilon states with the mixing angle $\theta = 51.69^\circ$ and the di-leptonic decay width of the mixed state is obtained as 0.14 keV which is very close to the width of $0.13 \pm 0.03 \text{ keV}$ reported by PDG. Present results indicates that addition of confinement to the instanton potential is crucial for the determination of the mass spectroscopy of heavy hadrons.

1 Introduction

Quarkonia are regarded as the simple and appropriate hadronic systems to explore the QCD aspects at the low energy regimes through its spectroscopy [1–4]. Since the discovery of $\Upsilon(1S)$, $\Upsilon(2S)$ and $\Upsilon(3S)$ at the Fermilab by E288 collaboration [5,6]; many of its orbital excited states such as $\chi_{b0,1,2}(1P)$ and $\chi_{b0,1,2}(2P)$ etc. are also discovered subsequently [7–10]. Further, continuous progress has been achieved in the bottomonium spectroscopy by the discovery of $\Upsilon(4S)$, $\Upsilon(10,860)$ and $\Upsilon(11,020)$ [11,12]. BABAR

collaboration discovered the spin singlet low lying $\eta_b(1S)$ [13] and by the efforts of CLEO, Belle and BABAR its average mass is found to be $9398.7 \pm 2.0 \text{ MeV}$ [14–16]. CLEO has given the successful observation of the $\eta_b(2S)$ with $9974.6 \pm 2.3 \pm 2.1 \text{ MeV}$ [17] but BELLE recorded the state having a mass of $9999.0 \pm 3.5 \text{ MeV}$ [18] after performing the experiment by taking much larger data sets than [17]. $\Upsilon(1^3D_2)$ also well established [19,20] and the low lying $h_b(1P)$ and $h_b(2P)$ are also reported by BABAR [21]. Recently, LHCb collaboration found the mass of $\chi_{b1}(3P)$ as $10,515^{+2.2}_{-3.9} \text{ MeV}$ [22]. Two other charged states in the bottom family, $Z_b(10,610)$ and $Z_b(10,650)$ are also being reported recently [23]. Their electric charge suggest that they do not belong to conventional bottomonium [24,25]. So we do not include them in the present work.

Looking to the advances in the experimental side, it is necessary to review earlier theoretical attempts to understand the quarkonia systems. Although there are many different approaches like Lattice QCD methods [26–29], NRQCD methods [30–32], Light front quark model [33–36], various coupled channel quark models [37–39], Effective Lagrangian approach [40] etc., employed to study them, still there exist no consensus and discrepancy persist in the predictions. For instance, authors [41] have used nonrelativistic constituent quark model wherein the lowest lying bottomonium states $\Upsilon(1S)$ and $\eta_b(1S)$ are about 50 MeV higher than the PDG [43] data but the same model give fine agreement for the higher excited state $\Upsilon(6S)$ while the relativized quark model [42] provides very good description of the low-lying states but higher excited $\Upsilon(6S)$ is overestimated by 100 MeV .

It is believed that most reliable description of the bound system can be obtained from Bethe–Salpeter formalism. It is a relativistic quantum field theory but ambiguity is that the interaction kernel is not solvable from QCD and it becomes very tough to extract hadronic information [44]. In EFTs, threshold region is still disappointing [46]. Lattice regularized QCD beyond threshold and full understanding of bot-

^a e-mail: bhumispandya@gmail.com (corresponding author)

^b e-mail: mshah09@gmail.com

^c e-mail: p.c.vinodkumar@gmail.com

tomonium seems difficult [47–54] even though very recently they have predicted few radial excitations.

Currently, computational complexity of other theoretical approaches makes the potential model approach as the most reliable which is able to meet the expectations. Potential models provide quantitative and qualitative analysis of the quarkonia states. Potential models account for both the shorter and larger distance behaviour of the $q\bar{q}$ interaction. The most promising non-perturbative contribution is the “Quark Confinement” which can be incorporated by the “Wilson loop”. Such non-perturbative effects will increase at the larger distances [2,55]. And such attempts are key to the lattice QCD calculations. Apart from the Wilson loop, the contribution from Instantons [56] is also considered for the non-perturbative effects at the larger distance. Instantons, which are the large fluctuations of the gluon field and corresponds to the tunneling from one minimum of the energy to the neighbouring one. Similar fluctuations but having tunneling in the opposite side is called “anti-instantons”.

In the present work we aim to find the effect of the instantons on the heaviest quarkonium system. As the form of the instanton potential contains the term which gives the non-perturbative effect at the larger distance as well as coulombian type behaviour of the shorter distance together make it suitable for the study of the spectroscopy of bottomonium. Additionally, we have also incorporated the spin dependent part of one gluon exchange potential. The detailed description of the form of instanton potential for heavy quark can be found in [57] where authors have added the spin dependent attributes into it. This potential has a long history [59–61] and its central part was derived long ago [62] based on the instanton liquid model [63,66,67] for the QCD vacuum. With the use of instanton potential acting between quark and antiquark pair, we solved the non-relativistic Schrödinger equation by variational method. To test the predictivity of the potential we have further calculated various decay properties like vector and pseudoscalar decay constants, di-leptonic, di-gamma, di-gluon as well as tri-gluon decays together with radiative transition decay widths. The computed results are tabulated and compared with other available theoretical and experimental data and finally we draw important conclusions.

2 Theoretical framework to compute mass spectra

Evaluation of the bound state mass spectra demands the solution of the Schrödinger equation with the potential from the instanton vacuum. Unfortunately, the Schrödinger equation is not exactly solvable for most of the systems. So, generally one adopts approximate methods like perturbation, variational or WKB methods. In the present work, the variational method is employed to calculate the ground and excited state mass spectra. For the treatment of bottomonium on the non-

relativistic footing the Hamiltonian can be written as

$$H = \frac{P^2}{2M} + V(r) \quad (1)$$

Here, P is the relative momenta, M represents reduced mass, $V(r)$ is the potential acting between quark and antiquark. The QCD instanton effects are important to apprehend the phenomena in the non-perturbative region in QCD. It is believed that instantons are present in the QCD vacuum and they are working effectively in reproducing many remarkable features of the strong interactions [64]. The most significant success of instantons is their capacity to provide a microscopic mechanism of the spontaneous chiral symmetry breaking [64,65]. But there exist lack of successful numerical predictions of the mass spectroscopy and decay properties based on the instanton inspired models. A recent attempt by Yakhshiev et al. [57] have obtained the mass spectra of low lying quarkonium states. For the case of ground state of bottomonium they found differences of 1006 MeV and 945 MeV for $\Upsilon(1S)$ and $\eta_b(1S)$ respectively as compared to PDG. For the $(1P)$ state difference was more than 1200 MeV. Also, the hyperfine contribution was very much less in their study as compared to the experimental mass splittings between spin triplet and spin singlet states. According to Schäfer and Shuryak in a review paper on “Instantons in QCD”, it is clearly stated that for heavy flavour hadrons and for higher states, the confinement effect play an important role [58] following this suggestions from [58] and looking at the results of the attempt made by [57] we found that stronger confinement effects are needed to improve the results. Thus, we have chosen instanton potential incorporating additional confinement part in the potential to predict the mass spectra and other decay properties in the bottom sector. So, for present study the form of the potential we have used is instanton induced potential according to the Instanton Liquid Model (ILM) [57]. The ILM has two important parameters portraying the diluteness of the instanton liquid [63,68]: one is the average size of instanton ($\bar{\rho}$) and other is the average distance between instantons (\bar{R}). Numerical values of these parameters are different for different approaches. For example, Shuryak [68] proposed the values of these parameters as $\bar{\rho} \simeq 0.33$ fm and $\bar{R} \simeq 1$ fm. In ref [69–71], $\bar{\rho} \simeq 0.35$ fm and $\bar{R} \simeq 0.856$ fm were used in their $1/N_c$ meson loop contribution for the light quark sector. And in this framework the range of the potential is identified only by the size of the instanton ($\bar{\rho}$).

For $r \ll \bar{\rho}$ i.e., when the distance between quark-antiquark is smaller than the average size of instanton, the central potential is given by [57]

$$V(r) \simeq \frac{4\pi\bar{\rho}^3}{R^4 N_c} \left(1.345 \frac{r^2}{\bar{\rho}^2} - 0.501 \frac{r^4}{\bar{\rho}^4} \right) \quad (2)$$

here, $N_c = 3$ represents the colour degrees of freedom.

For $r \gg \bar{\rho}$ i.e., the distance between quark-antiquark is higher than the size of instanton, the central part of the potential is given by [57]

$$V(r) \simeq 2\Delta M_Q - \frac{g_{np}}{r} \quad (3)$$

here, ΔM_Q is the correction to heavy-quark mass from the instanton vacuum and is given by [57].

$$\Delta M_Q = -\frac{4\pi^4 \bar{\rho}^3}{3\bar{R}^4 N_c} \left(J_0(\pi) + \frac{1}{\pi} J_1(\pi) \right) \quad (4)$$

where, J_0 and J_1 are the Bessel functions. The coupling constant in Eq. (3) is defined as $g_{np} = 2\pi^3 \bar{\rho}^4 / N_c \bar{R}^4$ and can be considered as nonperturbative correction to the strong coupling constant $\alpha_s(r)$. When r reaches to infinity, the potential is saturated at $2\Delta M_Q$ and it signifies that instanton vacuum cannot explain quark confinement [62]. Thus for the present study, we have added a state dependent confinement potential V_0 which has the form

$$\begin{aligned} V_0(z) &= b \ln(z) + a \\ &= b \ln(2N + l - 1) + a \end{aligned} \quad (5)$$

Where, $z = (2N + 3) - (4 - l)$ and $N = 2n + l$. Here, n is the radial quantum number and l is the orbital quantum number. z represents the eigenvalues corresponds to modified harmonic oscillator like potential. The modification can also be attributed to the centre of mass corrections which is state dependent. V_0 is then a logarithmically dependent eigenvalue corresponds to a modified three dimensional harmonic oscillator potential. The parameters a and b are our fitted model parameters which are deduced by fitting the masses of low lying bottomonium states based on instanton potential framework. The numerical values of the constants are $a = 211.1$ MeV and $b = 290.48$ MeV. For instanton liquid model framework, potential is defined only for $r \ll \bar{\rho}$ and $r \gg \bar{\rho}$ and to remove the discontinuity at $r = \bar{\rho}$ we have further added $V_1 = 61.81$ MeV to the long range part of the interaction represented by Eq. (3). It is justifiable from the QCD point of view that hadronic bound states has contributions from the long range part of the interactions. This inclusion changes the potential range from $r \ll \bar{\rho}$ to $r \leq \bar{\rho}$ and $r \gg \bar{\rho}$ to $r \geq \bar{\rho}$. The instanton potential parameters as $\bar{\rho} \simeq 0.36$ fm and $\bar{R} \simeq 0.89$ fm from the lattice simulation study of Instanton vacuum [72, 73] are being used.

Figure 1 shows the potential as a function of radial distance for few low lying states incorporating the addition of 61.81 MeV to Eq. (3). Exclusion of this addition of 61.81 MeV to Eq. (3) leads to kink at $r = 0.36$ fm in the behaviour of the potential. The potential corresponds to the instanton liquid model as a continuous function of the radial separation with and without the state dependent confinement part are plotted in Fig. 1. It is to be noticed that the addition of

confinement provides a shift to the potential while retaining the shape.

The harmonic oscillator trial wave-function used here is

$$R_{nl}(r) = N_{nl}(\mu r)^l \exp\left(-\mu^2 r^2/2\right) L_{n-1}^{l+1/2}(\mu^2 r^2) \quad (6)$$

with N_{nl} being the normalization constant expressed as

$$N_{nl} = \sqrt{\frac{2\mu^3 (n-1)!}{\Gamma(n+l+1/2)}} \quad (7)$$

Also, μ is the variational parameter and $L_{n-1}^{l+1/2}(\mu^2 r^2)$ is the associated Laguerre polynomial.

We employ the Ritz variational method by minimizing the expectation value of the Hamiltonian to obtain the variational parameter μ corresponds to the ground state. In the case of excited states, we have employed the Virial theorem ($\langle T \rangle = \frac{1}{2} \langle r \frac{\partial V(r)}{\partial r} \rangle$) to estimate the variational parameter. The deduced variational parameter correspond to few low-lying states are given in Table 1. Based on the above theoretical formalism, we have computed the spectroscopic masses of the S , P and D wave bottomonium. The spin averaged masses of the radial (nS) and orbital excited states (nP , nD) are computed as

$$M_{SA} = m_b + m_{\bar{b}} + \langle H \rangle \quad (8)$$

Further spin dependent confined one gluon exchange interactions are used to remove the mass degeneracy of the bottomonium states. The S wave degeneracy is removed by introducing the hyperfine interaction given by

$$V_{Q\bar{Q}}^{SS}(r) = \frac{8}{9} \frac{\alpha_s}{m_Q m_{\bar{Q}}} \mathbf{S}_Q \cdot \mathbf{S}_{\bar{Q}} 4\pi \delta^3(r) \quad (9)$$

For the masses of P and D waves, we have incorporated the contribution of the spin-orbit and tensor part of the confined one gluon exchange potential (COGEP) of the form given by [77, 78, 93, 110–112]

$$\begin{aligned} V_{Q\bar{Q}}^{LS}(r) &= \frac{\alpha_s}{4} \frac{N_Q^2 N_{\bar{Q}}^2}{m_Q m_{\bar{Q}}} \frac{\lambda_Q \cdot \lambda_{\bar{Q}}}{2r} \\ &\otimes \left[\left[\mathbf{r} \times (\hat{p}_Q - \hat{p}_{\bar{Q}}) \cdot (\sigma_Q + \sigma_{\bar{Q}}) \right] (D'_0(r) + 2D'_1(r)) \right. \\ &\quad \left. + \left[\mathbf{r} \times (\hat{p}_Q + \hat{p}_{\bar{Q}}) \cdot (\sigma_i - \sigma_j) \right] (D'_0(r) - D'_1(r)) \right] \end{aligned} \quad (10)$$

and

$$\begin{aligned} V_{Q\bar{Q}}^T(r) &= -\frac{\alpha_s}{4} \frac{N_Q^2 N_{\bar{Q}}^2}{m_Q m_{\bar{Q}}} \\ &\otimes \lambda_Q \cdot \lambda_{\bar{Q}} \left(\left(\frac{D'_1(r)}{3} - \frac{D'_1(r)}{3r} \right) S_{Q\bar{Q}} \right) \end{aligned} \quad (11)$$

where $S_{Q\bar{Q}} = [3(\sigma_Q \cdot \hat{r})(\sigma_{\bar{Q}} \cdot \hat{r}) - \sigma_Q \cdot \sigma_{\bar{Q}}]$ with $\hat{r} = \hat{r}_Q - \hat{r}_{\bar{Q}}$ as the unit vector in the direction of \mathbf{r} . N_Q and $N_{\bar{Q}}$ are the normalization constants of the wave functions for quark and

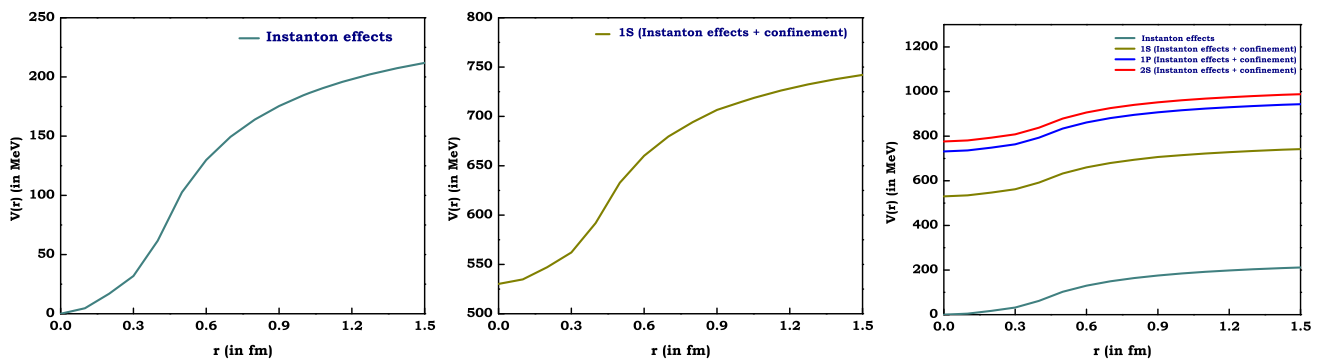


Fig. 1 The heavy quark potential from instanton vacuum for the $r \leq \bar{\rho}$ and $r \geq \bar{\rho}$ without confinement (matte green) and with confinement for 1S (dark yellow), 2S (red) and 1P (blue)

Table 1 The variational parameter (μ), spin-average mass and expectation values of spin-dependent contribution for few low-lying bottomonium bound states employing only the instanton effects without addition of confinement (a) and instanton effects with confinement (b)

State	Variational parameter μ (in MeV)	M_{SA} (in MeV)	$\langle V_{SS} \rangle$ (in MeV)	$\langle V_S \rangle$ (in MeV)	$\langle V_T \rangle$ (in MeV)
(a) Instanton potential without confinement term (IMWOC)					
1S	479.9	8526.19	2.36
2S	410.6	8641.34	2.24
1P	460.13	8613.87	...	0.85	0.24
2P	384.78	8681.89	...	0.60	0.17
1D	432.52	8675.38	...	0.30	0.08
2D	368.42	8719.95	...	0.21	0.05
(b) Instanton potential with confinement term (IMWC)					
1S	1233	9448.61	48.53
2S	995.2	10018.54	30.74
1P	1123	9874.56	...	7.78	2.35
2P	941.9	10,270.00	...	5.43	1.65
1D	1035	10,153.80	...	2.46	0.72
2D	892.6	10,456.6	...	1.77	0.50

antiquark. $\lambda_Q \lambda_{\bar{Q}}$ is the color factor and can be calculated separately as it is independent of the flavour contents of the quarks. For mesons $\langle \lambda_i \lambda_j \rangle_{q\bar{q}} = -\frac{4}{3}$ [115].

Expression for the confined gluon propagators are taken from [77,93] as,

$$D_0(r) = \left(\frac{\alpha_1}{r} + \alpha_2 \right) \exp(-r^2 c_0^2/2) \quad (12)$$

$$D_1(r) = \frac{\gamma}{r} \exp(-r^2 c_1^2/2) \quad (13)$$

where, $\alpha_1 = 1.035$, $\alpha_2 = 0.3977$ GeV, $c_0 = 0.3418$ GeV, $\gamma = 0.8639$ and $c_2 = 0.4123$ GeV are used here as the same as used in previous study of quarkonium states [93].

The strong running coupling constant α_s is calculated using the form,

$$\alpha_s(M^2) = \frac{4\pi}{\left(11 - \frac{2}{3}n_f\right) \left(\ln \frac{M^2}{\Lambda^2}\right)} \quad (14)$$

where, Λ_{QCD} is taken as 0.156 GeV which will give the PDG [43] listed value of α_s at the Z meson mass (91 GeV) range of 0.0118. Employing $\Lambda_{QCD} = 0.156$ GeV and putting spin average mass (M_{SA}) of respective state in Eq. (14) gives α_s for corresponding state. Here, n_f represents the number of flavours and for bottom sector it is 4.

To assess the importance of confinement in the instanton models, we have computed the spin average masses of the few low-lying states of bottomonium without considering the confinement part (Eqs. (2) and (3)) and with incorporating confinement part (Eqs. (2), (3) and (5)). The results are presented here in Table 1. It is seen that the predicted low-lying spin-average masses without confinement are off by about 900 MeV or more as found in the study reported by Ref. [57]. While the Instanton potential with confinement (IMWC) generates the spin average mass very close to the experimentally deduced values. It is observed that the average hyperfine and fine structure contribution in the case of instan-

Table 2 S-wave mass spectrum of bottomonium bound states (in MeV). Results are compared with Experimental data [43] Relativistic Dirac model [93] Relativistic functional approach [107] Constituent quark model [41] Relativistic quark model [42] Non- relativistic potential model [101] Relativistic quark model [108] Relativistic potential model [100]

nS	M_{SA}	State	J^{PC}	Spin-Spin contribution	Present	PDG [43]	[93]	[107]	[41]	[42]	[101]	[108]	[100]
1S	9448.61	1^3S_1	1^{--}	12.13	9460.75	9460.30 ± 0.26	9460.99	9490	9502	9465	9460	9460	9460
		1^1S_0	0^{-+}	- 36.39	9412.22	9398.7 ± 2.0	9390.7	9414	9455	9402	9392	9398	9393
2S	10,018.54	2^3S_1	1^{--}	7.68	10,026.22	$10,023.26 \pm 0.31$	10,024.1	10,089	10,015	10,003	10,024	10,023	10,023
		2^1S_0	0^{-+}	- 23.05	9995.48	...	9999.3	9987	9990	9976	9991	9990	9987
3S	10,358.24	3^3S_1	1^{--}	6.41	10,364.65	$10,355.2 \pm 0.5$	10,356.2	10,327	10,349	10,354	10,346	10,355	10,364
		3^1S_0	0^{-+}	- 19.24	10,339.00	...	10,325.3	...	10,330	10,336	10,323	10,329	10,345
4S	10,588.98	4^3S_1	1^{--}	5.49	10,594.47	$10,579.4 \pm 1.2$	10,576.2	...	10,607	10,635	10,575	10,586	10,643
		4^1S_0	0^{-+}	- 16.48	10,572.49	...	10,554.4	10,523	10,558	10,573	10,364
5S	10,761.30	5^3S_1	1^{--}	4.84	10,766.14	$10,885.2^{+2.6}_{-1.6}$	10,758.5	...	10,818	10,878	10,755	10,869	...
		5^1S_0	0^{-+}	- 14.53	10,746.76	...	10,738.4	10,869	10,741	10,851	...
6S	11,077.39	6^3S_1	1^{--}	4.31	11,081.70	$11,000 \pm 4$	10,995	11,243	10,904	11,088	...
		6^1S_0	0^{-+}	- 12.92	11,064.47	11,226	10,892	11,061	...

ton model without confinement (IMWOC) is very small compared to the case with confinement. For example, the ground state (1^3S_1) mass of the bottomonium system in the case of IMWOC is obtained as 8326 MeV as against the experimental mass of 9460 MeV. Further the energy difference between $\Upsilon(1S)$ and $\Upsilon(2S)$ in the case of IMWOC is just 115 MeV as against the experimental mass difference of 565 MeV. In is very clear then that, the instanton model without confinement term does not provide the masses in accordance with the experimental values. Hence, in the present study of bottomonium states we incorporate confinement along with the Instanton potential.

Results obtained here are tabulated in the Table 2 for S wave masses and in Table 3 for P and D wave masses. We have compared our data set with the theoretical model predictions such as relativistic Dirac Model [93], QCD Relativistic functional approach [107] where authors have used two different methods rainbow-ladder truncation of Dyson–Schwinger and Bethe–Salpeter in search of the effects of the varying shapes of the effective running coupling on ground as well as excited states in the channels having quantum numbers J less than or equal to 3, Constituent quark model [41], Relativistic quark model [42, 108], Relativistic potential model [100] and with the recent experimental data listed by PDG [43]. Also, Fig. 2 shows the energy level diagram of the bottomonium states comparing results with the PDG [43] reported values.

3 Decay properties of heavy quarkonia $Q\bar{Q}$

For any potential model, apart from the mass spectra other observables like radiative decays and annihilation decays are

important testing ground to know the inter-quark interactions. Keeping this view, we have computed the decay properties of $b\bar{b}$ states with no additional parameters. Further, We have incorporated these decay properties with and without QCD correction.

3.1 Pseudoscalar and vector decay constants

Estimation of the decay constants of mesons constituting heavy quarks is essential part of the study as it offers the information of the CKM (Cabibbo–Kobayashi–Maskawa) matrix elements. The conventional formula within the non-relativistic limit for the pseudoscalar and vector states of the S wave heavy quarkonia is the Van Royen–Weiskopf formula [74] given by

$$f_{P/V}^2(nS) = \frac{3|R_{nS}(0)|^2}{\pi M_{nS}} \bar{c}^2(\alpha_s) \quad (15)$$

Where, the first order QCD correction factor $\bar{c}^2(\alpha_s)$ is expressed as [75, 76]

$$\bar{c}(\alpha_s) = 1 - \frac{\alpha_s}{\pi} \left(\delta^{P,V} - \frac{m_1 - m_2}{m_1 + m_2} \ln \frac{m_1}{m_2} \right) \quad (16)$$

In the case of quarkonia, the second term inside the bracket in Eq. (16) vanishes. The term $\delta^V=8/3$ in the case of vector state and $\delta^P=2$ for the pseudoscalar state. The calculated vector and pseudoscalar decay constants with and without the first order QCD corrections are presented in Tables 4 and 5 and are compared with other available model predictions.

Table 3 P-wave and D-wave mass spectra of bottomonium states (in MeV). The Experimental data are from [43]

nL	M_{SA}	State	J^{PC}	Spin-Orbit contribution	Tensor contribution	Present	PDG [43]	[93]	[107]	[41]	[42]	[101]	[108]	[100]	[102]
1P	9874.56	1^3P_2	2^{++}	7.78	-0.94	9881.40	9912.21 ± 0.26	9912.3	9906	9886	9897	9908	9912	9912	9918
		1^3P_1	1^{++}	-7.78	4.69	9871.47	9892.78 ± 0.26	9901.8	9842	9874	9876	9888	9892	9891	9897
		1^3P_0	0^{++}	-15.56	-9.38	9849.61	9859.44 ± 0.42	9889.2	9815	9855	9847	9862	9859	9861	9865
		1^1P_1	1^{+-}	0	0	9874.56	9899.3 ± 0.8	9854.1	9806	9879	9882	9896	9900	9900	9903
2P	10,270	2^3P_2	2^{++}	5.43	-0.66	10,274.77	$10,268.65 \pm 0.22$	10,265.9	...	10,246	10,261	10,268	10,268	10,271	10,269
		2^3P_1	1^{++}	-5.43	3.29	10,267.86	$10,255.46 \pm 0.22$	10,258.9	10,120	10,236	10,246	10,256	10,255	10,255	10,251
		2^3P_0	0^{++}	-10.86	-6.59	10,252.54	$10,232.50 \pm 0.40$	10,234.7	10,254	10,221	10,226	10,241	10,233	10,230	10,226
		2^1P_1	1^{+-}	0	0	10,270.00	...	10,264.9	10,154	10,240	10,250	10,261	10,260	10,262	10,256
3P	10,526.5	3^3P_2	2^{++}	4.23	-0.51	10,530.21	$10,524.0 \pm 3.8$	10,516.9	...	10,521	10,550	10,516	10,550	...	10,540
		3^3P_1	1^{++}	-4.23	2.57	10,524.84	$10,513.4 \pm 0.7$	10,508.8	10,303	10,513	10,538	10,507	10,541	...	10,529
		3^3P_0	0^{++}	-8.46	-5.15	10,512.88	...	10,497.6	...	10,500	10,552	10,511	10,521	...	10,502
		3^1P_1	1^{+-}	0	0	10,526.50	...	10,540.2	...	10,516	10,541	10,497	10,544	...	10,529
4P	10,714.8	4^3P_2	2^{++}	3.49	-0.43	10,717.86	...	10,707.0	...	10,744	10,798	...	10,812	...	10,767
		4^3P_1	1^{++}	-3.49	2.13	10,713.44	...	10,706.5	10,788	...	10,802	...	10,753
		4^3P_0	0^{++}	-6.98	-4.28	10,703.56	...	10,703.8	10,781	...	10,781	...	10,732
		4^1P_1	1^{+-}	0	0	10,714.80	...	10,704.6	10,790	...	10,804	...	10,757
5P	10,863	5^3P_2	2^{++}	2.98	-0.36	10,865.62	11,022	10,965
		5^3P_1	1^{++}	-2.98	1.82	10,861.83	11,014	10,951
		5^3P_0	0^{++}	-5.97	-3.64	10,853.38	11,004	10,933
		5^1P_1	1^{+-}	0	0	10,863	11,016	10,955
1D	10,153.8	1^3D_3	3^{--}	4.92	-0.41	10,158.31	...	10,140.4	10,232	10,127	10,115	10,177	10,166	10,163	10,156
		1^3D_2	2^{--}	-2.46	1.43	10,152.77	$10,163.7 \pm 1.4$	10,138.7	10,145	10,122	10,147	10,162	10,161	10,157	10,151
		1^3D_1	1^{--}	-7.38	-1.43	10,144.99	...	10,136.0	...	10,117	10,138	10,147	10,154	10,149	10,145
		1^1D_2	2^{-+}	0	0	10,153.80	...	10,068.2	10,194	10,123	10,148	10,166	10,163	10,158	10,152

Table 3 continued

nL	M_{SA}	State	J^{PC}	Spin-Orbit contribution	Tensor contribution	Present	PDG [43]	[93]	[107]	[41]	[42]	[101]	[108]	[100]	[102]
2D	10,456.6	2^3D_3	3^{--}	3.55	-0.29	10,459.85	...	10,398.7	...	10,422	10,455	10,447	10,449	10,456	10,442
		2^3D_2	2^{--}	-1.77	1.04	10,455.86	...	10,397.1	...	10,418	10,449	10,437	10,443	10,450	10,438
		2^3D_1	1^{--}	-5.32	-1.04	10,450.23	...	10,395.7	...	10,414	10,441	10,428	10,435	10,443	10,432
		2^1D_2	2^{-+}	0	0	10,456.60	...	10,336.0	...	10,419	10,450	10,440	10,445	10,452	10,439
3D	10,664.7	3^3D_3	3^{--}	2.79	-0.23	10,667.25	...	10,620.9	10,711	10,652	10,717	...	10,680
		3^3D_2	2^{--}	-1.39	0.83	10,664.12	...	10,619.3	10,705	10,645	10,711	...	10,676
		3^3D_1	1^{--}	-4.19	-0.82	10,659.68	...	10,616.8	10,698	10,637	10,704	...	10,670
		3^1D_2	2^{-+}	0	0	10,664.70	...	10,564.3	10,706	10,646	10,713	...	10,677
4D	10,823	4^3D_3	3^{--}	2.31	-0.19	10,825.12	...	10,820.9	10,939	10,817	10,963	...	10,886
		4^3D_2	2^{--}	-1.15	0.69	10,822.52	...	10,819.3	10,934	10,811	10,957	...	10,882
		4^3D_1	1^{--}	-3.48	-0.69	10,818.83	...	10,816.9	10,928	10,805	10,949	...	10,877
		4^1D_2	2^{-+}	0	0	10,823.00	...	10,768.8	10,935	10,813	10,959	...	10,883
5D	10,952.6	5^3D_3	3^{--}	1.99	-0.17	10,954.42	...	11,005.2	10,955	11,069
		5^3D_2	2^{--}	-0.99	0.59	10,951.59	...	11,003.7	10,950	11,065
		5^3D_1	1^{--}	-2.99	-0.59	10,949.01	...	11,001.4	10,945	11,060
		5^1D_2	2^{-+}	0	0	10,952.60	...	10,956.7	10,952	11,066
6D	11,057.6	6^3D_3	3^{--}	1.75	-0.15	11,059.20	11,075
		6^3D_2	2^{--}	-0.88	0.52	11,057.24	11,071
		6^3D_1	1^{--}	-2.63	-0.52	11,054.45	11,066
		6^1D_2	2^{-+}	0	0	11,057.60	11,072

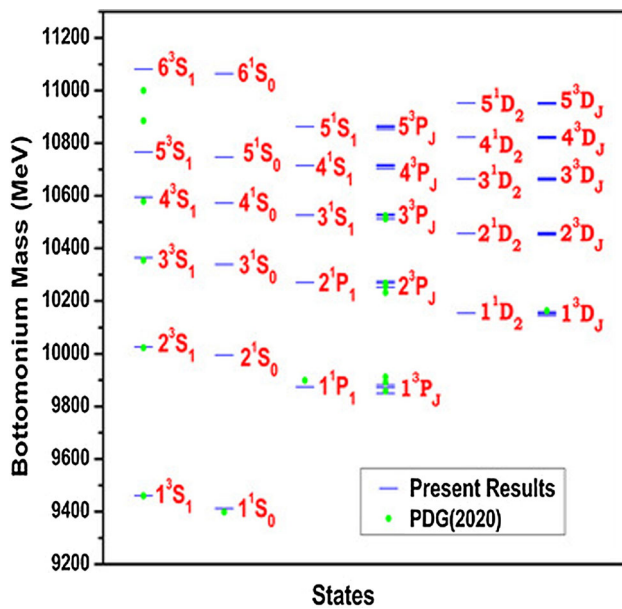


Fig. 2 Bottomonium ($b\bar{b}$) mass spectrum. Comparison of present results (blue horizontal lines) with experimental results by Particle Data Group [43] (green full circles)

3.2 The leptonic decays of bottomonium states

Descriptive study of leptonic decays of mesons not only provide underlying dynamics of quark anti-quark annihilation [80] but also used as a probe to study compactness of the pair of quark-antiquark bound states [79]. In addition to that

leptonic decay rates can also help us to distinguish conventional mesons and exotic states [81].

The vector mesons (1^{--}) annihilates leptonically via single virtual photon and leptonic partial decay width computed using Van Royen- Weisskopf formula [74] along with the QCD correction of one loop level given as [82,83]

$$\Gamma(n^3S_1 \rightarrow l^+l^-) = \frac{4\alpha_e^2 e_Q^2}{M^2(n^3S_1)} |R_{nl}(0)|^2 \left[1 - \frac{16\alpha_s}{3\pi} \right] \quad (17)$$

Here, $|R_{nl}(0)|^2$ is the square of the radial wave function at the origin. $\alpha_e = 1/137$ and α_s are the electromagnetic coupling constant and strong running coupling constant respectively. e_Q is the charge of the heavy quark in units of electron charge. For the D wave states n^3D_1 leptonic decay width can be obtained as [84],

$$\Gamma(n^3D_1 \rightarrow l^+l^-) = \frac{25\alpha_e^2 e_Q^2}{2m_b^4 M^2(n^3D_1)} |R''_{nl}(0)|^2 \left[1 - \frac{16\alpha_s}{3\pi} \right] \quad (18)$$

For the n^3D_1 bound states annihilation into l^+l^- have same first order QCD correction factor as of n^3S_1 bound states [114]. The leptonic decay widths are tabulated in Table 6 with and without QCD corrections. We have compared our results with other available experimental as well as theoretical data.

Table 4 The vector decay constant (in MeV) of the S wave bottomonium states. f_V and $f_V(C)$ are present results without and with QCD correction respectively. The Experimental data is taken from [43]

State	J^{PC}	f_V	$f_V(C)$	PDG [43]	[90]	[93]	[92]	[91]	[94]
1^3S_1	1^{--}	653.44	551.53	715 ± 5	649(31)	705.4	706	$498 \pm (20)$	831
2^3S_1	1^{--}	563.73	477.05	498 ± 8	481(39)	554.9	547	$366 \pm (27)$	566
3^3S_1	1^{--}	507.90	430.42	430 ± 4	...	436.8	484	$304 \pm (27)$	507
4^3S_1	1^{--}	466.60	395.80	336 ± 18	...	332.4	446	$259 \pm (22)$	481
5^3S_1	1^{--}	434.61	368.91	286.5	419	$228 \pm (16)$	458
6^3S_1	1^{--}	406.39	345.40

Table 5 The pseudoscalar decay constant (in MeV) of the S wave bottomonium states. f_P and $f_P(C)$ are present results without and with QCD correction respectively. There is no Experimental data available for comparison. Predictions from Lattice QCD and QCD sum rules are compared

State	J^{PC}	f_P	$f_P(C)$	[95]	[96]	Lattice [98]	[94]	QCD sum rules [99]	[97]
1^1S_0	0^{-+}	654.81	578.21	646.025	756	667	834	251 ± 0.072	1016.8
2^1S_0	0^{-+}	564.59	499.48	518.803	285	...	567	...	806.6
3^1S_0	0^{-+}	508.53	450.35	474.954	333	...	508	...	713.4
4^1S_0	0^{-+}	467.09	413.93	449.654	40
5^1S_0	0^{-+}	435.00	385.68	432.072
6^1S_0	0^{-+}	406.71	360.93

3.3 The Di-gamma decay of bottomonium states

Adopting the case of the para positronium (1S_0) state decays to two photons to that of the 1S_0 state of quarkonium with the inclusion of color factor, the expression to compute di-gamma decay width of pseudoscalar mesons can be written as [74]

$$\Gamma(n^1S_0 \rightarrow \gamma\gamma) = \frac{12\alpha_e^2 e_Q^4}{M^2(n^1S_0)} |R_{nl}(0)|^2 \times \left[1 - \frac{\alpha_s}{\pi} \left(\frac{20 - \pi^2}{3} \right) \right] \quad (19)$$

where bracketed quantity is the lowest order QCD correction [82]. In the case of P wave states, decay width depends on the first derivative of the radial wave function at the origin and is given by [82]

$$\Gamma(n^3P_0 \rightarrow \gamma\gamma) = \frac{27\alpha_e^2 2^4}{M(n^3P_0)^4} e_Q^4 |R'_{nl}(0)|^2 \times \left[1 + \frac{\alpha_s}{\pi} \left(\frac{\pi^3}{3} - \frac{28}{9} \right) \right] \quad (20)$$

$$\Gamma(n^3P_2 \rightarrow \gamma\gamma) = \frac{36\alpha_e^2 2^4}{5M(n^3P_2)^4} e_Q^4 |R'_{nl}(0)|^2 \times \left[1 - \left(\frac{16\alpha_s}{3\pi} \right) \right] \quad (21)$$

Table 7 summarises our computed results for the di-gamma decay width of n^1S_0 and $n^3P_{0,2}$ states and are compared with other available data.

3.4 The Di-gluon decay of bottomonium states

Similar to the di-photon decay, states holding even charge conjugation can decay into di-gluons. Expression used to compute S wave spin-singlet state decaying into two gluons together with the QCD correction is given by [82]

$$\Gamma(n^1S_0 \rightarrow gg) = \frac{2\alpha_s^2}{3m_Q^2} |R_{nl}(0)|^2 \left[1 + \frac{4.4\alpha_s}{\pi} \right] \quad (22)$$

Conventionally, theoretical expression for di-gluon decay width of the χ_{b0} and χ_{b2} states depend on the first derivative of the radial wave function at origin and the formula adopted here for the computation with QCD correction factor can be written as [82, 85, 86]

$$\Gamma(n^3P_0 \rightarrow gg) = \frac{6\alpha_s^2 2^4}{M(n^3P_0)^4} |R'_{nl}(0)|^2 \left[1 + \frac{10.0\alpha_s}{\pi} \right] \quad (23)$$

$$\Gamma(n^3P_2 \rightarrow gg) = \frac{8\alpha_s^2 2^4}{5M(n^3P_2)^4} |R'_{nl}(0)|^2 \left[1 - \frac{0.1\alpha_s}{\pi} \right] \quad (24)$$

The calculated di-gluon decay width with other available theoretical data are listed in Table 8.

3.5 The Tri-gluon decay of bottomonium states

The decay width of S wave vector state annihilates into three gluons along with QCD radiative correction is computed using the relation given by [82, 113]

$$\Gamma(n^3S_1 \rightarrow ggg) = \frac{10\alpha_s^3 2^2 (\pi^2 - 9)}{81\pi M(n^3S_1)^2} |R_{nl}(0)|^2 \times \left[1 - \frac{4.9\alpha_s}{\pi} \right] \quad (25)$$

And the state n^1P_1 , annihilates to the three gluons is given by [41, 82, 113]

$$\Gamma(n^1P_1 \rightarrow ggg) = \frac{20\alpha_s^3 2^4}{9\pi M(n^1P_1)^4} |R'_{nl}(0)|^2 \ln(m_Q \langle r \rangle) \quad (26)$$

In the case of D-wave, $n^3D_{1,3}$ states, the three gluon decay widths are computed using the expression given by [87]

$$\Gamma(n^3D_1 \rightarrow ggg) = \frac{760\alpha_s^3 2^6}{81\pi M(n^3D_1)^6} |R''_{nl}(0)|^2 \ln(4m_Q \langle r \rangle) \quad (27)$$

$$\Gamma(n^3D_3 \rightarrow ggg) = \frac{40\alpha_s^3 2^6}{9\pi M(n^3D_3)^6} |R''_{nl}(0)|^2 \ln(4m_Q \langle r \rangle) \quad (28)$$

The computed three gluon decay widths for S , P and D waves are given in Table 9 along with other model predictions.

3.6 Other annihilation channels of vector bottomonium states

Apart from the decays discussed above, there are other processes by which quarkonium states can annihilate. To elaborate specifically, the decay width of mixed strong and electromagnetic annihilation of n^3S_1 states into γgg photon and two gluons is given by [45]

$$\Gamma(n^3S_1 \rightarrow \gamma gg) = \frac{8(\pi^2 - 9)\alpha\alpha_s^2 e_Q^2 2^2}{9\pi M(n^3S_1)^2} |R_{nl}(0)|^2 \times \left[1 - \frac{7.4\alpha_s}{\pi} \right], \quad (29)$$

and that of $\gamma\gamma\gamma$ is given by [45]

$$\Gamma(n^3S_1 \rightarrow \gamma\gamma\gamma) = \frac{16(\pi^2 - 9)\alpha^3 e_Q^6 2^2}{3M(n^3S_1)^2} |R_{nl}(0)|^2 \times \left[1 - \frac{12.6\alpha_s}{\pi} \right], \quad (30)$$

Also, n^3P_1 state decay into light flavour mesons with a single gluon, the decay rate is computed using the relation

Table 6 The di-leptonic decay widths (in keV) of the S wave and D wave bottomonium states. $\Gamma_{l^+l^-}$ and $\Gamma_{l^+l^-}(C)$ are present results without and with QCD correction respectively

State	J^{PC}	$\Gamma_{l^+l^-}$	$\Gamma_{l^+l^-}(C)$	PDG [43]	[100]	[41]	[101]	[93]	[77]	[103]	[102]
1^3S_1	1^{--}	1.1191	0.7700	1.34 ± 0.018	1.33	0.71	1.20	1.30	1.809	1.3	1.60
2^3S_1	1^{--}	0.7859	0.5442	0.612 ± 0.011	0.62	0.37	0.52	0.76	0.797	0.5	0.64
3^3S_1	1^{--}	0.6171	0.4288	0.443 ± 0.008	0.48	0.27	0.33	0.45	0.618	...	0.44
4^3S_1	1^{--}	0.5096	0.3549	0.272 ± 0.029	0.40	0.21	0.24	0.26	0.541	...	0.35
5^3S_1	1^{--}	0.4350	0.3035	0.18	0.19	0.18	0.29
6^3S_1	1^{--}	0.3695	0.2586	0.15	0.15
1^3D_1	1^{--}	0.0072	0.0050	0.106
2^3D_1	1^{--}	0.0084	0.0058	0.078
3^3D_1	1^{--}	0.0085	0.0059	0.051
4^3D_1	1^{--}	0.0084	0.0058	0.042
5^3D_1	1^{--}	0.0082	0.0057	0.028
6^3D_1	1^{--}	0.0079	0.0055
7^3D_1	1^{--}	0.0076	0.0053

Table 7 The di-gamma decay widths (in keV) of the S wave and P wave bottomonium states. $\Gamma_{\gamma\gamma}$ and $\Gamma_{\gamma\gamma}(C)$ are present results without and with QCD correction respectively. There is no experimental data available for comparison

State	J^{PC}	$\Gamma_{\gamma\gamma}$	$\Gamma_{\gamma\gamma}(C)$	[95]	[41]	[97]	[101]	[102]	[103]	[1]
1^1S_0	0^{-+}	0.3782	0.3035	0.387	0.69	0.738	0.496	0.527	0.35	0.214
2^1S_0	0^{-+}	0.2636	0.2122	0.263	0.36	0.508	0.212	0.263	0.15	0.121
3^1S_0	0^{-+}	0.2067	0.1668	0.229	0.27	0.426	0.135	0.172	0.10	0.906
4^1S_0	0^{-+}	0.1705	0.1378	0.201	0.099	0.105	0.755
5^1S_0	0^{-+}	0.1455	0.1176	0.193	0.078	0.121	...
6^1S_0	0^{-+}	0.1235	0.1000
1^3P_0	0^{++}	0.0811	0.1150	0.0196	0.12	0.050	0.038	0.0208
2^3P_0	0^{++}	0.0717	0.1014	0.0195	0.14	0.037	0.029	0.0227
3^3P_0	0^{++}	0.0620	0.0875	0.0194	0.15	0.037
4^3P_0	0^{++}	0.0545	0.0768	0.0192
5^3P_0	0^{++}	0.0487	0.0686
1^3P_2	2^{++}	0.0213	0.0147	0.0052	0.00308	0.0066	0.008	0.0051
2^3P_2	2^{++}	0.0189	0.0131	0.0052	0.00384	0.0067	0.006	0.0062
3^3P_2	2^{++}	0.0164	0.0114	0.0051	0.00410	0.0064
4^3P_2	2^{++}	0.0144	0.0100	0.0051
5^3P_2	2^{++}	0.0129	0.0090

given by [82]

$$\Gamma(n^3P_1 \rightarrow q\bar{q} + g) = \frac{8\alpha_s^3 n_f 2^4}{9\pi M(n^3P_1)^4} |R'_{nl}(0)|^2 \ln(m_Q(r)) \quad (31)$$

The computed results of these decays are summarised in Table 10 along with other available theoretical and experimental data.

3.7 The electromagnetic transition widths of bottomonium states

Bottomonium states possess more compactness in nature due to relatively heavier mass of the bottom quark. In such situation especially dealing with the radiative transition which are governed by an emission or absorption of gamma photon, the wave length of the photon is either larger or comparable to the size of the radiating bottomonium state. So, one expects radiative transition in $b\bar{b}$ dominates. The leading order elec-

Table 8 The di-gluon decay widths (in MeV) of the S wave and P wave bottomonium states. Γ_{gg} and $\Gamma_{gg}(C)$ are present results without and with QCD correction respectively. There is no experimental data available for comparison

State	J^{PC}	Γ_{gg}	$\Gamma_{gg}(C)$	[95]	[41]	[105]	[104]	[106]
1^1S_0	0^{-+}	5.4496	6.8520	5.448	20.18	17.945	11.49	12.46
2^1S_0	0^{-+}	4.1775	5.2374	3.710	10.64	...	5.16	...
3^1S_0	0^{-+}	3.4499	4.3182	3.229	7.94	...	3.80	...
4^1S_0	0^{-+}	2.9454	3.6829	2.985
5^1S_0	0^{-+}	2.5769	3.2196	2.832
6^1S_0	0^{-+}	2.2859	2.8519
1^3P_0	0^{++}	0.9056	1.4297	0.276	2.00	5.250	0.96	2.15
2^3P_0	0^{++}	0.7855	1.2358	0.275	2.37	...	0.99	...
3^3P_0	0^{++}	0.6713	1.0539	0.273	2.46
4^3P_0	0^{++}	0.5853	0.9175	0.271
5^3P_0	0^{++}	0.5192	0.8127
1^3P_2	2^{++}	0.2384	0.2370	0.073	0.0836	0.822	0.33	0.22
2^3P_2	2^{++}	0.2076	0.2064	0.073	0.1042	...	0.35	...
3^3P_2	2^{++}	0.1778	0.1767	0.072	0.1114
4^3P_2	2^{++}	0.1552	0.1543	0.072
5^3P_2	2^{++}	0.1378	0.1370

Table 9 The Tri-gluon decay widths (in MeV) of the S , P and D wave bottomonium states. Γ_{ggg} and $\Gamma_{ggg}(C)$ are present results without and with QCD correction respectively

State	J^{PC}	Γ_{ggg}	$\Gamma_{ggg}(C)$	PDG [43]	[41]
1^3S_1	1^{--}	0.0400	0.0285	0.0441	0.0416
2^3S_1	1^{--}	0.0269	0.0193	0.0188	0.0242
3^3S_1	1^{--}	0.0206	0.0148	0.0072	0.01876
4^3S_1	1^{--}	0.0168	0.0121
5^3S_1	1^{--}	0.0141	0.0102
6^3S_1	1^{--}	0.0117	0.0085
1^1P_1	1^{+-}	0.0357	0.03526
2^1P_1	1^{+-}	0.0346	0.0527
3^1P_1	1^{+-}	0.0331	0.0621
4^1P_1	1^{+-}	0.0327
5^1P_1	1^{+-}	0.0309
1^3D_1	1^{--}	0.0106	0.0099
2^3D_1	1^{--}	0.0119	0.0096
3^3D_1	1^{--}	0.0118
4^3D_1	1^{--}	0.0113
5^3D_1	1^{--}	0.0108
1^3D_3	3^{--}	0.0060	0.00022
2^3D_3	3^{--}	0.0056	0.00125
3^3D_3	3^{--}	0.0055
4^3D_3	3^{--}	0.0053
5^3D_3	3^{--}	0.0051

Table 10 Some other annihilation decay widths (in keV) of the S and P wave bottomonium states. Γ and $\Gamma(C)$ are present results without and with QCD correction respectively

Transition	Γ	$\Gamma(C)$	PDG [43]	[41]
$1^3S_1 \rightarrow \gamma\gamma\gamma$	4.6775×10^{-4}	1.2312×10^{-4}	...	3.44×10^{-6}
$2^3S_1 \rightarrow \gamma\gamma\gamma$	3.2849×10^{-4}	8.9895×10^{-5}	...	2.00×10^{-6}
$3^3S_1 \rightarrow \gamma\gamma\gamma$	2.5794×10^{-4}	7.2037×10^{-5}	...	1.55×10^{-6}
$4^3S_1 \rightarrow \gamma\gamma\gamma$	2.1298×10^{-4}	6.0336×10^{-5}	...	1.29×10^{-6}
$5^3S_1 \rightarrow \gamma\gamma\gamma$	1.8183×10^{-4}	5.2021×10^{-5}	...	1.10×10^{-6}
$6^3S_1 \rightarrow \gamma\gamma\gamma$	1.5446×10^{-4}	4.4933×10^{-5}	...	9.56×10^{-7}
$1^3S_1 \rightarrow \gamma gg$	1.2730	0.7220	1.18	0.79
$2^3S_1 \rightarrow \gamma gg$	0.8689	0.4982	0.59	0.46
$3^3S_1 \rightarrow \gamma gg$	0.6718	0.3874	0.0097	0.36
$4^3S_1 \rightarrow \gamma gg$	0.5485	0.3176	...	0.30
$5^3S_1 \rightarrow \gamma gg$	0.4646	0.2698	...	0.25
$6^3S_1 \rightarrow \gamma gg$	0.3894	0.2272	...	0.22
$1^3P_1 \rightarrow q\bar{q} + g$	57.9585	71.53
$2^3P_1 \rightarrow q\bar{q} + g$	55.3966	106.14
$3^3P_1 \rightarrow q\bar{q} + g$	52.9585	124.53
$4^3P_1 \rightarrow q\bar{q} + g$	52.4466
$5^3P_1 \rightarrow q\bar{q} + g$	49.5181

Table 11 The root mean square radii (in fm) for bottomonium states

State	$\langle r^2 \rangle^{\frac{1}{2}}$
1S	0.19
2S	0.37
3S	0.53
4S	0.68
5S	0.83
1P	0.27
2P	0.44
3P	0.60
4P	0.75
1D	0.35
2D	0.51
3D	0.67
4D	0.82
5D	0.97

tromagnetic transitions are electric dipole ($E1$) and magnetic dipole ($M1$) transitions.

The selection rules for electric dipole transition ($E1$) are $\Delta l = \pm 1$, $\Delta s = 0$. In contrast, for magnetic dipole transitions ($M1$), $\Delta l = 0$, $\Delta s = \pm 1$. Within non relativistic limit, the decay width of the $E1$ transition from the initial state $n_i^{(2s_i+1)}l_{iJ_i}$ to final state $n_f^{(2s_f+1)}l_{fJ_f}$ can be obtained as [2]

$$\Gamma(i \xrightarrow{E1} f + \gamma) = \frac{4\alpha e_Q^2}{3} (2J' + 1) S_{if}^E \omega^3 |\mathcal{E}_{if}|^2 \times \frac{E_f}{M_i} \quad (32)$$

where

$$\omega = \frac{M_i^2 - M_f^2}{2M_i} \quad (33)$$

is the photon energy. S_{if}^E is the statistical factor and \mathcal{E}_{if} is the overlap integral which can be computed using initial and final state wave functions as

$$\mathcal{E}_{if} = \frac{3}{\omega} \int_0^\infty R_{n\ell}(r) R_{n'\ell'}(r) r^2 dr \times \left[\frac{\omega r}{2} j_0\left(\frac{\omega r}{2}\right) - j_1\left(\frac{\omega r}{2}\right) \right] \quad (34)$$

and

$$S_{if}^E = \max(\ell, \ell') \left\{ \begin{matrix} J & 1 & J' \\ \ell' & s & \ell \end{matrix} \right\}^2 \quad (35)$$

The formula used for the ($M1$) transition from initial state to final state for the quarkonium system can be computed as [88, 89]

$$\Gamma(i \xrightarrow{M1} f + \gamma) = \frac{4\alpha e_Q^2}{3m_Q^2} (2J' + 1) S_{if}^M \times \omega^3 |\langle f | j_0\left(\frac{\omega r}{2}\right) | i \rangle|^2 \times \frac{E_f}{M_i} \quad (36)$$

The statistical factor for the ($M1$) transition is given as

$$S_{if}^M = 6(2s + 1)(2s' + 1)$$

$$\left\{ \begin{matrix} J & 1 & J' \\ s' & l & s \end{matrix} \right\}^2 \left\{ \begin{matrix} 1 & \frac{1}{2} & \frac{1}{2} \\ \frac{1}{2} & s' & s \end{matrix} \right\}^2 \quad (37)$$

It is to be noted that the term E_f/M_i is acting as the relativistic correction factor to the radiative transition width where E_f is the energy of the final state and M_i being the mass of the initial state. The criteria for validity of these transitions to take place within long wave length approximation (LWLA) is specified as $\omega(2\langle r^2 \rangle_f)^{\frac{1}{2}} < 1$ [116]. Table 12 shows calculated values for $\omega(2\langle r^2 \rangle_f)^{\frac{1}{2}}$ using our predicted masses of the bottomonium states. Possible $E1$ and $M1$ transitions including relativistic correction factor and satisfying the LWLA are tabulated in Tables 13 and 14 along with the available theoretical and experimental data for comparison.

4 Results and discussion

The spectroscopic masses of S, P and D waves of the bottomonia are computed based on the instanton induced potential in the non-relativistic frame work with the addition of stronger confinement effects. The present results are compared with available experimental as well as with other model predictions in Table 2 for S wave masses and in Table 3 for P wave and D wave masses. Our results are found to be in very good agreement with the experimental values of the respective states. Our estimation for the 1^3S_1 is 9460.75 MeV which is in excellent agreement with the PDG listed mass of $1^3S_1(9460.30 \pm 0.26)$ MeV. For the case of spin singlet pseudoscalar state 1^1S_0 our finding is 9412.22 MeV which is roughly 13 MeV higher than PDG listed mass (9398.7 ± 2.0 MeV). The mass splitting of 1S state ($1^3S_1 - 1^1S_0$) is found to be 48 MeV which is relatively consistent with the PDG listed mass split for 1S bottomonium (62.3 ± 3.2 MeV). We observe that the S wave mass predictions upto the 4S states are very close to the experimental values (PDG average). Below the $B\bar{B}$ threshold, the spectrum of the bottomonium is well understood. The vector states, n^3S_1 for $n = 1, 2, 3, 4$ are well established states. But there exist discrepancy pertaining to the true identification of $\Upsilon(10,860)$ and $\Upsilon(11,020)$ states. Their status as pure 5^3S_1 and 6^3S_1 of bottomonium states is not clear. It is also evident from the fact that the leptonic decay width (0.31 ± 0.07 keV) of $\Upsilon(10,860)$ is higher than that of $\Upsilon(4S)$ (0.272 ± 0.029 keV) state [43]. The disparity is understood by treating such state of quarkonia as admixture of the S state with the nearby D state. Therefore, identification of these states beyond 4S with $J^{PC} = 1^{--}$ of quarkonia necessitates the considerations of S wave and D wave admixtures.

Such mixed state R_{nJ} in terms of the mixing angle θ can be represented as

$$R_{nJ}(0) = \cos\theta R_{nS}(0) - \sin\theta R_{n'D}(0)$$

Table 12 Photon energy ω (in MeV) calculated from the predicted masses of the states and the LWLA validity factor $\omega(2\langle r^2 \rangle_f)^{\frac{1}{2}}$ for (E1) and (M1) transitions

$i \rightarrow f + \gamma$	ω	$\omega(2\langle r^2 \rangle_f)^{\frac{1}{2}}$
$1^3P_2 \rightarrow 1^3S_1 + \gamma$	411.69	0.7928
$1^3P_1 \rightarrow 1^3S_1 + \gamma$	405.13	0.7801
$1^3P_0 \rightarrow 1^3S_1 + \gamma$	381.18	0.7340
$1^1P_1 \rightarrow 1^1S_0 + \gamma$	451.51	0.8695
$1^3D_1 \rightarrow 1^3P_0 + \gamma$	291.07	0.7965
$1^3D_1 \rightarrow 1^3P_1 + \gamma$	269.83	0.7384
$1^3D_1 \rightarrow 1^3P_2 + \gamma$	260.16	0.7119
$1^3D_2 \rightarrow 1^3P_1 + \gamma$	277.51	0.7594
$1^3D_2 \rightarrow 1^3P_2 + \gamma$	267.85	0.7329
$1^3D_3 \rightarrow 1^3P_2 + \gamma$	273.13	0.7474
$1^1D_2 \rightarrow 1^1P_1 + \gamma$	275.40	0.7536
$2^3S_1 \rightarrow 1^3P_0 + \gamma$	175.05	0.4790
$2^3S_1 \rightarrow 1^3P_1 + \gamma$	153.55	0.4202
$2^3S_1 \rightarrow 1^3P_2 + \gamma$	143.77	0.3934
$2^1S_0 \rightarrow 1^1P_1 + \gamma$	120.18	0.3289
$2^3P_2 \rightarrow 2^3S_1 + \gamma$	245.54	1.0950
$2^3P_1 \rightarrow 2^3S_1 + \gamma$	238.79	1.0649
$2^3P_0 \rightarrow 2^3S_1 + \gamma$	223.82	1.0001
$2^1P_1 \rightarrow 2^1S_0 + \gamma$	270.85	1.2078
$2^3P_2 \rightarrow 1^3S_1 + \gamma$	781.77	1.5054
$2^3P_1 \rightarrow 1^3S_1 + \gamma$	775.38	1.4931
$2^3P_0 \rightarrow 1^3S_1 + \gamma$	761.21	1.4659
$2^1P_1 \rightarrow 1^1S_0 + \gamma$	821.95	1.5828
$3^3S_1 \rightarrow 2^3P_0 + \gamma$	111.50	0.4972
$3^3S_1 \rightarrow 2^3P_1 + \gamma$	96.33	0.4296
$3^3S_1 \rightarrow 2^3P_2 + \gamma$	89.49	0.3990
$3^1S_0 \rightarrow 2^1P_1 + \gamma$	68.76	0.3066
$3^3S_1 \rightarrow 1^3P_0 + \gamma$	502.24	1.3744
$3^3S_1 \rightarrow 1^3P_1 + \gamma$	481.44	1.3175
$3^3S_1 \rightarrow 1^3P_2 + \gamma$	471.98	1.2916
$3^1S_0 \rightarrow 1^1P_1 + \gamma$	454.00	1.2424
$1^3S_1 \rightarrow 1^1S_0 + \gamma$	18.39	0.0932
$1^3P_2 \rightarrow 1^1P_1 + \gamma$	3.69	0.0187
$1^1P_1 \rightarrow 1^3P_1 + \gamma$	1.66	0.0084
$1^1P_1 \rightarrow 1^3P_0 + \gamma$	13.45	0.0681
$2^3S_1 \rightarrow 2^1S_0 + \gamma$	22.71	0.1151
$3^3S_1 \rightarrow 3^1S_0 + \gamma$	27.15	0.1376

The mixing angle θ is determined in terms of the S wave mass and the nearby D wave mass as

$$M_{nJ} = |a|^2 M_{nS} + (1 - |a|^2) M_{n'D}$$

Table 13 The E1 transition decay widths (in keV) of bottomonium states. Γ_{E1} and $\Gamma_{E1}(R)$ are present results without and with relativistic correction factor respectively. Experimental data is taken from [43]

Initial	Final	Γ_{E1}	$\Gamma_{E1}(C)$	PDG [43]	[95]	[109]	[41]	[42]
1^3P_2	1^3S_1	15.7013	15.0471	...	57.530	31.8	39.15	32.8
1^3P_1	1^3S_1	14.6574	14.0602	...	54.927	31.9	35.66	29.5
1^3P_0	1^3S_1	12.5171	12.0327	...	49.530	27.5	28.07	23.8
1^1P_1	1^1S_0	19.4593	18.5864	...	72.094	35.8	43.66	35.7
1^3D_1	1^3P_0	4.4757	4.3472	...	9.670	19.8	20.98	16.5
1^3D_1	1^3P_1	2.6836	2.6123	...	6.313	13.3	12.29	9.7
1^3D_1	1^3P_2	0.1606	0.1564	...	0.394	1.02	0.65	0.56
1^3D_2	1^3P_1	5.2421	5.0988	...	11.489	21.8	21.95	19.2
1^3D_2	1^3P_2	1.5736	1.5321	...	3.583	7.23	6.23	5.6
1^3D_3	1^3P_2	6.6766	6.4976	...	14.013	32.1	24.74	24.3
1^1D_2	1^1P_1	6.8415	6.6567	...	14.821	30.3	17.23	24.9
2^3S_1	1^3P_0	0.3021	0.2968	1.22 ± 0.11	2.377	1.09	1.09	0.91
2^3S_1	1^3P_1	0.6140	0.6045	2.21 ± 0.19	5.689	2.17	1.84	1.63
2^3S_1	1^3P_2	0.8412	0.8291	2.29 ± 0.20	8.486	2.62	2.08	1.88
2^1S_0	1^1P_1	0.8877	0.8770	...	10.181	3.41	2.85	2.48
3^3S_1	2^3P_0	0.2151	0.2128	1.20 ± 0.12	3.330	1.21	1.21	1.03
3^3S_1	2^3P_1	0.4175	0.4137	2.56 ± 0.26	7.936	2.61	2.13	1.91
3^3S_1	2^3P_2	0.5585	0.5536	2.66 ± 0.27	11.447	3.16	2.56	2.30
3^1S_0	2^1P_1	0.4586	0.4555	...	13.981	4.25	2.60	2.96

Table 14 The M1 transition decay widths (in eV) of bottomonium states. Γ_{M1} and $\Gamma_{M1}(C)$ are present results without and with relativistic correction factor respectively. No Experimental data is available comparison

Initial	Final	Γ_{M1}	$\Gamma_{M1}(C)$	[95]	[109]	[41]	[42]
1^3S_1	1^1S_0	3.7986	3.7827	37.668	10	9.34	10
1^3P_2	1^1P_1	0.0197	0.0197	...	0.095	0.089	0.096
1^1P_1	1^3P_1	0.0018	0.0018	...	0.0094	0.0115	0.012
1^1P_1	1^3P_0	0.3191	0.3183	...	0.90	0.86	0.89
2^3S_1	2^1S_0	1.7856	1.7801	5.619	0.59	0.58	0.59
3^3S_1	3^1S_0	1.0341	1.0315	2.849	3.9	0.66	0.25

Table 15 $S - D$ mixing parameters for $\Upsilon(10,860)$ and $\Upsilon(11,020)$ states and their predicted leptonic decay widths

Exp. state	Mixed state configuration	Mixing angle θ	Mixed state di-leptonic width (in keV)	Exp. decay width (in keV)
10,885.2	5^3S_1 and 4^3D_1	Not possible	...	0.31 ± 0.07
10,885.2	5^3S_1 and 5^3D_1	53.79°	0.15	0.31 ± 0.07
10,885.2	5^3S_1 and 6^3D_1	39.98°	0.25	0.31 ± 0.07
11,000.0	6^3S_1 and 5^3D_1	51.69°	0.14	0.13 ± 0.03
11,000.0	6^3S_1 and 6^3D_1	Not possible	...	0.13 ± 0.03
11,000.0	6^3S_1 and 7^3D_1	Not possible	...	0.13 ± 0.03

Table 16 Spin average masses and mass splittings of bottomonium states compared with PDG [43] collected data

State	Mass _{Present}	Mass _{Exp} [43]
$1\bar{S}$	9448.61	9444.97
$2\bar{S}$	10018.53	10017.19
$1\bar{P}-1\bar{S}$	425.95	454.9
$1^3S_1 - 1^1S_0$	48.53	62.3 ± 3.2
$2^3S_1 - 2^1S_0$	30.74	$24.3 \pm 3.5^{+2.8}_{-1.9}$
$3^3S_1 - 2^3S_1$	338.43	$331.50 \pm 0.02 \pm 0.13$

Here, $|a|^2 = \cos^2\theta$. The value of $|a|^2$ gives the probability of being in the M_{nS} state while $(1 - |a|^2)$ gives the probability of being in the $M_{n'D}$ state with mixing angle θ . Table 15 shows possible configuration and the corresponding mixing angles for $\Upsilon(10,860)$ and $\Upsilon(11,020)$ states. For some cases mixing probability is greater than one, they are listed as “not possible” mixing configuration. Following this, $\Upsilon(10,860)$ with a mass of 10,885.2 MeV as reported by PDG is found to be the admixture of 5^3S_1 and 6^3D_1 states with the mixing angle $\theta = 39.98^\circ$. And the di-leptonic decay width of this mixed state is estimated as 0.25 keV which is very close to the PDG reported value of 0.31 ± 0.07 keV rather than the predicted 0.43 keV of pure 5^3S_1 state. Similarly, $\Upsilon(11,020)$ of mass 11,000.0 MeV listed by PDG is found to be the admixture of 6^3S_1 and 5^3D_1 with mixing angle $\theta = 51.69^\circ$ and di-leptonic decay of the mixed state is obtained as 0.14 keV which is in excellent agreement with the PDG listed value of 0.13 ± 0.03 keV. Fine agreements are also seen in the case of P wave states and the D waves with respect to the existing experimental values.

In the Table 16 we have compared the spin average masses and various mass splitting with PDG reported values. The spin average masses of the low-lying states $1\bar{S}$ from present study is 9448.61 MeV which is very close to 9444.97 MeV as observed from PDG listed masses of 1^3S_1 and 1^1S_0 . Considering 2^1S_0 mass of 9999.0 ± 3.5 MeV predicted by Belle Collaboration and its 2^3S_1 mass listed in PDG, we get the $2\bar{S}$ mass as 10,018.53 MeV which is very close to the predicted $2\bar{S}$ mass of 10,017.19 MeV. Similarly, the centroid masses

of the low-lying P- states listed in Table 17 are also in good accord with the values deduced from experimental values.

It is found that the modification incorporated to the instanton induced potential to make it a smoothly varying function of r as well as introducing a strong confinement part has led to the successful predictions of the masses of various states as well as the right amount of hyperfine splitting between the 3S_1 and 1S_0 states and the P and D wave splitting. It has also been observed that by the subtraction of 61.81 MeV from Eq. (2) (short range part of potential) instead of adding to the long range part of potential though removes the discontinuity at $r = 0.36$ fm but the predictions are not satisfactory as compared to the experimental values even after refitting of the parameters. It can be seen from the Table 11 for the root mean square radii, the rms radii of $1S$, $2S$ and $1P$ are nearer to 0.36 fm. So, this short range part of the potential though is small but not negligible for these low lying states. It is to be noticed that even though the contribution from short range part of the potential (Eq. (2)) is small compared to other components of the potential, their contributions to the hyperfine splitting through the variational parameter μ is not negligible. For instance, if we include only the long range and confinement part in our calculations then in that case, the hyperfine contribution is only 8 MeV which is very much lower than the experimental hyperfine mass splitting of 62.3 ± 3.2 MeV between $\Upsilon(1S)$ and $\eta_b(1S)$ states. So, short range part of the potential is crucial for the prediction of spin dependent contributions. Thus, the full potential (short range plus long range plus confinement) part is important for the determination of the wave functions of the hadronic states. These wave functions are crucial for the predictions of the hyperfine, fine structure as well as the decay properties of the mesonic states.

As far as the identification of the excited bottomonium states are concerned, the masses, decay constant, di gamma, di leptonic decay width are primarily important. Comparing the results of the vector decay constants by our formalism with the experimental result of PDG [43] we found that our results are fairly in good agreement. For the orbitally excited states our results with the radiative corrections are in very good agreement with PDG [43] reported values. For instance, the vector decay constant of 3S state, our predicted value of

Table 17 The theoretical masses of the ground and few excited states of P - wave states are compared with the spin-average centroid masses of the corresponding experimental data reported by PDG [43]

State	Mass (n^1P_1) _{Present}	Mass (n^1P_1) _{Exp} [43]	Mass (n^3P_J) _{Present}	Mass (n^3P_J) _{Exp} [43]
1P	9874.56	9899.3 ± 0.8	9874.56	9899.87 ± 0.27
2P	10,270.00	$10,259.8 \pm 1.2$	10,270.00	$10,260.20 \pm 0.36$
3P	10,526.50	...	10,526.50	$10,534 \pm 9$ $10,551 \pm 1$ [117] $10,530 \pm 4$ [118]

430.42 MeV is very close to the value reported by PDG [43] (430 ± 4 MeV). This indicates that radiative corrections play an important role in the decay mechanism.

In the case of pseudoscalar decay constant, due to the unavailability of PDG data we have compared our data set with the results of other theoretical model [95], Lattice QCD [98] and QCD sum rules [99]. We found our results in good agreement with the Cornell potential predictions [95]. Also in the case of Lattice QCD, the pseudoscalar decay constant for 1^1S_0 state is (667 MeV) which is in close agreement with our prediction (654.81 MeV). One can see from the Table 5 that predictions from QCD sum rule [99] is roughly three times lower than all other estimations for 1^1S_0 state. Also, the results in the case of [97] (potential model study) are considerably higher than other estimations of the pseudoscalar decay constant. It is important to note that both the vector and the pseudoscalar decay constants rely upon the numerical value of the square of the wave function at the origin, so the choice of the potential and parameters may affect on the predictions. We have also computed the Di-lepton, Di-gamma and Di-gluon decay widths and results are summarized in Table 6, 7 and 8. Our results are compared with the respective values reported in PDG [43]. In the case of di-leptonic decay of the n^3S_1 it is observed that the radiative corrections are important for higher ($n > 2$) radial excited states. The results predicted by [100] (Relativistic potential model) are much closer to our predictions with radiative corrections. Besides that our predicted results with radiative correction are also in a close agreement with [102] (Screened potential model). In general our results are comparable with other predictions. We have been able to compute the n^3D_1 annihilation into electron positron pair but we don't find more estimations in literature for its comparison.

The di-gamma decay widths are consistent with the outcomes of Cornell potential model [95] for pseudoscalar n^1S_0 states. It is found that the predictions reported by [97] are almost two times higher than all the other reported values. For the n^3P_0 states annihilation into $\gamma\gamma$, our results are comparable with other predictions. Our predictions are found to be higher than other results in the case of n^3P_2 states.

For n^3S_0 decays into gg our estimated results without radiative correction are consistent with [95] and considerably lower than that from the [41, 104–106]. For n^3P_0 states, our computed results agree well with the [104] without QCD correction while they are in accordance with results of [41, 106] with radiative correction. For these states estimations of [95] are comparable with our predictions. For the n^3P_2 state, our predictions are in line with that of [106] and [104]. For the S wave vector state decaying into ggg our results are in agreement with the decay width listed by PDG [43]. Also they are comparable with [41] especially for 1^3S_1 state where our prediction is 0.040 MeV without radiative correction while their prediction is 0.041 MeV. For the n^1P_1 , n^3D_1 and 1^3D_3

we do not include the radiative correction because such corrections are very small. Interestingly, for 1^1P_1 stat, our predicted tri-gluon decay width is 0.035 MeV which is the same as reported by [41].

Table 10 summaries some of the other annihilation decay of the n^3S_1 into $\gamma\gamma\gamma$ and γgg . As for the $\gamma\gamma\gamma$ is concerned, we find that present results are considerably higher than [41]. For γgg it is again consistent with the PDG [43] and [41]. The radiatively corrected results are almost matching with that of [41]. Results for $n^3P_1 \rightarrow q\bar{q} + g$ are few keV lower than [41] but still comparable. For all these decays one has to wait for the experimental confirmation. In general, We can conclude that our instanton potential predictions and those from the constituent quark model predictions [41] are in a good agreement for the ggg as well as γgg decays.

Table 12 shows the photon energy computed using predicted masses of the bottomonia states and the corresponding long wavelength approximation (LWLA) validity factor for various electromagnetic transitions. $\omega(2\langle r^2 \rangle_f)^{\frac{1}{2}}$. From the table one can find that for transitions $2P \rightarrow 1S$, $2P \rightarrow 2S$ and $3S \rightarrow 1P$ the value of $\omega(2\langle r^2 \rangle_f)^{\frac{1}{2}}$ is greater than 1. So, for these transition long wavelength approximation is not applicable and we have not computed their transition widths here. Very recently, Bruschini and Gonzalez [116] worked on the radiative decays in bottomonium beyond the long wavelength approximation and corresponding study are underway. We present some of the allowed electric dipole transitions (E1) and magnetic dipole transitions (M1) as per LWLA in Tables 13 and 14 respectively. Our predicted transition widths are compared with the available other predictions and experimental values wherever it is available. Looking to the results, we find that for $1P \rightarrow 1S$ results are much lower than that reported in [95]. And similarly we find that our results deviates from [95] for every transitions except from $1D \rightarrow 1P$ where the predictions do agree. On the experimental side, transition width from $1P \rightarrow 1S$, $1D \rightarrow 1P$ is not yet listed in PDG. For the magnetic dipole transitions our predictions are more or less comparable with other data set. The discrepancies in the theoretical predictions may vary from one model to the other due to the choice of potential which plays an important role in the predictions of mass which in turn creates differences in phase space which affects the transition widths. Also sometimes it is a choice of different wave function which effects the predictions of the transition widths. To improve the results one need to consider either the numerical solutions or the perturbative effects on the wave functions for the computation of the electromagnetic transition form factors. Also, instanton vacuum potential plays a vital role in obtaining mass spectroscopy and other relevant properties of bottomonium. Present study re-iterates the importance of the addition of confinement in heavy sector while employing

instanton effects as highlighted by Schäfer and Shuryak in their review paper [58].

Finally, we hope that our predicted results using instanton effects on heavy quarks will be useful in the identification of new quarkonium states that will be observed in future experiments.

Acknowledgements We acknowledge the partial support from DST-SERB, India through the major research project: (SERB/F/8749/2015-16).

Data Availability Statement This manuscript has no associated data or the data will not be deposited. [Authors' comment: All data generated or analyzed during this study are included in this published article and the paper has no associated data since it is theoretical work.]

Open Access This article is licensed under a Creative Commons Attribution 4.0 International License, which permits use, sharing, adaptation, distribution and reproduction in any medium or format, as long as you give appropriate credit to the original author(s) and the source, provide a link to the Creative Commons licence, and indicate if changes were made. The images or other third party material in this article are included in the article's Creative Commons licence, unless indicated otherwise in a credit line to the material. If material is not included in the article's Creative Commons licence and your intended use is not permitted by statutory regulation or exceeds the permitted use, you will need to obtain permission directly from the copyright holder. To view a copy of this licence, visit <http://creativecommons.org/licenses/by/4.0/>. Funded by SCOAP³.

References

1. S. Godfrey, N. Isgur, Phys. Rev. D **32**, 189 (1985)
2. E. Eichten, K. Gottfried, T. Kinoshita, K.D. Lane, T.M. Yan, Phys. Rev. D **17**, 3090 (1978)
3. A. Garmash, EPJ Web Conf. **96**, 01014 (2015)
4. E. Eichten, S. Godfrey, H. Mahlke, J.L. Rosner, Rev. Mod. Phys. **80**, 1161 (2008)
5. S. Herb et al., Phys. Rev. Lett. **39**, 252 (1977)
6. W.R. Innes et al., Phys. Rev. Lett. **39**, 1240 (1977)
7. K. Han et al., Phys. Rev. Lett. **49**, 1612 (1982)
8. G. Eigen et al., Phys. Rev. Lett. **49**, 1616 (1982)
9. C. Klopfenstein et al., Phys. Rev. Lett. **51**, 160 (1983)
10. F. Pauss et al., Phys. Lett. B **130**, 439 (1983)
11. D. Lovelock et al., Phys. Rev. Lett. **54**, 377 (1985)
12. D. Besson et al. (CLEO), Phys. Rev. Lett. **54**, 381 (1985)
13. B. Aubert et al. (BABAR), Phys. Rev. Lett. **101**, 071801 (2008)
14. G. Bonvicini et al. (CLEO), Phys. Rev. D **81**, 031104 (2010)
15. R. Mizuk et al. (Belle), Phys. Rev. Lett. **109**, 232002 (2012)
16. J. Lees et al. (BABAR), Phys. Rev. D **84**, 072002 (2011)
17. S. Dobbs, Z. Metreveli, K. Seth, A. Tomaradze, T. Xiao, Phys. Rev. Lett. **109**, 082001 (2012)
18. S. Sandilya et al. (Belle), Phys. Rev. Lett. **111**, 112001 (2013)
19. G. Bonvicini et al. (CLEO), Phys. Rev. D **70**, 032001 (2004)
20. P. del Amo Sanchez et al. (BABAR), Phys. Rev. D **82**, 111102 (2010)
21. J. Lees et al. (BABAR), Phys. Rev. D **84**, 091101 (2011)
22. R. Aaij et al. (LHCb), J. High Energy Phys. **10**, 005 (2014)
23. A. Bondar et al. (Belle), Phys. Rev. Lett. **108**, 122001 (2012)
24. T. Guo, L. Cao, M. Zhou, H. Chen, [arXiv:1106.2284](https://arxiv.org/abs/1106.2284)
25. I. Danilkin, V. Orlovsky, Y. Simonov, Phys. Rev. D **85**, 034012 (2012)
26. C. Hughes, R.J. Downdall, C.T.H. Davies, R.R. Horgan, G. von Hippel, M. Wingate, Phys. Rev. D **84**, 094501 (2015)
27. D. Becirevi, M. Kruse, F. Sanfilippo, J. High Energy Phys. **05**, 014 (2011)
28. R. Lewis, R.M. Woloshyn, Phys. Rev. D **84**, 094501 (2011)
29. M. Baker, A.A. Penin, D. Seidel, N. Zerf, Phys. Rev. D **92**, 054502 (2015)
30. N. Brambilla, Y. Jia, A. Vairo, Phys. Rev. D **73**, 054005 (2006)
31. N. Brambilla, P. Pietrulewicz, A. Vairo, Phys. Rev. D **85**, 094005 (2012)
32. A. Pineda, J. Segovia, Phys. Rev. D **87**, 074024 (2013)
33. H.M. Choi, Phys. Rev. D **75**, 073016 (2007)
34. H.W. Ke, X.Q. Li, Z.T. Wei, X. Liu, Phys. Rev. D **82**, 034023 (2010)
35. H.W. Ke, X.Q. Li, Y.L. Shi, Phys. Rev. D **87**, 054022 (2013)
36. H.W. Ke, X. Q. Li, X. Liu, [arXiv:1002.1187](https://arxiv.org/abs/1002.1187)
37. J. Ferretti, E. Santopinto, Phys. Rev. D **90**, 094022 (2014)
38. J. Ferretti, G. Galata, E. Santopinto, Phys. Rev. D **90**, 054010 (2014)
39. Y. Lu, M.N. Anwar, B.S. Zou, Phys. Rev. D **94**, 034021 (2016)
40. F. Fazio, Phys. Rev. D **79**, 054015 (2009)
41. J. Segovia, P.G. Ortega, D.R. Entem, F. Fernández, Phys. Rev. D **93**, 074027 (2016)
42. S. Godfrey, K. Moats, Phys. Rev. D **92**, 054034 (2015)
43. P.A. Zyla et al., Prog. Theor. Exp. Phys. **2020**, 083C01 (2020)
44. W. Lucha, F. Schöberl, D. Gromes, Phys. Rep. (Rev. Sect. Phys. Lett.) **4**, 200 (2018)
45. W. Kwong, J.L. Rosner, C. Quigg, Annu. Rev. Nucl. Part. Sci. **37**, 325 (1987)
46. N. Brambilla et al., Eur. Phys. J. C **71**, 1534 (2011)
47. J.J. Dudek, R.G. Edwards, N. Mathur, D.G. Richards, Phys. Rev. D **77**, 034501 (2008)
48. S. Meinel, Phys. Rev. D **79**, 094501 (2009)
49. A. Gray, I. Allison, C.T.H. Davies, E. Gulez, G.P. Lepage, J. Shigemitsu, M. Wingate, Phys. Rev. D **72**, 094507 (2005)
50. T. Burch et al., Phys. Rev. D **81**, 034508 (2010)
51. S. Meinel, Phys. Rev. D **82**, 114502 (2010)
52. R. Lewis, R. Woloshyn, Phys. Rev. D **85**, 114509 (2012)
53. M. Wurtz, R. Lewis, R.M. Woloshyn, Phys. Rev. D **92**, 054504 (2015)
54. R. Dowdall et al., Phys. Rev. D **85**, 054509 (2012)
55. E. Eichten, K. Gottfried, T. Kinoshita, J.B. Kogut, K.D. Lane, T.M. Yan, Phys. Rev. Lett. **34**, 369 (1975)
56. A.A. Belavin, A.M. Polyakov, A.S. Schwartz, Y.S. Tyupkin, Phys. Lett. B **59**, 85 (1975)
57. U.T. Yakhshiev, H.C. Kim, B. Turimov, M.M. Musakhanov, Chin. Phys. C **41**, 083102 (2017)
58. T. Schäfer, E.V. Shuryak, Instantons in QCD. Rev. Mod. Phys. **70**(2) (1998)
59. F. Wilczek, A. Zee, Phys. Rev. Lett. **40**, 83 (1978)
60. C.G. Callan, R.F. Dashen, D.J. Gross, F. Wilczek, A. Zee, Phys. Rev. D **18**, 4684 (1978)
61. E. Eichten, F. Feinberg, Phys. Rev. D **23**, 2724 (1981)
62. D. Diakonov, V.Y. Petrov, P.V. Pobylitsa, Phys. Lett. B **226**, 372 (1989)
63. D. Diakonov, V.Y. Petrov, Nucl. Phys. B **245**, 259 (1984)
64. D. Diakonov, Prog. Part. Nucl. Phys. **51**, 173 (2003), and references therein
65. P. Pobylitsa, Phys. Lett. B **226**, 387 (1989), and references therein
66. D. Diakonov, V.Y. Petrov, Nucl. Phys. B **272**, 457 (1986)
67. D. Diakonov, Prog. Part. Nucl. Phys. **226** (1989)
68. E.V. Shuryak, Nucl. Phys. B **203**, 93 (1982)
69. H.-Ch. Kim, M. Musakhanov, M. Siddikov, Phys. Lett. B **633**, 701 (2006)

70. K. Goetze, H.-Ch. Kim, M. Musakhanov, M. Siddikov, Phys. Rev. D **76**, 116007 (2007)
71. K. Goetze, M. Musakhanov, M. Siddikov, Phys. Rev. D **76**, 076007 (2007)
72. T.A. DeGrand, Phys. Rev. D **64**, 094508 (2001)
73. P. Faccioli, T.A. DeGrand, Phys. Rev. Lett. **91**, 182001 (2003)
74. R. Van Royen, V.F. Weisskopf, Nuovo Cim A **50**, 617 (1967)
75. E. Braaten, S. Fleming, Phys. Rev. D **52**, 181 (1995)
76. A.V. Berezhnoy, V.V. Kiselev, A.K. Likhoded, Z. Phys. A **356**, 89 (1996)
77. P.C. Vinodkumar, K.B. Vijayakumar, S.B. Khadkikar, Pramana J. Phys. **39**, 47 (1992)
78. A.P. Monterio, K.B. Vijayakumar, Nat. Sci. **2**, 1292 (2010)
79. J.L. Rosner et al. (CLEO Collaboration), Phys. Rev. Lett. **96**, 092003 (2006)
80. S. Bhatnagar, S. Li, J. Phys. G Nucl. Part. Phys. **32**, 949 (2006)
81. A. Badalyan, B. Ioffe, A.V. Smilga, Nucl. Phys. B **281**, 85 (1987)
82. W. Kwong, P.B. Mackenzie, R. Rosenfeld, Phys. Rev. D **37**, 3210 (1988)
83. R. Barbieri, R. Gatto, R. Kögerler, Z. Kunszt, Phys. Lett. B **57**, 455 (1975)
84. V.A. Novikov, L.B. Okun, M.A. Shifman, A.I. Vainshtein, M.B. Voloshin, V.I. Zakharov, Phys. Rep. **41**, 1 (1978)
85. J.P. Lansberg, T.N. Pham, Phys. Rev. D **79**, 094016 (2009)
86. R. Barbieri, M. Caffo, R. Gatto, E. Remiddi, Nucl. Phys. B **192**, 61 (1981)
87. G. Belanger, P. Moxhay, Phys. Lett. B **199**, 575 (1987)
88. S.F. Radford, W.W. Repko, Phys. Rev. D **75**, 074031 (2007)
89. T.A. Lahde, Nucl. Phys. A **714**, 183 (2003)
90. B. Colquhoun, R.J. Dowdall, C. Davies, K. Hornbosteland, G.P. Lepage, Phys. Rev. D **91**(7), 074514 (2015)
91. G.L. Wang, Phys. Lett. B **633**, 492 (2006)
92. B. Patel, P.C. Vinodkumar, J. Phys. G Nucl. Part. **36**, 035003 (2009)
93. T. Bhavsar, M.N. Shah, P.C. Vinodkumar, Eur. Phys. J. C **78**, 227 (2018)
94. K. Bhaghyesh, B. Vijaya Kumar, A.P. Monterio, Phys. G. Nucl. Part. **38**, 085001 (2011)
95. N.R. Soni, B.R. Joshi, R.P. Shah, H.R. Chauhan, J.N. Pandya, Eur. Phys. J. C **78**, 592 (2018)
96. A. Krassnigg, M. Gomez-Rocha, T. Hilger, J. Phys. Conf. Ser. **742**, 012032 (2016)
97. H. Negash, S. Bhatnagar, Int. J. Mod. Phys. E **25**(8), 1650056 (2016)
98. C. McNielle, C. Davies E. Follana, K. Hornbostel, G.P. Lepage (HPQCD Collab.), Phys. Rev. D **86**, 074503 (2012). [arXiv:1207.0994](https://arxiv.org/abs/1207.0994) [he-lat]
99. E. Veli Veliev, K. Azizi, H. Sundu, N. Aksit, J. Phys. G Nucl. Part. **39**, 015002 (2012)
100. S.F. Radford, W.W. Repko, Nucl. Phys. A **865**, 69 (2011)
101. M. Shah, A. Parmar, P.C. Vinodkumar, Phys. Rev. D **86**, 034015 (2012)
102. B.Q. Li, K.T. Chao, Commun. Theor. Phys. **52**, 653 (2009)
103. D. Ebert, R.N. Faustov, V.O. Galkin, Mod. Phys. Lett. A **18**, 601 (2003)
104. J.T. Lavery, S.F. Radford, W.W. Repko, [arXiv:0901.3917](https://arxiv.org/abs/0901.3917) (2009)
105. A. Parmar, B. Patel, P.C. Vinodkumar, Nucl. Phys. A **848**, 299 (2010)
106. S.N. Gupta, J.M. Johnson, W.W. Repko, Phys. Rev. D **54**, 2075 (1996)
107. C.S. Fischer, S. Kubrak, R. Williams, Eur. Phys. J. A **51**, 10 (2015)
108. D. Ebert, R.N. Faustov, V.O. Galkin, Eur. Phys. J. C **71**, 1825 (2011)
109. W.J. Deng, H. Liu, L.C. Gui, X.H. Zhong, Phys. Rev. D **95**, 074002 (2017)
110. M. Shah, B. Patel, P.C. Vinodkumar, Eur. Phys. J. C **76**, 36 (2016)
111. M. Shah, B. Patel, P.C. Vinodkumar, Phys. Rev. D **93**, 094028 (2016)
112. M. Shah, B. Patel, P.C. Vinodkumar, Phys. Rev. D **90**, 014009 (2014)
113. W. Kwong, J.L. Rosner, Phys. Rev. D **38**, 279 (1988)
114. A. Bradley, A. Khare, Z. Phys. C Part. Fields **8**, 131 (1981)
115. F.E. Close, *An Introduction to Quarks and Partons* (Academic Press, London, 1979)
116. R. Bruschini, P. Gonzalez, Phys. Rev. D **100**, 074001 (2019)
117. V.M. Abazov et al. (DØ), Phys. Rev. D **86**, 031103 (2012)
118. G. Aad et al. (ATLAS Collaboration), Phys. Rev. Lett. **108**, 152001 (2012)