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# Pion Showers in the CALICE AHCAL 2 Prototype

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5 In this note we present results from the data collected with the CALICE Anal-  
6ogue Hadronic Calorimeter (AHCAL) during the 2007 test beam at the CERN  
7 SPS. The data analyzed correspond to events collected using negative pion beams  
8 in the energy range 8 - 80 GeV. Detailed investigations of the shower properties in  
9 pion beam data are described and compared with results obtained with Geant 4  
10 based Monte Carlo simulations. This is possible due to the high granularity of the  
11 scintillator-steel AHCAL prototype. The pion interaction length in the material  
12 mix of the AHCAL has been compared between simulations with several Geant 4  
13 physics lists and measured data at different beam energies. A comparison between  
14 data and simulation in the longitudinal shower shape and the energy density is  
15 presented and the differences found are discussed.

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## 37 1 Introduction

38 Measurements at the planned  $e^+e^-$  International Linear Collider (ILC) require a very high  
39 precision in energy - roughly a factor two better than what was achieved by detectors so far.  
40 One way to achieve the aimed jet energy resolution of  $\sigma_E/E \approx 3 - 4\%$  is the Particle Flow  
41 approach [1, 2, 3]. Particles in a jet are measured depending on their nature in the detector  
42 providing the best energy resolution. For charged particles the tracker is used, for photons  
43 the electromagnetic calorimeter and for neutral hadrons the combined electromagnetic and  
44 hadronic calorimeter. In order to be able to distinguish single particles in the detector system  
45 and to minimize confusion in the assignment of energy very highly granular calorimeters  
46 are necessary. The optimization of calorimeter designs heavily depends on Monte Carlo  
47 simulations which have to be validated against real data.

48 The CALICE collaboration developed several calorimeter prototypes using different tech-  
49 niques proposed for calorimeters at a future  $e^+e^-$  collider. In a test beam program in 2006  
50 and 2007 at the CERN SPS H6 beam line data was taken in a combined calorimeter setup.

51 In this note results from data taken with the CALICE analogue hadronic calorimeter proto-  
52 type (AHCAL) are presented. The prototype and the combined test beam setup are described  
53 in section 2. Section 3 gives an overview of the data set investigated and the cuts applied to  
54 select events of interest. An algorithm to find the starting point of the hadronic shower is  
55 discussed. The tools used for the comparison of data and Monte Carlo simulation including  
56 a brief discussion of the Geant 4 [4] physics lists studied and a method used for the decom-  
57 position of deposited energy are introduced in 4. In section 5 the distribution of the initial  
58 point of the shower and longitudinal profiles are compared between data and simulation.

## 59 2 Experimental setup

60 The data investigated in this note have been obtained during a three-month test beam cam-  
61 paign at CERN in 2007. In this section we briefly introduce the AHCAL prototype and the  
62 test beam setup.

### 63 2.1 The AHCAL Prototype

64 The CALICE AHCAL prototype is a  $1\text{ m}^3$  sandwich calorimeter with 2 cm thick steel absorber  
65 plates and 38 active layers using 0.5 cm thick plastic scintillator tiles with varying areas of  
66  $3 \times 3\text{ cm}^2$  in the core up to  $12 \times 12\text{ cm}^2$  in the outer part. The 7608 tiles are read out with  
67 Silicon Photomultipliers [5] via wave length shifting fibers. More details on the calorimeter  
68 and its commissioning can be found in [6].

### 69 2.2 Setup at the CERN SPS H6 Test Beam

70 The combined CALICE setup consisted of the Si-W electromagnetic calorimeter (ECAL), the  
71 scintillator-steel AHCAL and the scintillator-steel tail-catcher and muon-tracker (TCMT).  
72 The setup is sketched in Fig. 1. The trigger for the experiment is provided by a coincident  
73 signal from two  $10 \times 10\text{ cm}^2$  plastic scintillators (Sc1 and Sc3 in the figure). Multi-particle  
74 events can be rejected offline using data from a  $20 \times 20\text{ cm}^2$  multiplicity scintillator counter  
75 (Sc2) or from a  $1 \times 1\text{ m}^2$  beam halo veto (Veto in the figure) with a central opening  $20 \times 20$   
76  $\text{cm}^2$  along the beam line. Separation of electrons and pions is performed with a Čerenkov

77 counter upstream the setup, and muons are tagged by a  $1 \times 1 \text{ m}^2$  muon wall (Mc1) in the rear  
 78 of the detectors. A more detailed description of the setup is available in [7].

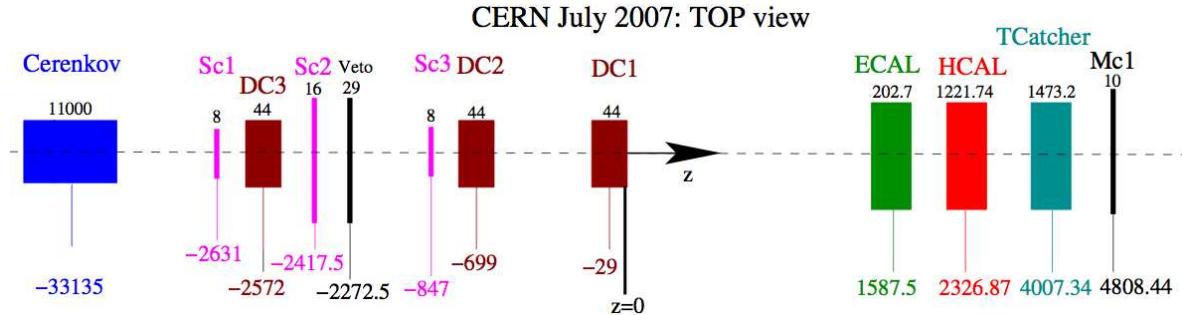


Figure 1: Setup during the 2007 data taking at CERN SPS H6 test beam site.

## 79 **3 Calibration and Event Selection**

### 80 **3.1 Calibration**

81 The equalization and calibration of calorimeter cell response including correction for temper-  
 82 ature effects is performed with the CALICE software v03-00 <sup>1</sup> and according to the procedure  
 83 described in [8]. The calibration chain can be summarized in the following steps:

- 84     • equalization of all cell responses
- 85     • correction of the non-linear SiPM response
- 86     • calibration with electromagnetic showers from test beam facilities

87 The factors for the equalization of cell response are obtained from muon runs acquired at  
 88 CERN. The same equalization factors and SiPM response correction curves are applied to  
 89 data and digitized Monte Carlo simulations (MC).

### 90 **Calibration Uncertainties**

91 The largest contribution to the calibration uncertainties is from the extraction of the MIP  
 92 values used for the cell equalization and is in the order of 3%. This is an uncertainty on  
 93 the energy scale and does not depend on the energy density. The value of 3% also includes  
 94 the uncertainty on corrected short term temperature variations. Uncertainties on the SiPM  
 95 non-linearity and the saturation point are instead energy density dependent. Their effect is  
 96 negligible at low energies and becomes relevant, i.e. comparable to the MIP uncertainty, only  
 97 for energies larger than 30 GeV.

98 Due to the complexity of the error propagation for these non-linear effects they are not  
 99 included in the present note. A more careful study on the effect of calibration uncertainties  
 100 is ongoing and the results will follow as an update to this note.

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<sup>1</sup>A custom installation equivalent to v03-00 has been used - cf. appendix C

101 **3.2 Event Selection**

102 In this study pion events are selected using the standard trigger coincidence between the two  
 103 trigger scintillators Sc1 and Sc2 (cf. section 2.2). Furthermore, pion showers are selected  
 104 which had a first hard interaction in the ten layers of the AHCAL after the first one (layers  
 105 number 2–11). The determination of the first hard interaction is performed using an algorithm  
 106 called Primary Track Finder (PTF). The details of the algorithm will be discussed in detail  
 107 in section 3.3. Events in which the shower starts in the first AHCAL layer are excluded to  
 108 minimize possible systematic effects introduced by this algorithm. The purpose of the limit  
 109 on the shower starting point to be in layers 2–11 is to minimize energy leakage from the  
 110 AHCAL.

111 The requirement on the PTF to find the first hard interaction in the AHCAL automatically  
 112 includes the rejection of muon events and leakage by electrons showering in the ECAL. Also  
 113 empty events caused by fake triggers and multi-particle events are rejected.

114 Table 1 shows the selection of runs chosen for this analysis. For all energies at least 100k  
 115 events have been recorded. At lower energies the amount of electrons in the beam is higher  
 116 than at high energies. This reduces the overall pion statistics in the low energy runs.

117 Fig. 2 shows the reconstructed energy in all three CALICE calorimeters before and after  
 118 applying the event selection requirements described above. After event selection, the energy  
 119 distribution in the ECAL is consistent with that of a minimum ionizing particle. A large  
 120 fraction of the pion energy is contained in the AHCAL, leading to the expected Gaussian  
 121 distribution for the bulk of the data. A small tail is still present in the AHCAL energy  
 122 distribution due to leakage to the TCMT. No further effort is made to select showers contained  
 123 in the AHCAL in order not to bias the longitudinal shower profiles to a particular type of  
 124 events. The bottom row shows the correlation between the energy deposited in the TCMT  
 125 and the sum of energy deposited in the ECAL and the AHCAL. A peak around 1-2 GeV  
 126 energy deposited in the ECAL and AHCAL is clearly visible before applying event selection  
 127 requirements. This peak is due to muons and pions traversing the AHCAL without interacting  
 128 and it is removed by the event selection requirements.

Run	Energy	Particle	Tot. ev.	Ev. after cuts	Efficiency
330334	8 GeV	$\pi^-$	105773	14494	13.7 %
330332	10 GeV	$\pi^-$	178504	26414	14.8 %
330330	12 GeV	$\pi^-$	261601	45113	17.2 %
330328	15 GeV	$\pi^-$	179131	36677	20.5 %
330327	18 GeV	$\pi^-$	178369	37662	21.1 %
330326	20 GeV	$\pi^-$	180279	38764	21.5 %
330325	25 GeV	$\pi^-$	177620	38492	21.7 %
330960	35 GeV	$\pi^-$	182907	34768	19.0 %
330961	45 GeV	$\pi^-$	174589	39074	22.4 %
330962	80 GeV	$\pi^-$	179777	40325	22.4 %

Table 1: Set of runs investigated in this note. The shower is required to start in the front part of the AHCAL. The number of events after this cut is given in column five and as a relative value in the last column.

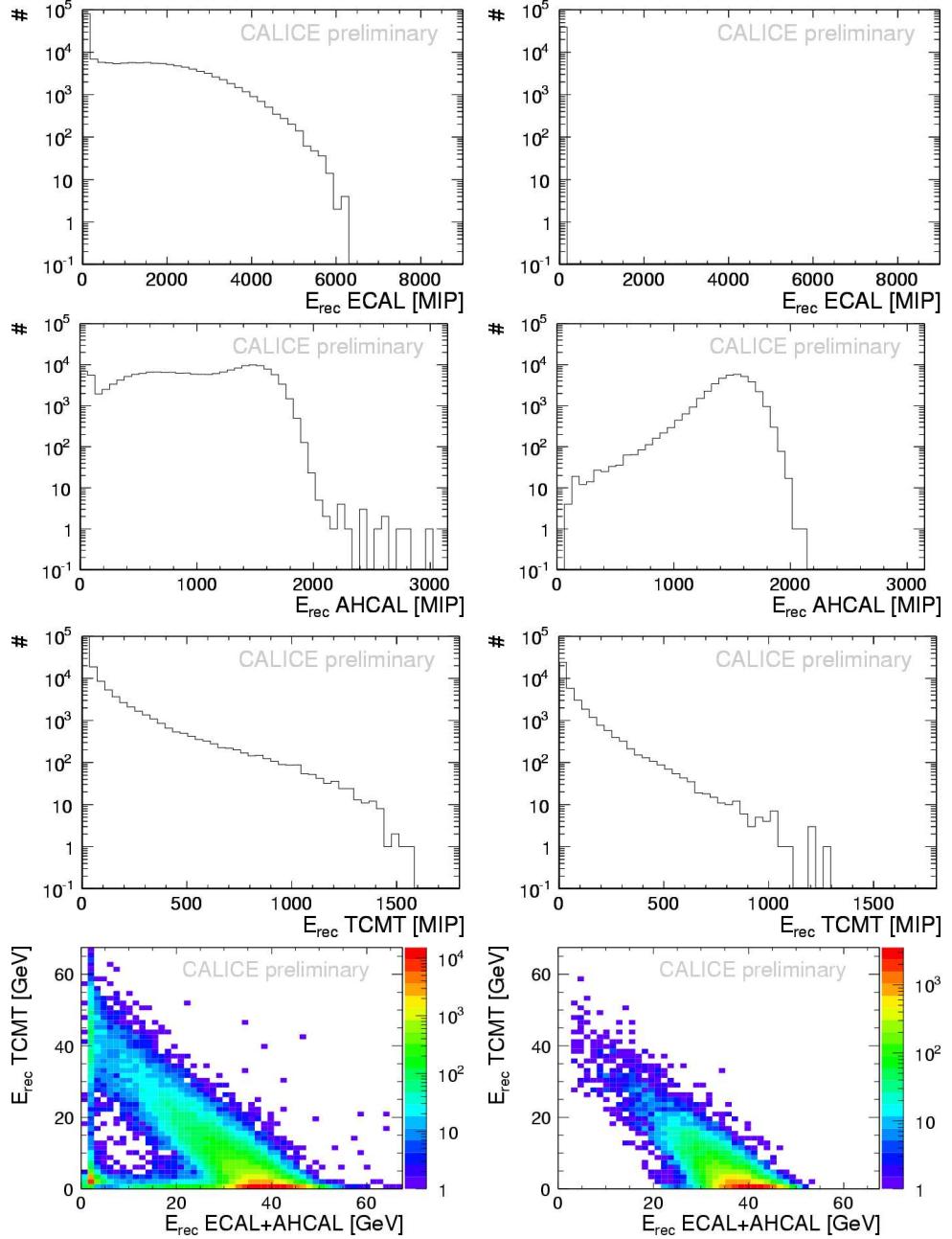


Figure 2: Distributions of reconstructed energy from 45 GeV pion showers in all three CALICE calorimeters before (left) and after (right) applying the event selection. In the bottom plots the correlation between the energy deposited in the TCMT and the sum of the energy deposited in the ECAL and AHCAL in GeV is shown.

129 **3.3 The Primary Track Finder Algorithm**

130 The Primary Track Finder algorithm (PTF) has been developed by M. Chadeeva [9]. Its pur-  
131 pose is to find the track corresponding to a primary particle entering the combined CALICE  
132 detector setup up to the starting point of its shower.

133 The starting layer is found using the accumulated average energy deposited,

$$A_i = \sum_{k=0}^i E_k / (i + 1). \quad (1)$$

134  $E_k$  being the energy deposited in layer  $k$  of the AHCAL.

135 The layer  $i$  where an hadronic shower starts is then defined as the first layer for which any  
136 of the following two criteria is fulfilled:

137 -  $(A_i + A_{i+1}) > 6.5 \text{ MIP}$  and  $N_i^{hit} + N_{i+1}^{hit} > 8$   
138 -  $E_{i+1} > E_{thr}$

139 where  $N_i^{hit}$  is the number of hits in layer  $i$ . The energy threshold  $E_{thr}$  expressed in units of  
140 MIP is heavily dependent on the beam energy. In the latest official version of the PTF this  
141 parameter is fixed for all events in one run, with a default value of 25 MIP. For this analysis  
142 a modified version of the PTF is used, with a threshold varying as a function of the energy  
143 deposited in the calorimeter system (sum of ECAL, AHCAL and TCMT). The information  
144 of the beam energy is not used by the algorithm. The energy dependence of the threshold  
145 value is parametrized by a first order polynomial. It is 14.5 MIP at 8 GeV and 32 MIP at 80  
146 GeV energy deposited in the calorimeters.

147 **Systematic uncertainty of the PTF algorithm**

148 In order to evaluate systematic effects due to the PTF algorithm, the algorithm-determined  
149 layer of the first hard interaction is compared with the true point of first hadronic interaction  
150 in the Mokka [10] simulation. If this point is either inside an active AHCAL layer or in the  
151 absorber plate just before it, that layer is assumed to be the true starting point.

152 Fig. 3 Shows the performance of the PTF for an 80 GeV  $\pi^-$  run simulated using the  
153 QGSP\_BERT physics list.

154 The performance of the PTF is dependent on the beam energy and the physics list used.  
155 For all physics lists and all energies for 74% (84%) of the events there is an agreement of  
156 better than  $\pm 1$  ( $\pm 2$ ) layers between the first interaction layer found by PTF and the MC  
157 truth. Averaged over all energies and all physics lists, there is a systematic shift of -0.2 layers  
158 and a correlation of 85.8%. Further tuning might improve the quality of the PTF.

159 **4 Simulation**

160 The simulation of the entire setup is performed using the program Mokka version 7.02 [10].  
161 Mokka is a Geant 4 based application able to simulate full ILC detector geometries as well  
162 as the 2007 CERN test beam setup. The geometry and materials used in the simulation are  
163 given in [7]. The simulated events from Mokka were passed through the CALICE digitization  
164 chain v03.01. The digitization includes emulation of light leakage across calorimeter cells and  
165 saturation effects in the AHCAL.

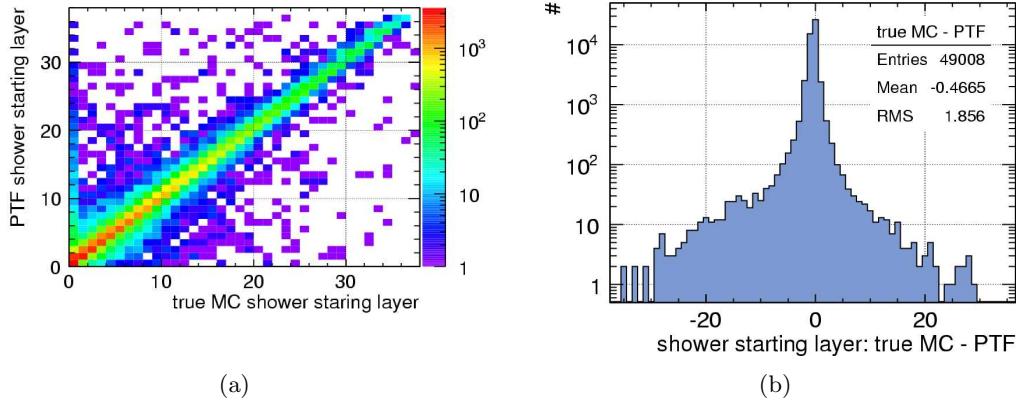


Figure 3: Performance of the Primary Track Finder (PTF) compared to MC true using the QGSP\_BERT physics list simulating an 80 GeV  $\pi^-$  beam. The correlation between the PTF and the true MC shower starting layer (a) is 97%. The average difference between the shower starting layer found by the PTF and the true one (b) is 0.5 AHCAL layers.

#### 166 4.1 AHCAL Digitization

167 In Mokka the AHCAL detector is simulated with  $1 \times 1 \text{ cm}^2$  scintillator tiles. In the digitization  
 168 these cells are then ganged to  $3 \times 3$ ,  $6 \times 6$  and  $12 \times 12 \text{ cm}^2$  cells according to the true geometry of  
 169 the AHCAL prototype. As next step in the digitization, detector effects such as light leakage  
 170 between tiles and non-linearity effects of the photo-detectors are simulated. For the current  
 171 studies a light leakage factor of 10% between calorimeter cells is assumed. Non-linearity  
 172 effects are simulated (and corrected during the reconstruction step) according to individual  
 173 measurements for each calorimeter cell. More details on the digitization and the simulated  
 174 detector effects can be found in [11].

#### 175 4.2 Physics Lists

176 For the simulation of hadronic showers several models exists which partially differ in their  
 177 predictions. The Monte Carlo simulation framework Geant 4 implements different models  
 178 [12] valid in certain energy ranges and combines them in so called physics lists:

179 **QGS** The Quark-Gluon-String model is used to simulate the interactions of protons, neutrons,  
 180 pions and kaons with nuclei in the energy range from 10 GeV to 50 TeV. It needs other  
 181 models (e.g. Pre-Compound) to de-excite and fragment the residual nuclei.

182 **FTF** In the Fritiof model string formation is performed via scattering of projectiles on nucle-  
 183 ons. It is valid for energies  $> 4 \text{ GeV}$ .

184 **LEP & HEP** are the low energy parameterized (LEP) and high energy parameterized (HEP)  
 185 models which have their origins in the GHEISHA hadronic package. They depend on  
 186 parameterized fits to measured data and are no detailed theory driven hadronic models.

Physics List	Model content (for $\pi^\pm$ )
LHEP	LEP (0–55 GeV); HEP ( $>25$ GeV)
QGSP_BERT	Bertini (0–9.9 GeV); LEP (9.5–25 GeV); QGSP ( $>12$ GeV)
FTF_BIC	BIC (0–5 GeV); FTFB ( $>4$ GeV)
CHIPS	applied over full energy range

Table 2: The table shows the physics models invoked for pion inelastic interactions in each physics list. Where ranges overlap, **GEANT4** chooses randomly between models, with probabilities varying linearly with energy over the range of overlap.

187 Energy and momentum are only conserved on average, but not event by event. These  
 188 models are outdated but are still commonly used in many Geant 4 physics lists for  
 189 particles or energy ranges for which no better model exists.

190 **Binary cascade** This model is used for low energies ( $< 10$  GeV). Nucleon-nucleon scattering  
 191 is done by resonance formation and decay.

192 **Bertini cascade** This model includes intra-nuclear cascades followed by pre-compound and  
 193 evaporation phases for the residual nucleus and is valid for low energies ( $< 10$  GeV).

194 **Pre-Compound model** Theory driven model. It is used as back-end for other models to  
 195 model pre-compound and evaporation phases of residual nuclei.

196 In the Geant 4 physics lists there are overlaps in the energy ranges the several models are  
 197 applied. In these transitions from one model to another model Geant 4 chooses randomly  
 198 between the two models on an event-by-event basis. The following physics lists [13] were  
 199 considered:

200 **LHEP** is a combination of LEP and HEP.

201 **QGSP\_BERT** is a combination of the QGS model with the Pre-Compound model above 12  
 202 GeV, the Bertini cascade below 9.9 GeV and LEP parameterization between 9.5 GeV  
 203 and 25 GeV.

204 **FTF\_BIC** uses the Binary cascade for protons, neutrons, pions and kaons for energies below  
 205 5 GeV. For energies above 4 GeV the Fritiof model is used. The BIC model is again  
 206 used for the re-scattering of secondaries in this case.

207 **CHIPS** The experimental Chiral Invariant Phase Space (CHIPS) physics list which simulates  
 208 all elastic and inelastic hadron-nuclear reactions, photo/lepto-nuclear reactions, the  
 209 stopping of hadrons and synchrotron radiation. For the simulation of events with this  
 210 physics list the patched Geant 4.9.3.p01 version was used.

211 The physics content of these physics lists for pions is summarized in Table 2.

### 212 4.3 Energy Decomposition

213 Mokka allows to determine the energy deposited by a given particle type in a certain simulated  
 214 material volume. However, this information is only available in the first stage of the simulation  
 215 chain, still at the  $1 \times 1 \text{ cm}^2$  tile level, and it is lost in the first step of the digitization, i.e. the

216 ganging. Due to the non-linear detector effects, threshold effects and non-Gaussian calibration  
 217 smearing, data and simulation comparison can only be performed at the digitized MIP level.  
 218 In the following we describe the method developed to retrieve single particle energy deposition  
 219 in the scintillator (visible energy) event-by-event through the digitization chain - see also  
 220 Fig. 4.

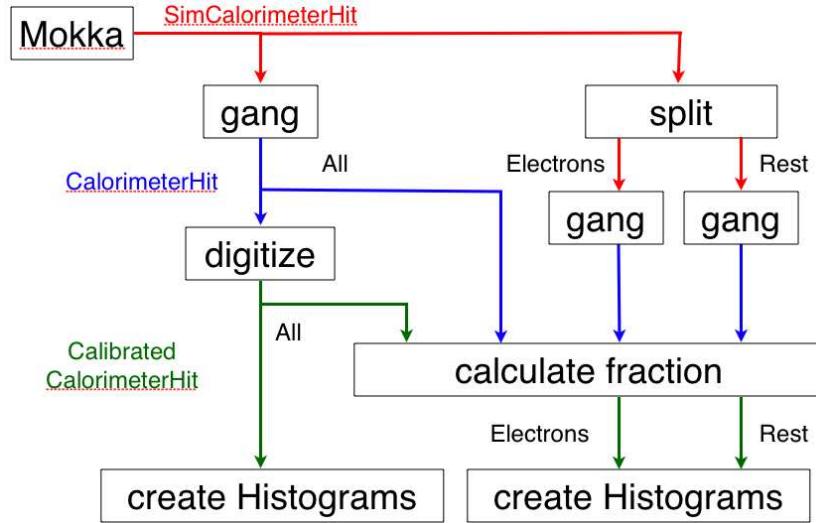


Figure 4: Flow chart of the developed technique to access the energy deposited by single particles in Mokka. For simplicity the process is only shown for one particle type (Electrons) and all other particles (Rest).

221 • A collection of visible energy is created for each particle type. The collections are filled  
 222 event-by-event at the  $1 \times 1 \text{ cm}^2$  tile level.

223 • The  $1 \times 1 \text{ cm}^2$  hits divided in particle-type specific collections are then ganged, and new  
 224 collections are created containing information at the real size tile level - the same way  
 225 as for the hits not decomposed by particle type.

226 • For the merged hits the ratio between energy deposited by a certain particle type and  
 227 the total energy are calculated. These ratios are used as weights to obtain, after the  
 228 digitization steps, the fractions of particles of various type.

229 • The standard AHCAL digitization is applied to all hits, including a threshold cut of 0.5  
 230 MIP to reject noise.

231 • After digitization the total energy at the MIP scale is multiplied by the weights previ-  
 232 ously obtained to get the digitized fractional energy of each particle type.

233 The result of the visible energy decomposition is shown in Fig. 5 for 80 GeV pions. A  
 234 difference of  $\approx 7\%$  between the peak of the energy sum from data and the one from the  
 235 FTF\_BIC prediction (filled histogram) is visible. The observed difference is consistent with  
 236 another independent investigation [14]. The other histograms in Fig. 5 are the distributions

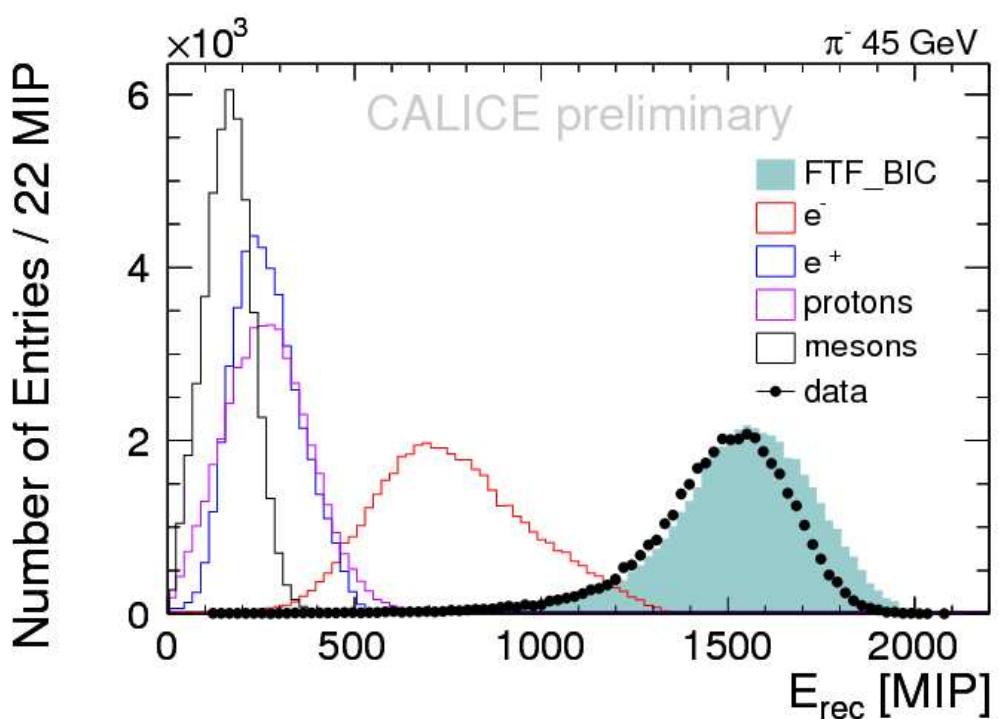


Figure 5: The breakdown of the reconstructed visible energy into the contributions from various particles:  $e^\pm$ , p, and mesons ( $\pi^\pm$  and  $K^\pm$ ).

237 of visible energy deposited by individual particle types, i.e.  $e^\pm$ , p, and mesons ( $\pi^\pm$  and  $K^\pm$ ).  
238 The information can be used to investigate disagreements between data and simulation in  
239 same observables to a deeper level, as it will be shown in the following.

## 240 5 Results

241 In this section we present the results of the comparison between data and simulations for  
242 various observables. We try to outline which physics list is closest to the data in the energy  
243 range covered by these studies. All results apply the calibration procedure and the event  
244 selection criteria described in section 3.

### 245 5.1 Total visible energy

246 The ratio between the reconstructed energy for the simulated and for the negative pion  
247 showers is shown in figure 6 at beam energies of 8, 18 and 80 GeV. The CHIPS physics list  
248 shows an energy independent overestimation of roughly 8%, while the response of the other  
249 physics lists varies. This overestimation is expected, since the low energy neutron cross-  
250 sections are not yet properly implemented in CHIPS [15]. LHEP predicts 9% too low energy  
251 deposition at 8 and 18 GeV (LEP parameterization) and 4% too low energy deposition at 80  
252 GeV (HEP parameterization). The QGSP\_BERT physics lists agrees with the data within  
253 the uncertainty at 8 and 18 GeV (Bertini Cascade and LEP), while at 80 GeV the QGS  
254 high energy model overestimates the energy by 6%. The FTF\_BIC physics list behaves very  
255 similar: it underestimates the energy by 4% at 8 GeV (Binary cascade and FTF), agrees  
256 within the uncertainty at 18 GeV and gives 4% to high energy deposition at 80 GeV (at both  
257 higher energies FTF dominates the description of the first hard interaction).

### 258 5.2 First hard interaction

259 Exploiting the high granularity of the AHCAL, the position of the first hard interaction of  
260 hadrons in the calorimeter is determined. This information is used to deconvolute the true  
261 longitudinal development of the hadronic showers from the starting point. In figure 7(a) the  
262 shower profile from the calorimeter front face (filled histogram) and the one relative to the  
263 first interaction point (black line) are shown.

264 If the fluctuations of the shower starting position are removed the shower looks shorter  
265 and more similar to an electromagnetic one. As it will be discussed later in section 5.3,  
266 electromagnetic processes indeed dominate the shower shape.

267 From the distribution of the shower starting position one can directly extract the effective  
268 nuclear interaction length of pions in the material mix of the AHCAL. An example of the  
269 distribution and the applied exponential fit is shown in figure 7(b). This is a consistency check  
270 of the validity of the PTF algorithm applied on both data and MC. The effects introduced by  
271 the uncertainty of the PTF algorithm are largest in the first two and the last eight calorimeter  
272 layers. These layers are excluded from the fit. Figure 8 shows the values of the pion interaction  
273 length  $\lambda_{int}^\pi$  in cm extracted from the data and from the various physics lists.

274 The statistical error on  $\lambda_{int}^\pi$  extracted from the fits is below 1%. The uncertainty of the  
275 algorithm distorts the exponential form of the shower starting point distribution and trans-  
276 lates into a systematic uncertainty on the interaction length extracted from the fits. This  
277 uncertainty is dependent on the physics list used and the beam energy. To estimate this

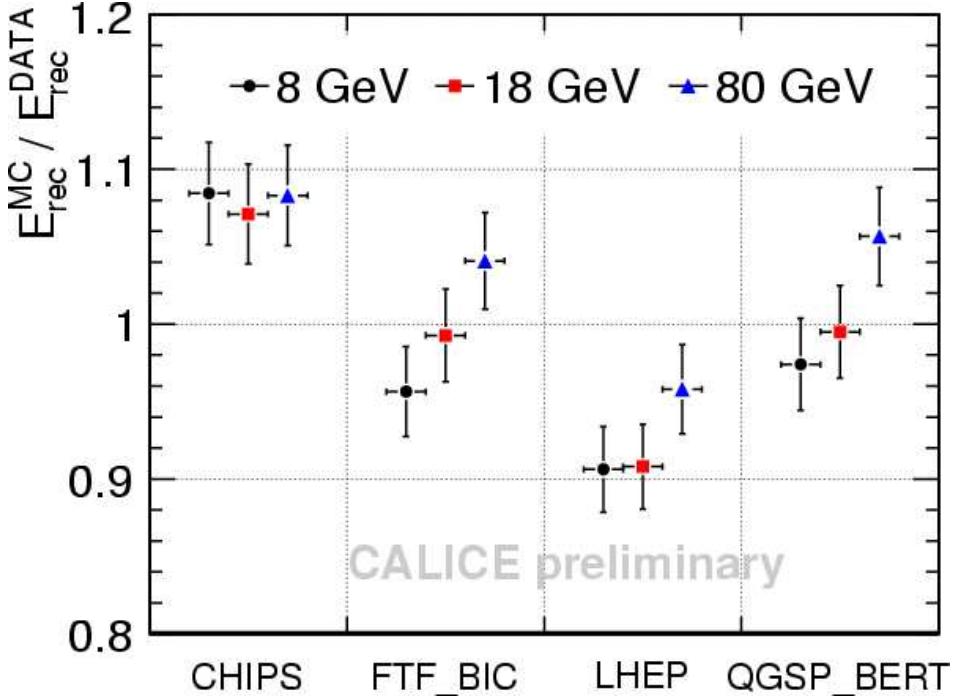


Figure 6: Ratio between the reconstructed energy in the AHCAL from simulation and from data for negative pion showers as a function of beam energy. All physics lists except CHIPS show an energy dependent behavior.

uncertainty two fits are performed, one to the true MC distribution and one to the distribution obtained determining the first interaction layer with the PTF algorithm. The RMS of the difference between the two fit values for each beam energy is used as uncertainty for the interaction length. The uncertainty varies from roughly 5% at 8 GeV to less than 1% at 80 GeV. In addition, a systematic increase of 4-5 cm is observed for all physics lists in the  $\lambda_{\text{int}}^{\pi}$  obtained using the algorithm with respect to the true MC one (c.f. Appendix B). The same systematic errors are assumed for the data.

Within uncertainties the fit yields the same effective nuclear interaction length,  $\lambda_{\text{int}} \sim 30$  cm for all the physics lists which use the same pion cross section (all apart from LHEP which has a larger cross-section and a  $\lambda_{\text{int}} \sim 26$  cm, and CHIPS which has a smaller cross-section and a  $\lambda_{\text{int}} \sim 31$  cm). Data are found to be consistent with the majority of the models yielding a value of  $\lambda_{\text{int}} \sim 29 \pm 1$  cm. The effect of the transition from QGS to LEP in the QGSP\_BERT physics lists is visible below 25 GeV. Above 25 GeV, where the QGS model is applied, the QGSP\_BERT physics lists agrees with the data.

### 5.3 Longitudinal Shower Profiles

Thanks to the very high granularity of the CALICE calorimeters, the longitudinal profile of hadronic showers can be investigated with an unprecedented accuracy. In particular, the true longitudinal development can be de-convolved from the distribution of the shower starting points. When measured from the position of the first hadronic interaction, as opposed to from the calorimeter front layer, hadronic showers are shorter and any layer-to-layer fluctuations

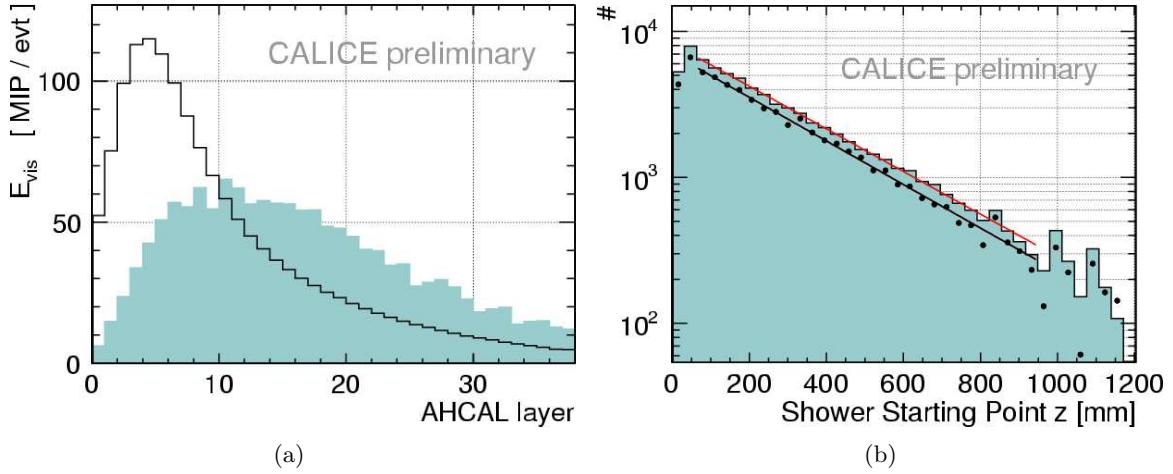


Figure 7: (a) Longitudinal profile for a 45 GeV  $\pi^-$  run relative to the calorimeter front (filled histogram) and relative to the first interaction (black line). (b) Distribution of shower start detected with the PTF in the GEANT4 simulation for the QGSP\_BERT physics list (filled histogram) and 45 GeV pions data (filled circles).

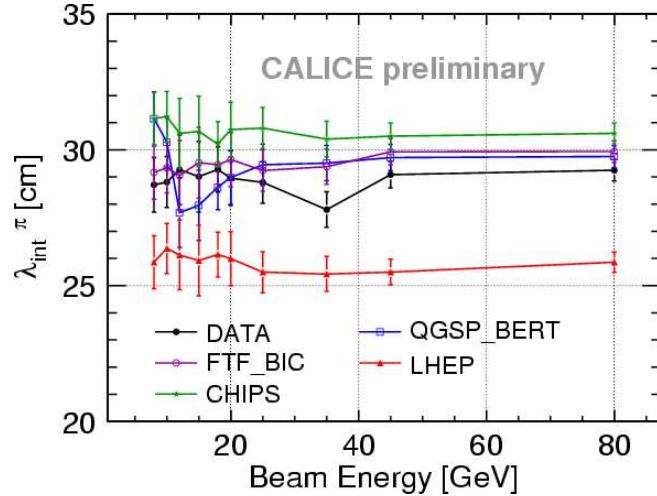


Figure 8: Extracted values for the interaction length in cm from data and MC models. The error bars shown are the uncertainty from the fitting added in quadrature with the uncertainty of the algorithm to determine the first interaction layer (see text). The 35 GeV run is currently under investigation.

298 introduced by calibration and dead channel effects are washed away.

299 The breakdown of the energy contribution from various particles in the shower ( $e^\pm, p, \pi^\pm, \mu^\pm$ ,  
300 can be determined for the simulated events. This additional information helps to un-  
301 derstand which physics processes contribute most in which phase of the shower development.  
302 This is clearly visible comparing the electron and positron contribution to the entire pro-  
303 file. The hadrons only contribute significantly in the very first layers after the first hard

304 interaction.

305 In this analysis pion events collected at three different energies have been chosen to compare  
306 the longitudinal shower profile measured in data to that obtained from MC. The various MC  
307 models playing a role at low, medium and high energies are investigated.

308 Fig. 9 shows the longitudinal shower profiles for pions of 8, 18 and 80 GeV. The black  
309 points are data which are compared to the various MC models shown as filled histograms.  
310 The error bars show the effect due to the uncertainty in the determination of the first hard  
311 interaction only, and were estimated from comparison to profiles relative to the true first  
312 interaction layer in MC. The dominant systematic uncertainty from calibration is a 3% error  
313 on the energy scale. This would result in a coherent scaling of the whole spectrum and is  
314 therefore not shown here.

315 The position of the shower maximum is in general quite well simulated. All models except  
316 CHIPS tend to underestimate the tails of the showers seen in data.

317 At 8 GeV LHEP overestimates the deposited energy in the first layer. After the fifth layer  
318 it predicts to little energy. The QGSP\_BERT physics lists agrees with the data within the  
319 uncertainty. The FTF\_BIC simulation is similar to the one using QGSP\_BERT. The predicted  
320 showera are slightly to short than in data. In layers 1-3 CHIPS agrees with the data, but the  
321 showers are too long.

322 All physics are close to data until approximately one  $\lambda_{int}^\pi$  after the first hard interaction  
323 at 18 GeV beam energy. Bigger differences become visible after this point for CHIPS (up to  
324 30% overestimation) and LHEP (down to 30% too little energy deposition). The data are well  
325 described by QGSP\_BERT, besides slightly too short showers. The performance of FTF\_BIC  
326 is a little worse, but comparable to QGSP\_BERT. Too much energy is deposited by CHIPS  
327 in the shower maximum and the showers are too long.

328 At 80 GeV LHEP simulates the tail of the shower profile quite well, but the energy de-  
329 position in the shower maximum is too low by almost 20%. The other three physics lists  
330 overestimate the energy in the shower maximum by 20 – 25%. The highest energy deposition  
331 is predicted by QGSP\_BERT, followed by FTF\_BIC. Both simulate too short showers. The  
332 longitudinal shower shape predicted by CHIPS at this energy is closer to data. Still theres  
333 is too much energy in the shower maximum and the showers are too long. Since the energy  
334 deposition in the shower maximum is dominated by electrons and positrons, the overshoot in  
335 the energy there predicted by QGSP\_BERT, FTF\_BIC and CHIPS could be a hint to a too  
336 high electromagnetic fraction at 80 GeV.

## 337 **6 Summary and Conclusions**

338 In this note several variables related to the shower physics have been investigated in order  
339 to compare Geant 4 models and to validate them against data from real pion showers in the  
340 AHCAL detector in the range from 8 - 80 GeV. Events have been simulated using several  
341 Geant 4.9.3 physics lists as well as the CHIPS physics lists from the patched version Geant  
342 4.9.3.p01.

343 The calorimeter response predicted by four physics lists has been investigated and compared  
344 to data. The QGSP\_BERT, FTF\_BIC and LHEP physics lists show an energy dependent  
345 deviation from data. The CHIPS physics lists estimates roughly 8% to high response, but  
346 looks promising since it does not show a dependency on the beam energy. The overestimation  
347 of the response is expected by the developers due to the low energy neutron cross sections

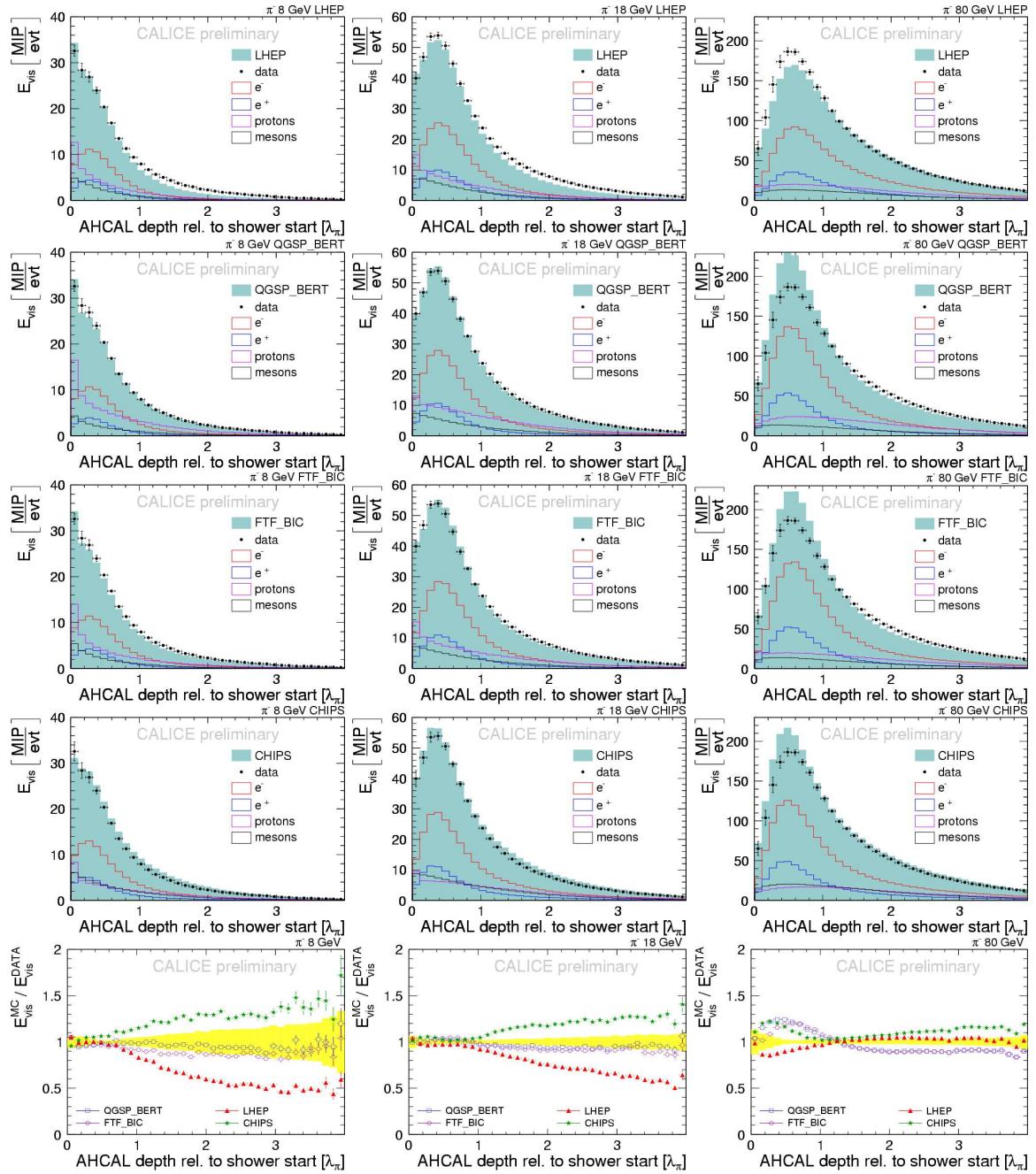


Figure 9: Longitudinal shower profile for 8, 18 and 80 GeV pions. The error bars include the statistical uncertainty and the uncertainty introduced by the determination of the first hard interaction. The breakdown of the energy contribution from various particles in the shower ( $e^\pm, p, \pi^\pm, \mu^\pm$ ) is shown. The bottom row of plots shows the ratio of simulation and data.

348 not yet implemented properly.

349 The nuclear interaction length  $\lambda_{int}^\pi$  for pions in the AHCAL have been extracted from the  
 350 distribution of the first hard interaction position found with the Primary Track Finder algo-

351 rithm. The values for  $\lambda_{int}^\pi$  extracted from the Monte Carlo simulations have been compared  
352 to the ones for the data at different beam energies. This is a test of the pion cross sections  
353 implemented in the physics lists. Good agreement between the FTF\_BIC physics list and  
354 data is found. LHEP simulates a higher cross section than the other physics lists which re-  
355 sults in an too low interaction length compared to data. The application of the LEP model  
356 in the QGSP\_BERT physics list at energies below 20 GeV is clearly visible. Above 20 GeV  
357 QGSP\_BERT describes the DATA as well as FTF\_BIC. For the CHIPS physics list a too high  
358 interaction length compared to data is found. This was expected by the developers due to  
359 the quasi-elastic and diffractive cross-sections used.

360 The last variable studied are highly granular longitudinal shower profiles relative to the  
361 position of the first hadronic interaction. The Mokka simulation allows to access the contri-  
362 bution of single particle species to the calorimeter hits which might give a hint for the reason  
363 of the disagreement between simulated and real showers. Profiles at three beam energies  
364 where different MC models in the physics lists come into play are presented. Showers in  
365 data are systematically longer than the predictions by QGSP\_BERT, FTF\_BIC and LHEP.  
366 Too long showers are found for the CHIPS and have been predicted by the developers. At  
367 high energies all physics list but LHEP simulate a too high energy deposition in the shower  
368 maximum.

369 The Primary Track Finder used for the studies in this note is still in development and results  
370 might further improve. The correct treatment of calibration uncertainties is still missing and  
371 will follow as an update to this note.

## 372 A Additional Simulations

373 More simulations with additional physics lists (all Geant 4.9.3) have been done and are pre-  
 374 sented in this appendix. Longitudinal profiles for 8, 18 and 25 GeV for the following physics  
 375 lists are shown in fig. 10:

376 **QGSP\_FTFP\_BERT** This is a variant of the QGSP\_BERT physics list where the LEP pa-  
 377 rameterization is replaced by the FTF model in the interval ( $6 \text{ GeV} < E < 25 \text{ GeV}$ ).  
 378 The Bertini models overlaps with FTF in the range ( $6 \text{ GeV} < E < 8 \text{ GeV}$ ), while the  
 379 QGS model overlaps with FTF in the range ( $12 \text{ GeV} < E < 25 \text{ GeV}$ ).

380 **QGS(P)\_BIC** Combination of the QGS model at high energy with the Binary cascade model  
 381 at low energy and the usual bridge between the two covered by LHEP. In the version  
 382 with the Pre-Compound model turned on (QGSP\_BIC) the BIC model is not used for  
 383 pions (only for neutrons and protons).

384 **FTFP\_BERT\_TRV** Uses the Pre-compound model and the Bertini cascade for the low energy  
 385 region. For the high energy region, the Fritiof model is used. FTFP\_BERT\_TRV is a  
 386 variant of FTFP\_BERT with the transition region between the BERT and the FTF  
 387 model shifted from 4-5 GeV (default) to 6-8 GeV.

388 The physics content of these models for pions is summarized in Table 3.

Physics List	Model content (for $\pi^\pm$ )
QGSP_FTFP_BERT	Bertini (0–8 GeV); FTFP (6–25 GeV); QGSP (>12 GeV)
QGS_BIC	BIC (0–1.3 GeV); LEP (1.2–25 GeV); QGSB (>12 GeV)
QGSP_BIC	LEP (<25 GeV); QGSP (>12 GeV)
FTFP_BERT_TRV	Bertini (0–8 GeV); FTFP (>6 GeV)

Table 3: The table shows the physics models invoked for pion inelastic interactions in each physics list. Where ranges overlap, GEANT4 chooses randomly between models, with probabilities varying linearly with energy over the range of overlap.

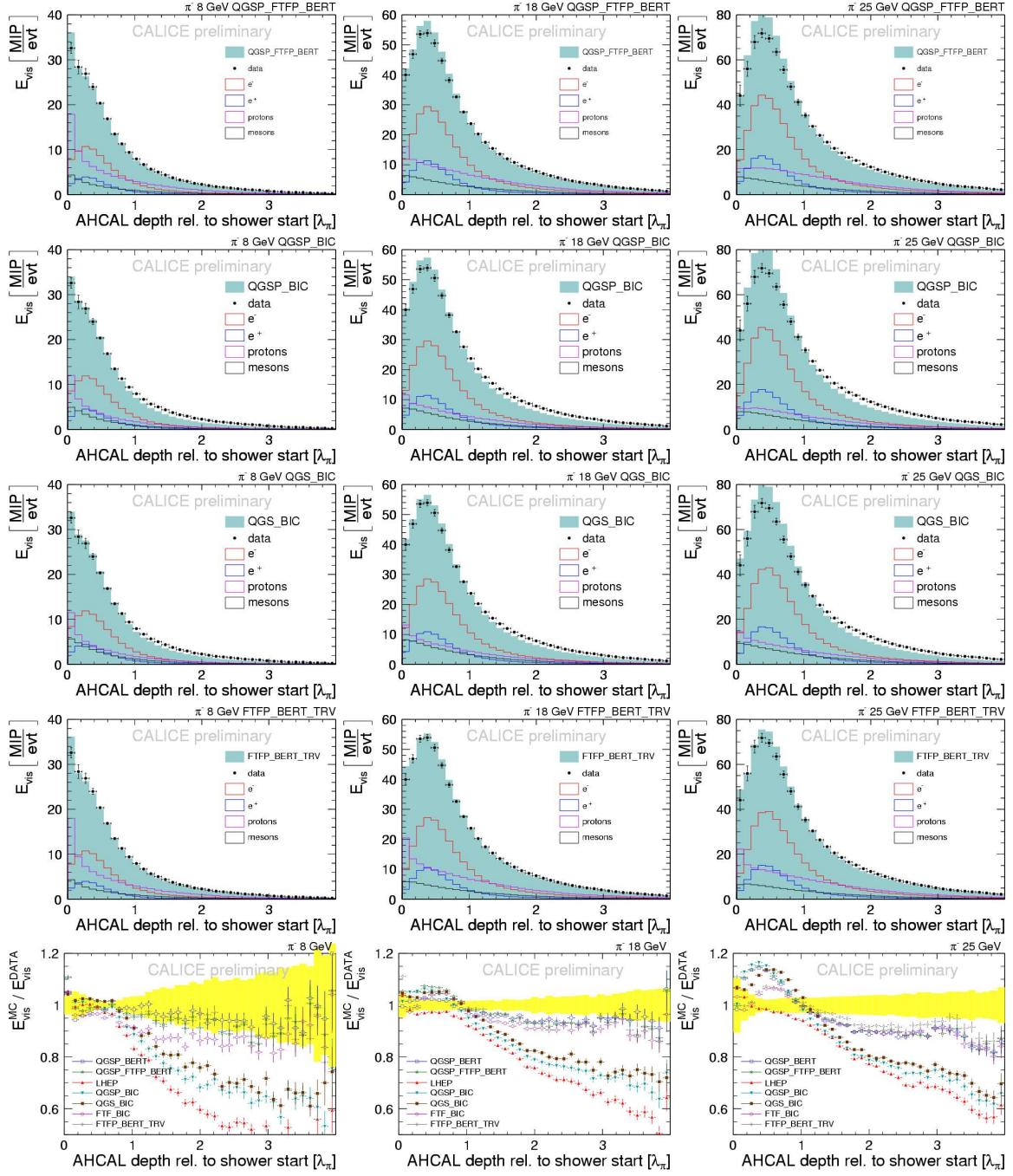


Figure 10: Longitudinal shower profile for 8, 18 and 25 GeV pions. The error bars include the statistical uncertainty and the uncertainty introduced by the PTF.

## 389 B Nuclear Interaction Length for Pions from true MC

390 To estimate the uncertainty on  $\lambda_{int}^\pi$  found with the PTF, also the true interaction length  
 391  $\lambda_{int}^{\pi, true}$  using the real first interaction layer available in MC has been obtained. It has been  
 392 extracted using the same procedure as described in section 5. The extracted values are shown  
 393 in Fig. 11 for all physics lists (a), a zoomed picture is given in (b). All physics lists (excluding  
 394 LHEP) use the same cross section for the pion interactions and thus the same interaction  
 395 length within the uncertainty from the fit is found. Since no distortion of the exponential  
 396 distribution for the shower starting layer due to the PTF is present, the absolute values of  
 397  $\lambda_{int}^{\pi, true}$  are systematically smaller by 4-5 cm than the ones found by the PTF. The extracted  
 398 values of the pion interaction length found by the PTF are shown in fig. 12.

399 The difference between the interaction length found using the true first interaction length  
 400 and the one found using the PTF has been average over all physics lists. The RMS of this  
 401 value is a measure for the uncertainty  $\lambda_{int}^\pi$  found with the PTF. In table 4 the mean shift and  
 402 uncertainty for each energy point is summarized.

Beam Energy [GeV]	$\langle \lambda_{int}^\pi \rangle$ [mm]	$\langle \lambda_{int}^\pi - \lambda_{int}^{\pi, true} \rangle$ [mm]	RMS [mm]	$\Delta(\lambda_{int}^\pi)$
8	288	49	15	5 %
10	286	44	11	4 %
12	282	39	7	3 %
15	286	39	7	2 %
18	289	40	4	1 %
20	291	41	3	1 %
25	294	43	2	1 %
35	295	40	1	1 %
45	297	40	1	1 %
80	298	39	2	1 %

Table 4: Uncertainty introduced by the PTF on extracted interaction length averaged over all physics lists. The mean interaction length found by the PTF (1st column), the mean difference between the interaction length found using PTF and true MC (2nd column), the RMS of this value (3rd column), and the relative uncertainty on  $\langle \lambda_{int}^\pi \rangle$  in percent (last column).

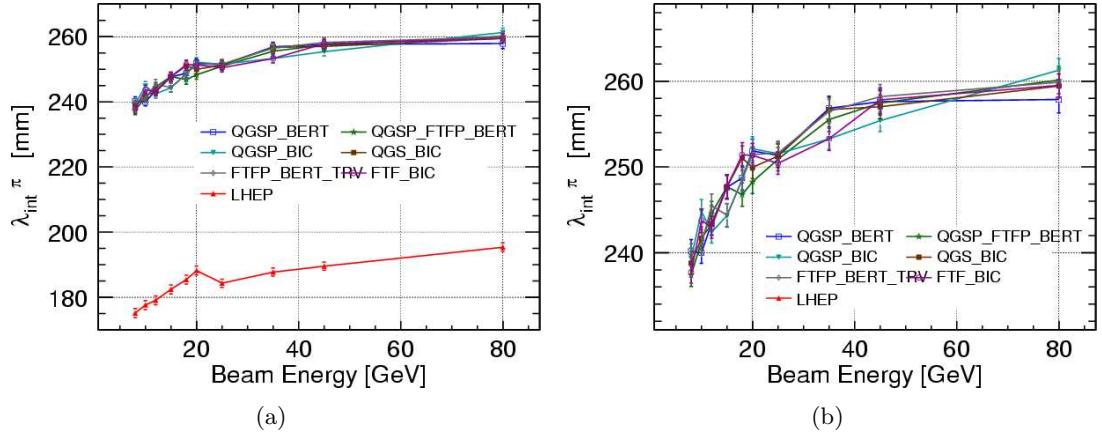


Figure 11: Pion Interaction length extracted using the true shower starting layer in Monte Carlo simulation for all physics lists (a) and zoomed in (b). Error bars shown are the fit uncertainty.

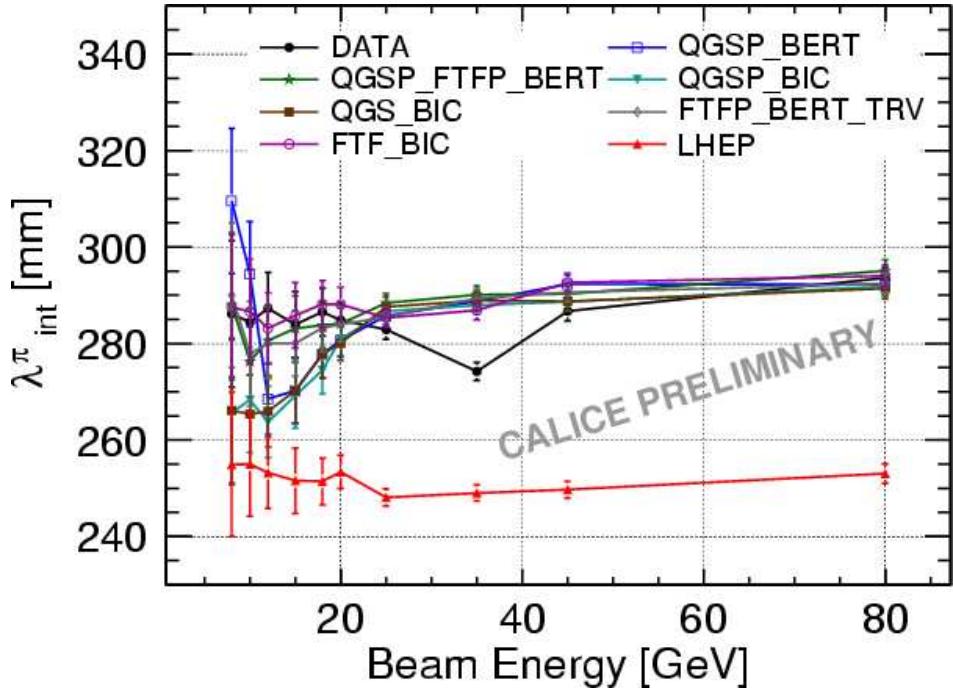


Figure 12: Extracted values for the interaction length in cm from data and MC models. Error shown are the uncertainty from the fitting added in quadrature with the uncertainty of the PTF to find the first interaction layer (see text). There is a problem with the 35 GeV run currently under investigation.

## 403 C Software Versions and Calibration Constants

404 The data and Monte Carlo simulations have been reconstructed and digitized using a custom  
 405 software version since there was no official tagged version available at the time the data was  
 406 processed. The software is nevertheless very close to the official tagged version v03-00 and  
 407 should be equivalent.

408 Table 5 shows the database folder used during the reconstruction and digitization.

Collection Name	Data Base Path	Tag
AhcMipConstants	/cd_calice/Hcal/mip_constants	ahc_mip_constants_002
AhcMipSlopes	/cd_calice/Hcal/mip_slopes	ahc_mip_slopes_002
AhcGainConstants	/cd_calice/Hcal/gain_constants	ahc_gain_constants_002
AhcGainSlopes	/cd_calice/Hcal/gain_slopes	ahc_gain_slopes_002
AhcInterConstants	/cd_calice/Hcal/ic_constants	ahc_ic_constants_002
AhcCellQuality	/cd_calice/Ahc/BadCellMap	ahc_bad_cell_map_001
AhcTempSensorCalib	/cd_calice/Hcal/tempSensors	ahc_tempSensors_001
AhcSroModMapping	/cd_calice_cernbeam/Hcal/AhcSroModMapping	ahc_AhcSroModMapping_001
AhcResponseCurve	/cd_calice/Ahc/ResponseCurve	ahc_response_curve_002
AhcResponseScaling	/cd_calice/Ahc/ResponseCurveScaling	ahc_response_curve_scaling_001

Table 5: Database folders and versions tags of the calibration constants used.

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