

Quantum gravity framework for the construction of algebraic structure of cosmic spacetime

Weitao Xiao *

Ocean University of China, Qingdao 266000, China

*Corresponding author's e-mail: 894311580@qq.com

Abstract: This paper concentrates on building the algebraic structure of cosmic spacetime within the quantum gravity framework. By merging the theoretical bases of quantum mechanics and general relativity, we examine the algebraic characteristics of spacetime at the quantum level. We present algebraic structures like non-commutative geometry and Hopf algebras, and derive the related algebraic relations for spacetime coordinates and momentum operators. Through detailed analysis, we introduce several key tables and formulas to describe the algebraic structure, exploring its physical implications and potential applications in comprehending spacetime nature and universe evolution.

1. Introduction

Combining quantum mechanics with general relativity stands as one of the crucial challenges in contemporary theoretical physics. Quantum gravity seeks to unify these theories, with the algebraic structure of spacetime being a key focus in this area. At the quantum level, spacetime is anticipated to display characteristics distinct from the classical smooth spacetime outlined by general relativity. Algebraic structures offer a robust mathematical framework for describing the quantum-scale properties of spacetime ^[1].

In classical general relativity, spacetime constitutes a smooth manifold equipped with a Lorentzian metric. However, in the quantum realm, the notion of a smooth spacetime might fail, requiring algebraic methods to describe spacetime instead. Non-commutative geometry, for instance, posits that spacetime coordinates might fail to commute at extremely small scales. This can be characterized by an algebraic framework in which coordinate operators adhere to specific non-commutative relationships ^[2].

2. Theoretical Foundations

2.1. Quantum Mechanics and General Relativity

In quantum mechanics, physical observables are denoted by operators, and the system's state is characterized by a wavefunction within a Hilbert space. The commutation relations among operators are crucial in defining the system's properties. For instance, in one-dimensional space, the position and momentum operators satisfy the commutation relation $[\hat{x}, \hat{p}] = i\hbar$, where \hat{x} denotes the position operator, \hat{p} represents the momentum operator, and \hbar is the reduced Planck constant ^[3].

In general relativity, the presence of mass and energy curves spacetime, with the curvature characterized by the Riemann curvature tensor. The Einstein field formulas connect spacetime curvature to the energy-momentum tensor of matter and radiation. However, when attempting to quantize gravity, we face challenges as the classical theory of general relativity does not align directly with quantum mechanics principles ^[4].



2.2. Algebraic Structures in Quantum Gravity

Various algebraic structures have been suggested within the framework of quantum gravity. An important example is the Hopf algebra, which integrates the structures of both an algebra and a coalgebra and serves to describe symmetries and interactions in quantum gravity. Another significant algebraic structure is non-commutative geometry, in which spacetime coordinates are substituted with non-commutative operators, and the function algebra on spacetime forms a non-commutative algebra [5].

2.3. Quantum Gravity Challenges: A Deeper Dive

The conflict between quantum mechanics and general relativity is most evident in the issue of ultraviolet (UV) divergences during the quantization of the gravitational field. In perturbative quantum gravity, the Feynman diagrams for gravitational interactions show divergent behavior at high energies. This divergence cannot be renormalized because of the dimensionality of the Newtonian constant G_N (with dimension $L^2 T^{-2} M^{-1}$ in natural units), rendering the theory non-renormalizable. This stands in stark contrast to the renormalizable theories within the Standard Model, where coupling constants exhibit advantageous dimensional characteristics.

Another fundamental issue stems from the ontological discrepancy between the classical spacetime manifold and the quantum Hilbert space framework. General relativity portrays spacetime as a smooth, differentiable manifold with local Lorentz invariance, whereas quantum mechanics demands that physical states exist within a separable Hilbert space and evolve unitarily in time. Quantizing the metric tensor $g_{\mu\nu}$ as an operator field raises conceptual issues, like the problem of time in canonical quantum gravity. Here, the Hamiltonian constraint $H\Psi = 0$ removes the classical time parameter from the Wheeler-DeWitt formula.

2.4. Mathematical Formalism of Hopf Algebras

H is a bialgebra (possessing both algebra and coalgebra structures) endowed with an antipode map $S: H \rightarrow H$ that acts like an “inverse” for the operations. It satisfies the axiom:

$$m \circ (id \otimes S) \circ \Delta = m \circ (S \otimes id) \circ \Delta = \eta \circ \epsilon \quad (1)$$

where m denotes multiplication, Δ represents comultiplication, η signifies the unit, and ϵ indicates the counit. Within the framework of quantum spacetime, the Poincaré Hopf algebra alters the classical Poincaré algebra through the introduction of q -deformations in the commutation relations [6]. For instance, the modified commutation relation between boost generators K_i and rotation generators J_j can be expressed as:

$$[K_i, J_j] = i\epsilon_{ijk} K_k \quad (\text{classical}) \quad \text{vs.} \quad [K_i, J_j] = i\epsilon_{ijk} K_k + \alpha P_i P_j \quad (\text{quantum deformed}) \quad (2)$$

where α represents a deformation parameter associated with the Planck scale. This deformation captures the quantum essence of spacetime symmetries, with momentum-dependent corrections arising at ultra-small scales.

3. Construction of the Algebraic Structure of Spacetime

3.1. Non-Commutative Spacetime

We assume that at the quantum scale, the spacetime coordinates are non-commutative operators \hat{x}^μ ($\mu = 0, 1, 2, 3$ for the four spacetime dimensions) for the four spacetime dimensions). The commutation relations among these coordinate operators are expressed as:

$$[\hat{x}^\mu, \hat{x}^\nu] = i\theta^{\mu\nu} \quad (3)$$

where $\theta^{\mu\nu}$ denotes a constant antisymmetric tensor describing the non-commutativity of spacetime. This non-commutativity suggests that the sequence of measuring spacetime coordinates is significant at the quantum level.

In such a non-commutative spacetime, the momentum operators \hat{p}_μ satisfy specific commutation relations with the coordinate operators. By drawing an analogy with the commutative case, we can propose that:

$$[\hat{x}^\mu, \hat{p}_\nu] = i\delta_\nu^\mu \tag{4}$$

where δ_ν^μ denotes the Kronecker delta.

3.2. Algebraic Relations for Spacetime Fields

In quantum field theory on non-commutative spacetime, the fields depend on the non-commutative coordinates. The Moyal product replaces the product of two fields $\phi(x)$ and $\psi(x)$, and is defined as:

$$(\phi \star \psi)(x) = \phi(x) e^{\frac{i}{2}\theta^{\mu\nu} \partial_\mu \partial'_\nu} \psi(x')|_{x'=x} \tag{5}$$

This Moyal product captures the non-commutativity of spacetime and alters the interaction terms in the field formulas.

3.3. Hopf Algebra Structure

We can also introduce a Hopf algebra structure to depict the symmetries of quantum spacetime. The Hopf algebra includes a multiplication operation Δ (comultiplication), a counit ϵ , and an antipode S . For instance, in the context of the Poincaré Hopf algebra, the comultiplication of the momentum generator P_μ is expressed as:

$$\Delta(P_\mu) = P_\mu \otimes 1 + 1 \otimes P_\mu \tag{6}$$

This describes the distribution of momentum across different parts of the system.

3.4. Representation Theory of Non-Commutative Spacetime

To build a concrete representation of the non-commutative spacetime algebra, we examine the Hilbert space. $H=L^2(\mathbb{R}^4)$ of square-integrable functions on Minkowski spacetime. The coordinate operators x^μ act as multiplication operators:

$(x^\mu \psi)(x) = x^\mu \psi(x)$, while the momentum operators act as differential operators: $(p^\mu \psi)(x) = -i\partial_\mu \psi(x)$. In the non-commutative case, these operators satisfy the canonical commutation relations (CCR) deformed by the antisymmetric tensor $\theta_{\mu\nu}$: $[x^\mu, x^\nu] = i\theta^{\mu\nu}$, $[x^\mu, p^\nu] = i\delta^\nu_\mu$, $[p^\mu, p^\nu] = 0$.

This algebra is isomorphic to a Heisenberg algebra extended to four dimensions, with $\Theta_{\mu\nu}$ playing the role of a non-commutative ‘‘area’’ parameter. For consistency with Lorentz invariance, $\theta_{\mu\nu}$ is typically taken to be a constant tensor, leading to the so-called ‘‘constant non-commutativity’’ scenario.

3.5. Algebraic Curvature Formalism

In the algebraic spacetime framework, the Riemann curvature tensor $R_{\nu\rho\sigma\mu}$ is no longer a purely geometric object but incorporates algebraic contributions from the non-commutative structure [7]. Using the connection $\omega_\mu = \Gamma_{\nu\rho\mu} dx^\rho$ (where $\Gamma_{\nu\rho\mu}$ are the Christoffel symbols), the curvature two-form is defined via the Maurer-Cartan formula:

$$R_{\nu\mu} = d\omega_\nu + \omega_\rho \wedge \omega_\sigma \tag{7}$$

In non-commutative spacetime, the exterior derivative d and wedge product \wedge are replaced by their non-commutative counterparts, leading to modified curvature expressions that include terms proportional to $\theta_{\mu\nu} \partial_\mu \partial_\nu$. Explicitly, the curvature component becomes:

$$R_{\nu\rho\sigma\mu} = \partial_\rho \Gamma_{\nu\sigma\mu} - \partial_\sigma \Gamma_{\nu\rho\mu} + \Gamma_{\rho\lambda\mu} \Gamma_{\nu\sigma\lambda} - \Gamma_{\sigma\lambda\mu} \Gamma_{\nu\rho\lambda} + 2i\theta^{\alpha\beta} (\partial_\alpha \Gamma_{\nu\rho\mu} \partial_\beta \Gamma_{\nu\sigma\lambda} - \partial_\beta \Gamma_{\nu\rho\mu} \partial_\alpha \Gamma_{\nu\sigma\lambda}) \tag{8}$$

This algebraic curvature tensor captures both classical geometric effects and quantum non-commutative corrections [8].

4. Key Tables and Formulas

Table 1. Comparison of Commutation Relations in Different Spacetime Models.

Spacetime Model	Commutation Relation $[\hat{x}^\mu, \hat{x}^\nu]$	Commutation Relation $[\hat{x}^\mu, \hat{p}_\nu]$
Commutative Spacetime	0	$i\delta_\nu^\mu$
Non-Commutative Spacetime	$i\theta^{\mu\nu}$	$i\delta_\nu^\mu$

As shown in Table 1, the non-commutative spacetime model has a non-zero commutation relation for the coordinate operators, which is a key difference from the commutative spacetime model.

Table 2. Properties of Different Algebraic Structures in Quantum Gravity.

Algebraic Structure	Main Features	Applications in Quantum Gravity
Non-Commutative Geometry	Non-commuting spacetime coordinates	Description of spacetime at small scales
Hopf Algebra	Incorporates symmetries and coalgebra structure	Description of particle interactions and spacetime symmetries
Lie Algebra	Describes infinitesimal symmetries	Study of local symmetries in spacetime

Table 2 summarizes the main features and applications of different algebraic structures in quantum gravity, which helps us understand the role of each algebraic structure in the construction of the spacetime algebraic structure [9].

Formula 1: Moyal Product for Two Scalar Fields

$$(\phi \star \psi)(x) = \int \frac{d^4k}{(2\pi)^4} \int \frac{d^4p}{(2\pi)^4} \tilde{\phi}(k) \tilde{\psi}(p) e^{i(k+p) \cdot x} e^{\frac{i}{2} \theta^{\mu\nu} k_\mu p_\nu} \tag{9}$$

where $\tilde{\phi}(k)$ and $\tilde{\psi}(p)$ are the Fourier transforms of the fields $(\phi)(x)$ and $(\psi)(x)$ respectively.

Table 3. Energy-Momentum Tensor in Non-Commutative Spacetime.

Component	Expression in Non-Commutative Spacetime
$T^{\mu\nu}$	$\frac{1}{\sqrt{-g}} \left(\frac{\delta S}{\delta g_{\mu\nu}} \right)$ with fields multiplied by Moyal product

Table 3 shows the expression of the energy-momentum tensor in non-commutative spacetime, which is modified due to the non-commutative nature of spacetime.

As shown in Tables 4, 5, and 6, the algebraic structure of curvature tensors, deformation parameters in quantum gravity theories, and comparisons of spacetime quantization approaches collectively illustrate the diverse mathematical frameworks used to describe quantum spacetime.

Table 4. Curvature Tensor Components in the Algebraic Spacetime Structure.

Curvature Tensor Component	Algebraic Expression
$R^\mu_{\nu\rho\sigma}$	Involves derivatives of the connection and commutation relations of coordinate operators

Table 5. Deformation Parameters in Quantum Gravity Theories.

Theory	Deformation Parameter	Physical Interpretation	Scale of Effect
Non-Commutative Geometry	$\theta^{\mu\nu}$	Spacetime uncertainty scale	$ \theta \sim l_p^2$ (Planck area)
Loop Quantum Gravity	γ (Immirzi parameter)	Quantization of area/volume	$l_p = \sqrt{\hbar G_N / c^3}$
String Theory	α' (String scale)	Effective spacetime resolution	$\sqrt{\alpha'} \sim l_s$ (string length)

Annotation: This table connects the abstract deformation parameters from different quantum gravity theories to their physical interpretations and the scales at which they become relevant. These parameters quantify how much the quantum theory deviates from classical general relativity and offer potential targets for experimental searches for Lorentz violation or other quantum gravity phenomena.

Formula 2: Deformed Heisenberg Formula of Motion

In non-commutative quantum mechanics, the time evolution of an observable A is governed by the deformed Heisenberg formula:

$$i\hbar \frac{dA}{dt} = [A, H] + \frac{i}{2} \theta^{\mu\nu} \{ \partial_\mu A, \partial_\nu H \} \tag{10}$$

where $\{\cdot, \cdot\}$ denotes the Poisson bracket. This formula incorporates both the classical commutator and non-local contributions from spacetime non-commutativity.

Table 6. Comparison of Spacetime Quantization Approaches.

Approach	Spacetime Structure	Algebraic Formulation	Key Symmetry
Canonical Quantum Gravity	Functional on superspace	Wheeler-DeWitt formula	Diffeomorphism invariance
Non-Commutative Geometry	Operator algebra $\mathcal{A} = C_\theta^\infty(\mathbb{R}^4)$	Moyal-Weyl algebra	Deformed Poincaré symmetry
Loop Quantum Gravity	Graphs with quantum labels	Asymptotic safety via Hopf algebras	Quantum diffeomorphism group

5. Physical Implications

5.1. Spacetime Geometry at the Quantum Level

The non-commutative nature of spacetime suggests that, at the quantum scale, it forms a discrete and fuzzy structure, unlike the smooth manifold in classical general relativity. This may have significant implications for studying black holes, the early universe, and the origins of spacetime ^[10].

5.2. Influence of Quantum Gravity on Particle Physics

The algebraic framework of spacetime can alter interaction terms in particle physics models. For instance, the Moyal product incorporates non-local interactions, potentially resulting in novel physical phenomena like altered scattering amplitudes and varied particle decay rates.

5.3. Cosmological Applications

In cosmology, the algebraic structure of spacetime is employed to study the early universe, encompassing primordial quantum fluctuations and the cosmic inflation process. Quantizing spacetime as a non-commutative algebra replaces the classical Big Bang singularity with a quantum bounce, keeping the universe’s energy density finite through the generalized uncertainty principle (GUP). In Hopf algebraic cosmological models, the comultiplication of momentum operators represents the distribution of primordial matter perturbations, resulting in a modified power spectrum that may be observable in cosmic microwave background (CMB) polarization data. Non-commutative effects contribute to resolving the horizon problem by altering the effective spacetime resolution scale, which modifies the speed of sound in the primordial plasma and enhances causality at quantum scales. Recent simulations with algebraic curvature corrections reveal that these effects can produce distinctive bispectrum signatures in CMB anisotropies, providing a testable connection between quantum spacetime geometry and observational cosmology.

5.4. Modifications to the Uncertainty Principle from Quantum Gravity

The algebra of non-commutative spacetime leads to a generalized uncertainty principle (GUP), altering the Heisenberg uncertainty relation. For position and momentum operators, the GUP adopts the form:

$$\Delta x^\mu \Delta p_\nu \geq \frac{\hbar}{2} |\delta_\nu^\mu + \beta \theta^{\mu\rho} \theta_{\rho\nu}| \tag{11}$$

where β represents a dimensionless constant approximately equal to one. This relation establishes a minimum measurable length scale $l_{min} \sim \sqrt{\theta}$, implying that spacetime cannot be probed below the Planck scale, potentially resolving the singularity issue in black hole physics. In classical general relativity, the Schwarzschild singularity at $r = 0$ transforms into a quantum-corrected region where the GUP stops infinite curvature, as indicated in Table 7.

Table 7. Singularity Resolution in Different Theories.

Theory	Singularity Behavior	Mechanism
Classical GR	Curvature diverges at $r = 0$	No quantum effects
Non-Commutative GR	Curvature peaks at $r \sim l_p$	GUP-induced minimum length
Loop Quantum Gravity	Discrete spacetime prevents $r \rightarrow 0$	Quantum geometry operators

5.5. Phenomenological Prospects

Precision measurements of photon propagation in astrophysical environments could serve as experimental tests for spacetime non-commutativity. For instance, the time-of-flight difference among high-energy gamma rays from distant gamma-ray bursts (GRBs) might disclose a momentum-dependent speed of light, indicating deformed dispersion relations in non-commutative spacetime. The modified dispersion relation is expressed as:

$$E^2 = p^2 c^2 + m^2 c^4 + \lambda \frac{E^4}{M_p^2} \quad (12)$$

where λ represents a dimensionless non-commutativity parameter and $M_p = \sqrt{\hbar c / G_N}$ denotes the Planck mass. Current observations from telescopes such as Fermi-LAT constrain $\lambda < 10^{-4}$, establishing limits on the scale of non-commutative effects.

6. Advanced Mathematical Framework: Category Theory Viewpoint (New Section)

To consolidate the diverse algebraic structures, we utilize category theory, wherein quantum spacetime is regarded as an object within a tensor category \mathcal{C} , equipped with suitable braiding and duality axioms. A central idea is the monoidal functor, which associates the tensor product of Hilbert spaces in quantum mechanics with the algebraic product in non-commutative geometry. The category of Hopf algebras \mathcal{H} constitutes a monoidal category under the tensor product, where the Poincaré Hopf algebra serves as a pivotal object that encodes spacetime symmetries.

The Drinfeld double construction offers a universal approach to quantize Lie algebras into Hopf algebras, playing a key role in deforming classical symmetries. For the Poincaré algebra $D(\mathfrak{poinc})$, the Drinfeld double $D(\mathfrak{poinc})$ integrates both left and right coactions, resulting in a non-commutative spacetime algebra that remains covariant under deformed Lorentz transformations. This categorical method offers a unified language for comparing various quantum gravity frameworks, as outlined in Table 8.

Table 8. Categorical Structure of Quantum Gravity Theories.

Theory	Category Type	Braiding Operator	Key Functor
Non-Commutative Geometry	Tensor category of algebras	Moyal braiding B_θ	GNS representation functor
Loop Quantum Gravity	Category of graphs	Holonomy-flux braiding	Functor to Hilbert space of spin networks
Quantum Group Approach	Monoidal category of Hopf algebras	R-matrix braiding	Universal enveloping algebra functor

7. Conclusion

The extended algebraic framework introduced here connects quantum mechanics and general relativity through a spacetime algebra that integrates non-commutative geometry with Hopf algebra symmetries. The detailed mathematical formulations, including representation theory, curvature formalism, and category theory unification, offer a solid foundation for exploring quantum spacetime phenomena. The generalized uncertainty principle and deformed dispersion relations from this framework provide testable predictions for future experiments, with the resolution of classical singularities underscoring its potential in early universe cosmology. Placing the theory in a categorical context allows for cross-fertilization with other quantum gravity approaches, promoting a more unified understanding of spacetime's quantum nature.

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