

Tau Neutrino Physics at DUNE

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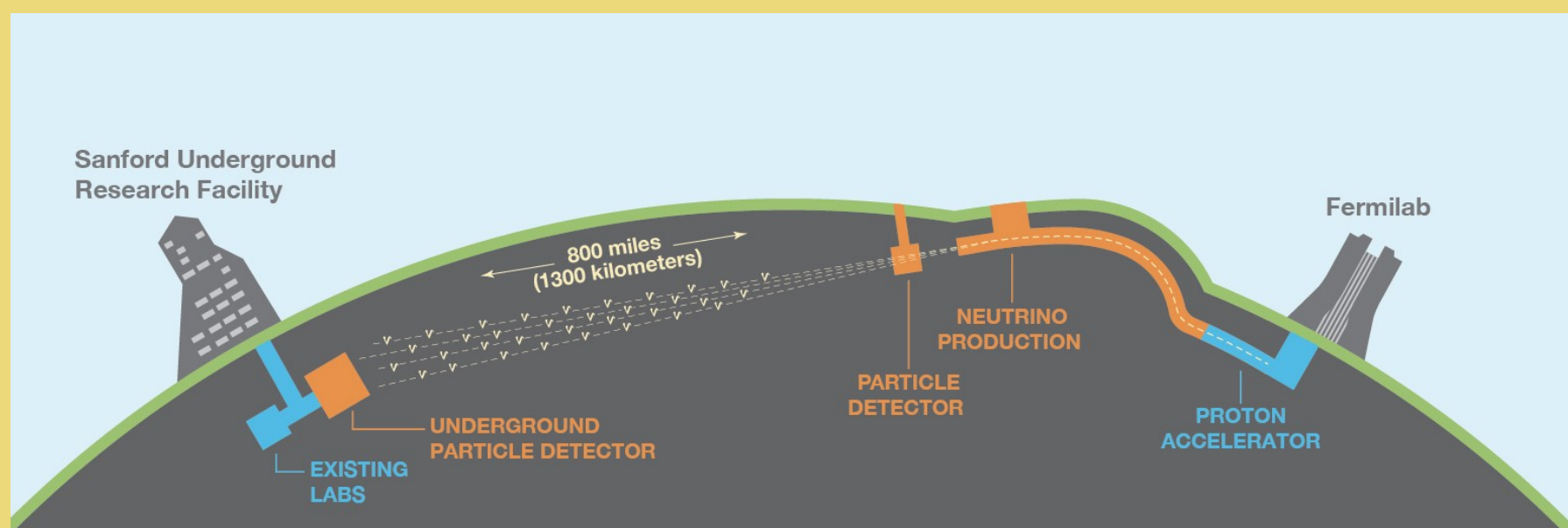


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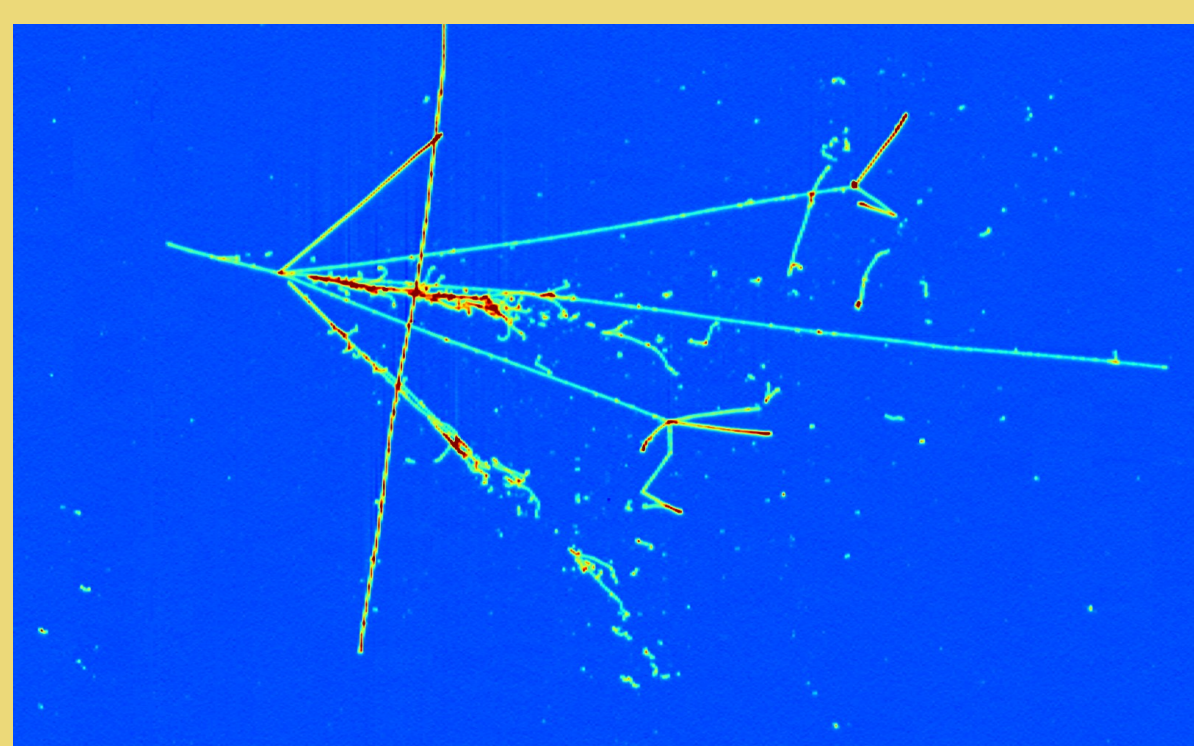


1) DUNE



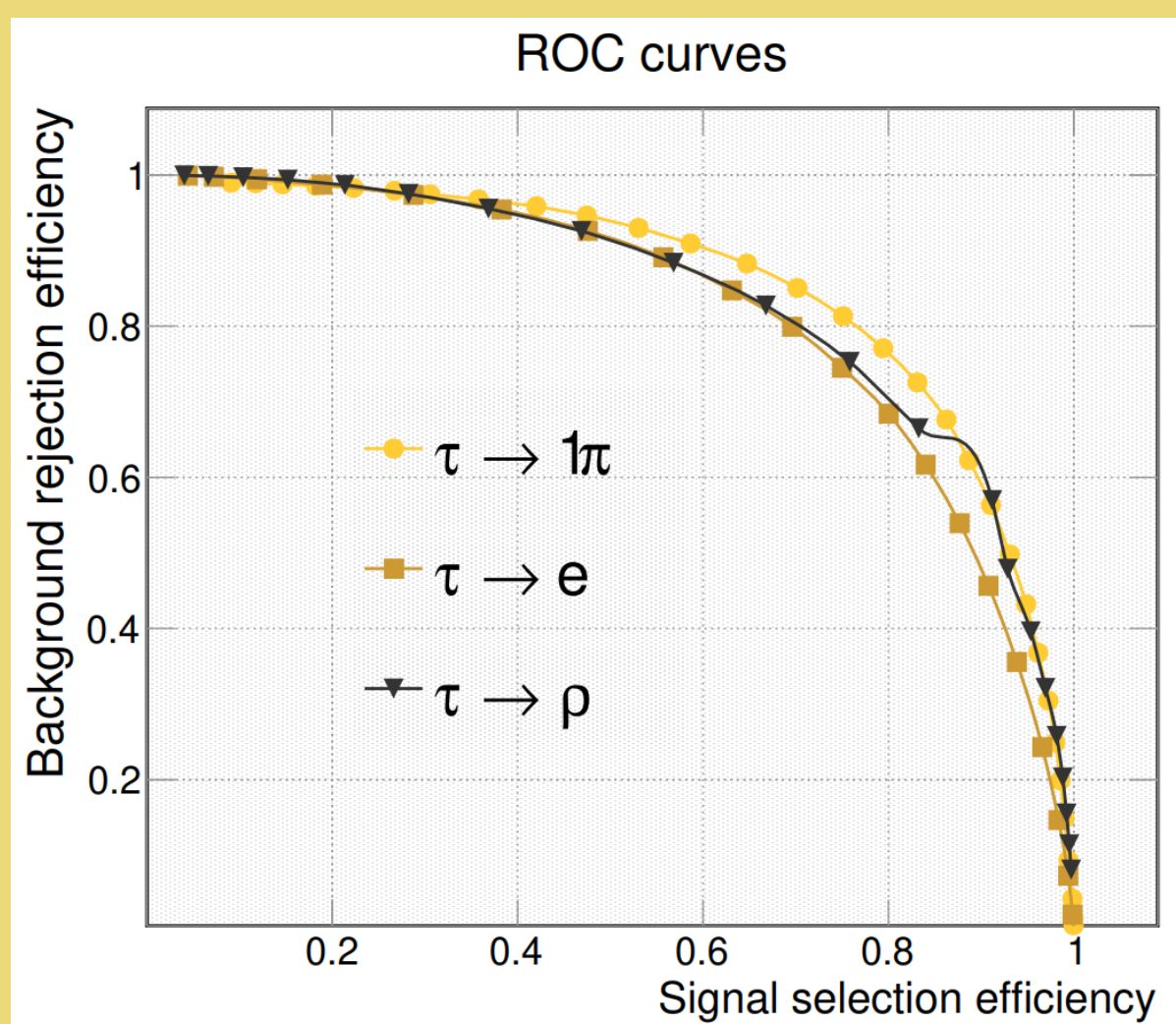
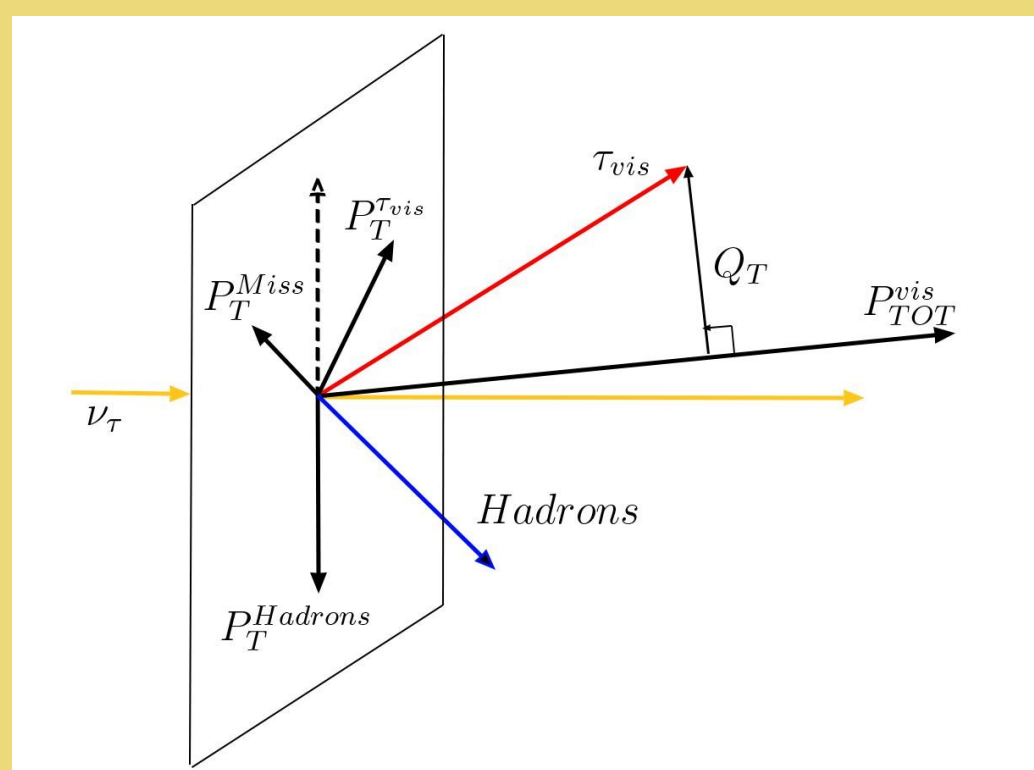
- Far Detector
 - 1300 km baseline
 - Liquid argon time projection chamber (LArTPC) technology for high resolution neutrino interaction imaging
 - 4x17 kton LArTPC modules
- Near Detector
 - 574 m baseline
 - Multiple detector systems
 - 147 ton LArTPC component

- DUNE is a long-baseline neutrino experiment currently under construction
- Will constrain the three flavor paradigm
 - Measure δ_{CP} and mass ordering by studying $\nu_\mu \rightarrow \nu_e$ oscillations

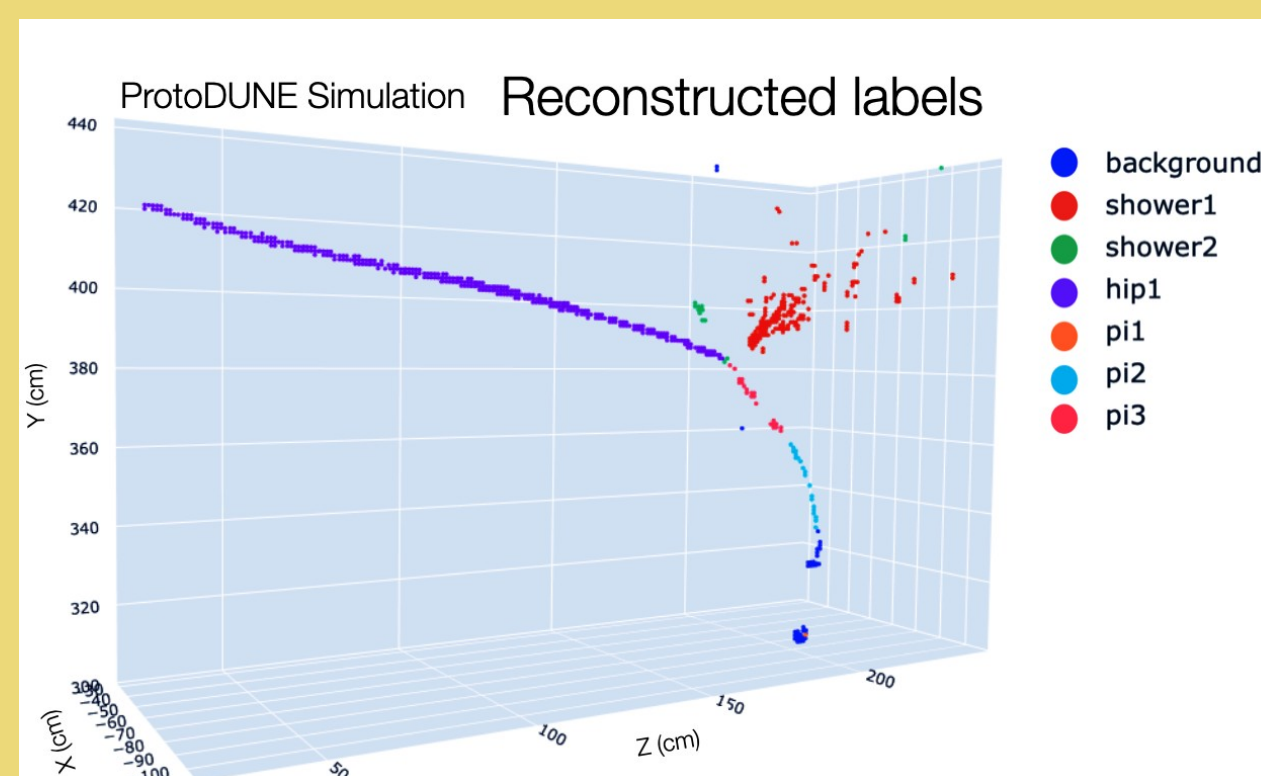
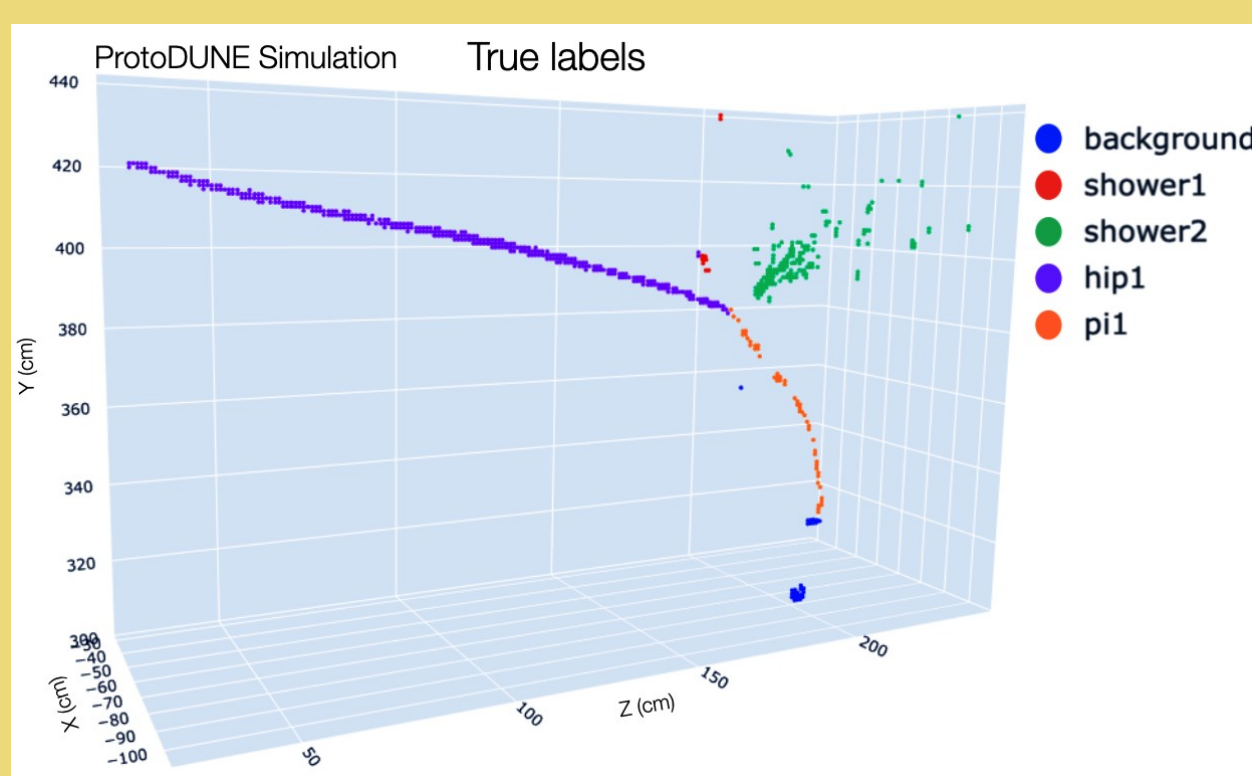


4) Selection

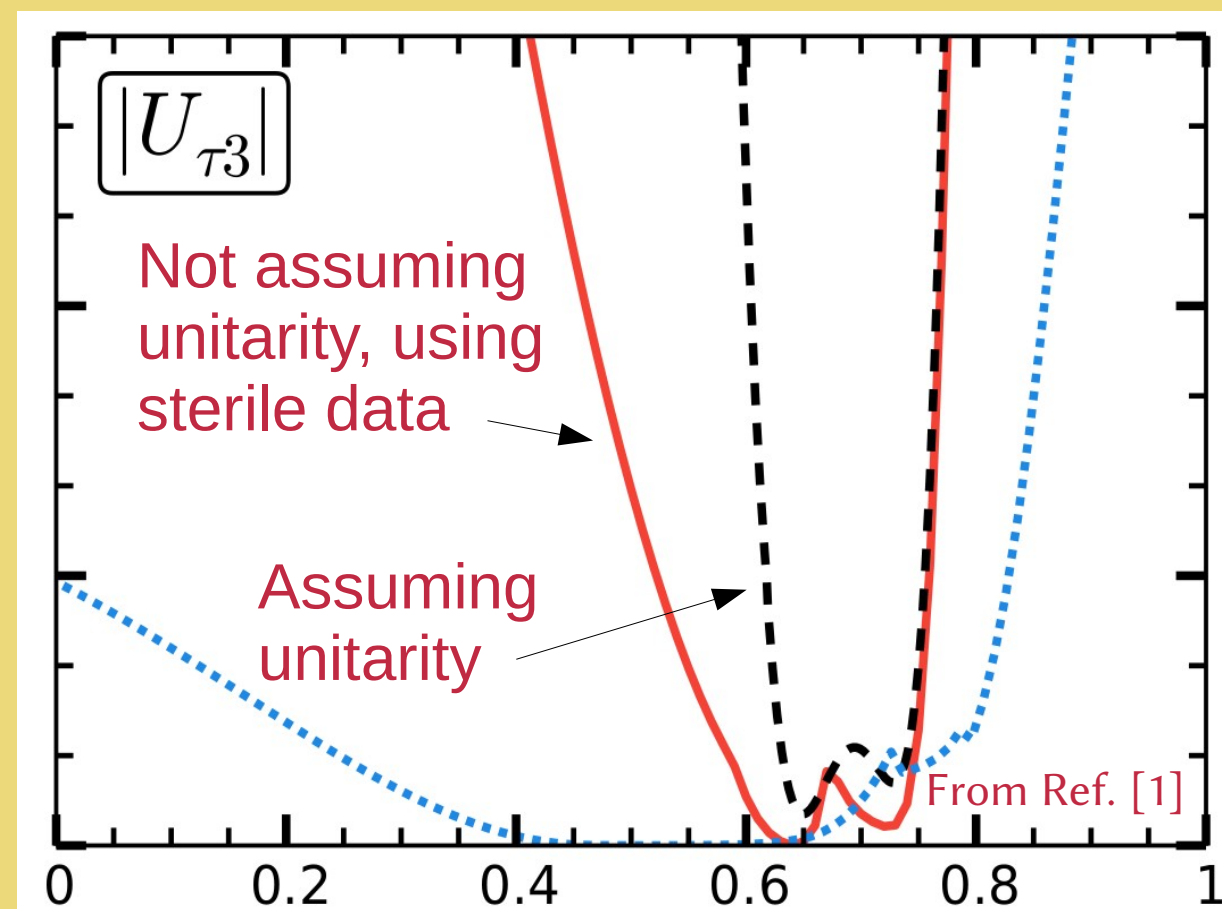
- Truth-level study of atmospheric tau neutrinos suggested excellent hadronic ν_τ /NC discrimination using simple kinematic cuts [2]
- Optimistic assumption: near perfect e/γ and μ/π discrimination in LArTPC
- Suggests 30% signal efficiency and 0.5% NC background efficiency possible
- Use as a first estimate of sensitivity



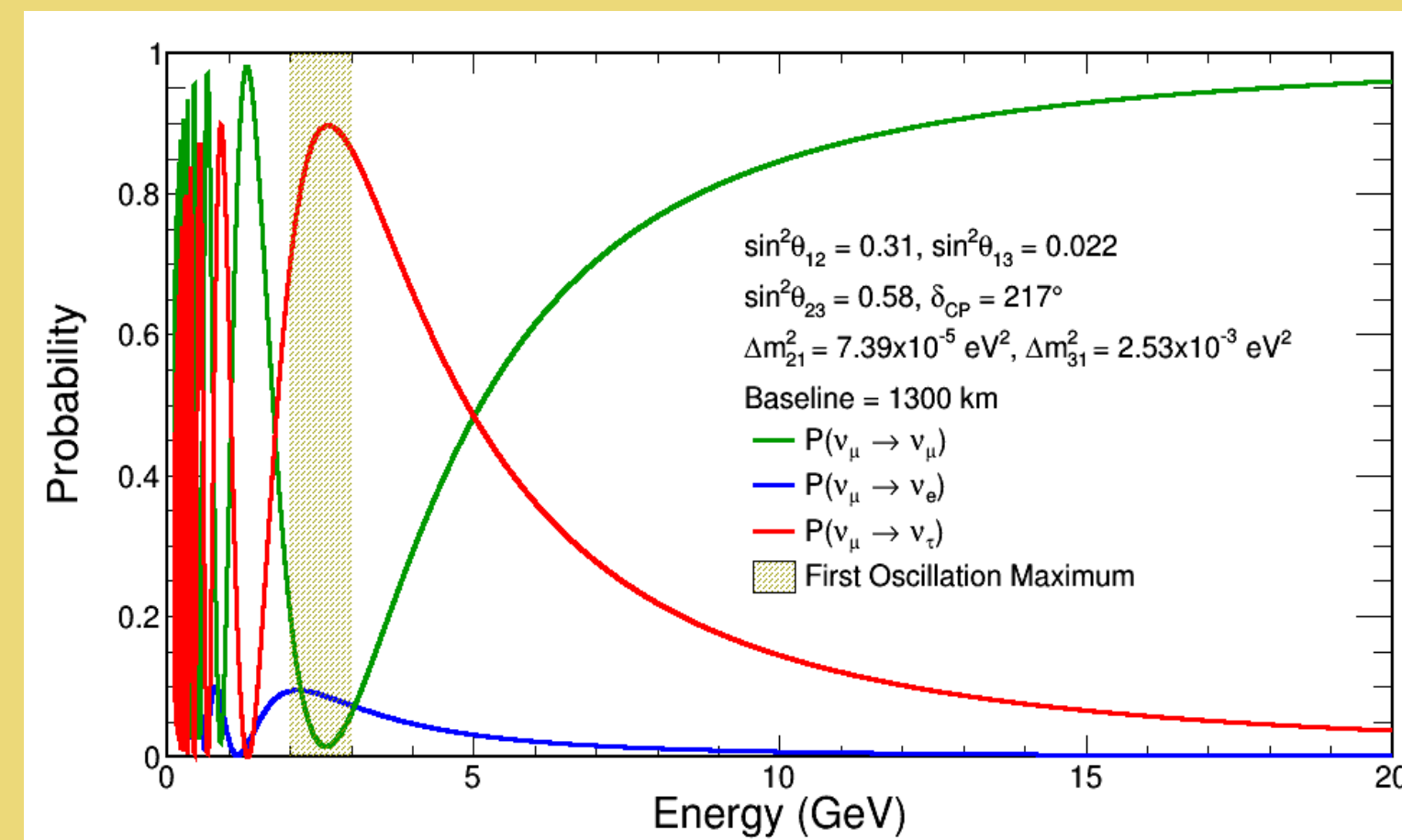
- More realistic reconstruction relies on the transverse plane kinematic approach suggested in Ref. [3]
- Can be applied to hadronic decays if the tau-lepton products can be identified
 - First attempt in Ref. [4] applied approach to $\tau \rightarrow \rho$ decays
- Investigating machine learning approaches for improved particle ID and reconstruction



2) Why Tau Neutrinos?

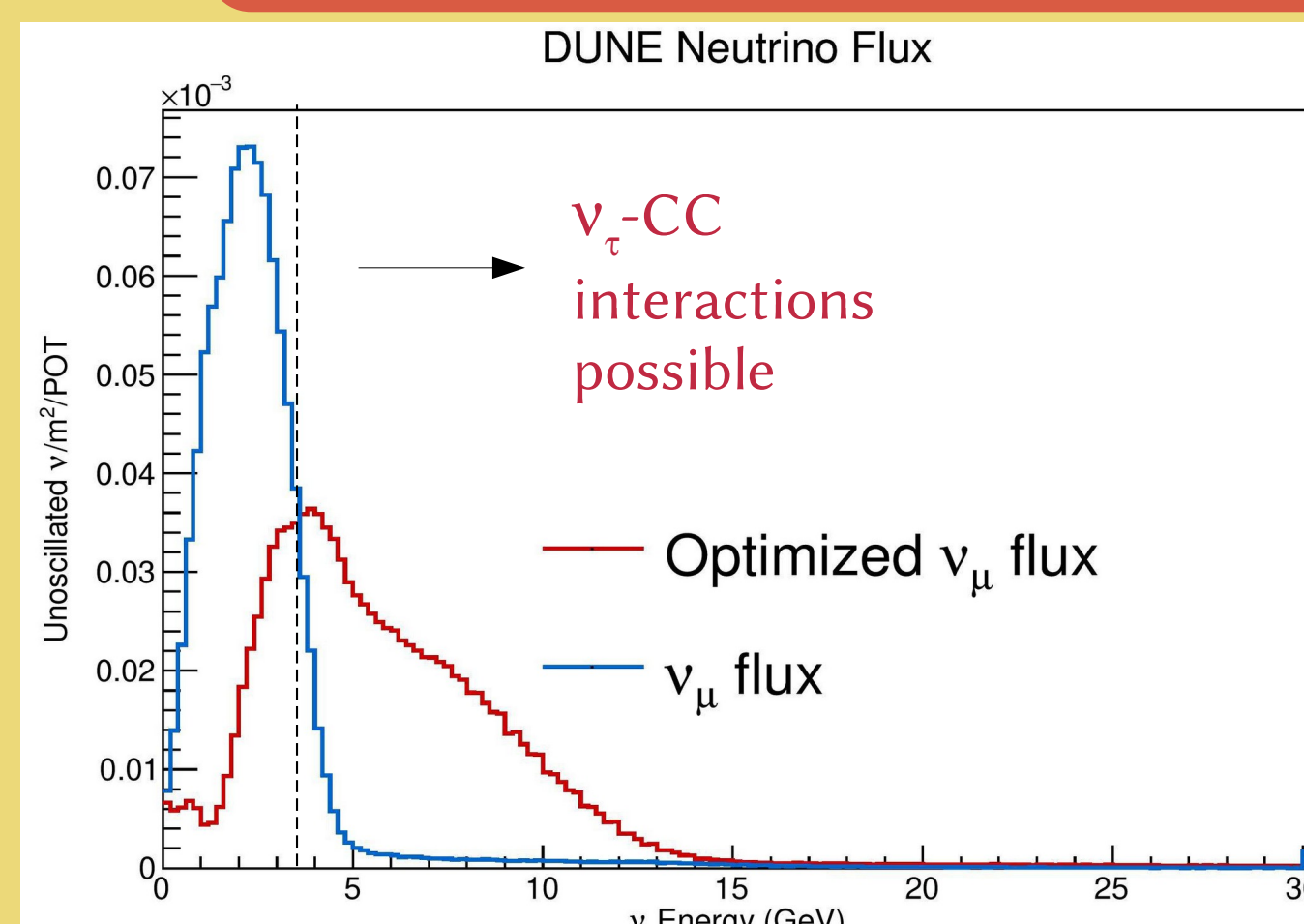


- Current generation of neutrino experiments provides nearly complete description of three flavor paradigm
- Almost all knowledge of tau neutrino sector is taken from
 - Lepton universality for cross sections
 - PMNS unitarity for oscillations
- Critical that these assumptions are tested

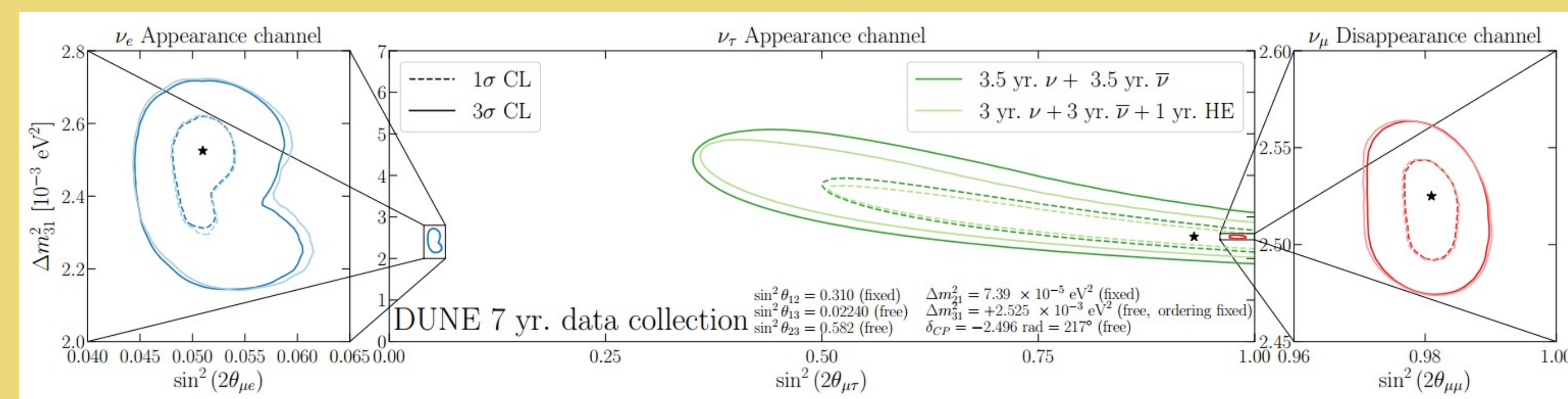


- At atmospheric maximum, almost all muon neutrinos oscillate to tau neutrinos
- Excellent opportunity to probe:
 - Unitarity by measuring all three oscillation modes
 - Standard model cross section assumptions

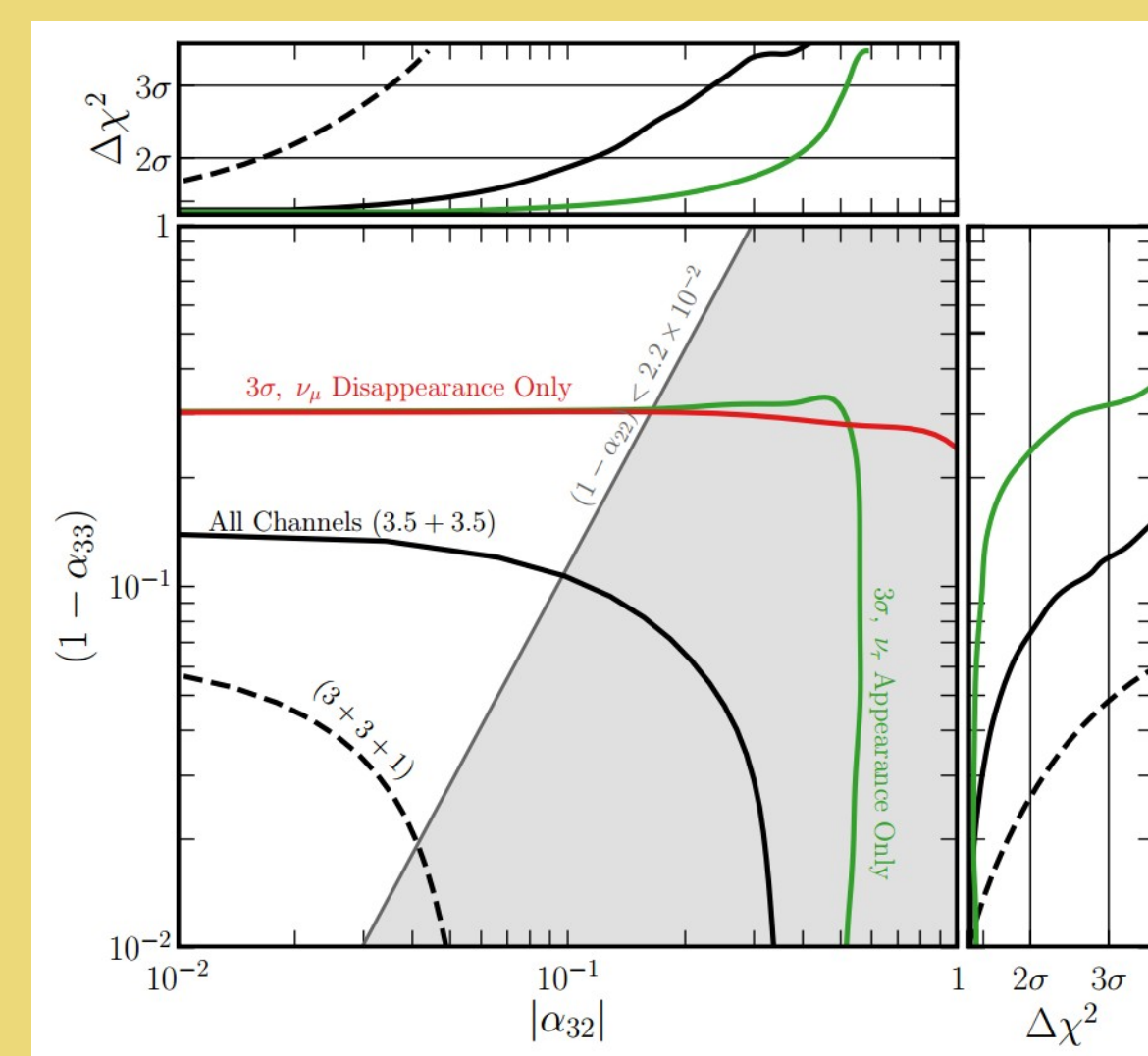
5) Long-Baseline Oscillations



- Default beam configuration peaks ~ 2 GeV to maximize sensitivity to CP violation
- High energy tail is above kinematic threshold
- Expected counts/year (1.2 MW beam)
 - ~ 130 ν_τ in neutrino mode
 - ~ 30 $\bar{\nu}_\tau$ in antineutrino mode
- Tau optimized configuration
 - Higher energy
 - Possible configuration after CP program
 - ~ 800 ν_τ per year



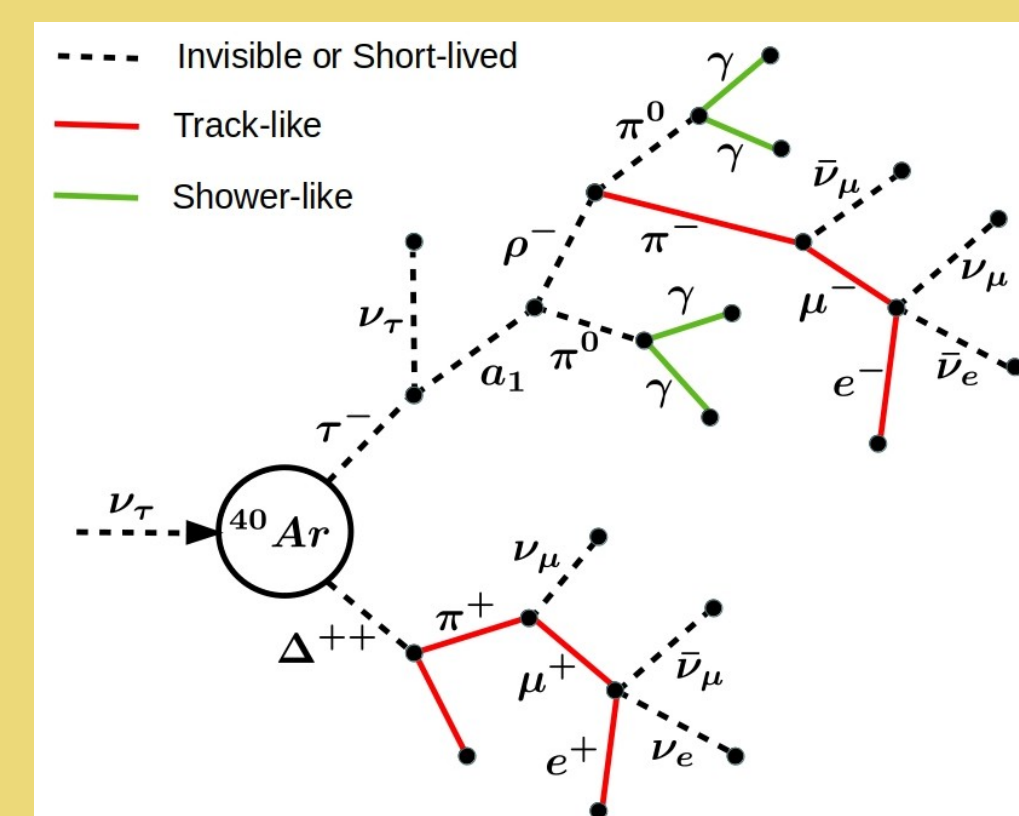
- Ref. [6] showed that DUNE can constrain normalization of 3^rd PMNS column to $\sim 5\%$ in model independent way
- If we assume non-unitarity is due to kinematically inaccessible sterile neutrinos, can set strong limits on α_{33}
- Tau optimized configuration is particularly useful for this



3) Tau Neutrino Challenges

τ^- Decay Mode	Branching Ratio
$\mu^- \bar{\nu}_\mu \nu_\tau$	17.4%
$e^- \bar{\nu}_e \nu_\tau$	17.8%
$\pi^- \nu_\tau$	10.8%
$\pi^- \pi^0 \nu_\tau$	25.5%
$\pi^- 2\pi^0 \nu_\tau$	9.3%
$2\pi^- \pi^0 \nu_\tau$	9.3%
$2\pi^- \pi^+ \pi^0 \nu_\tau$	4.6%

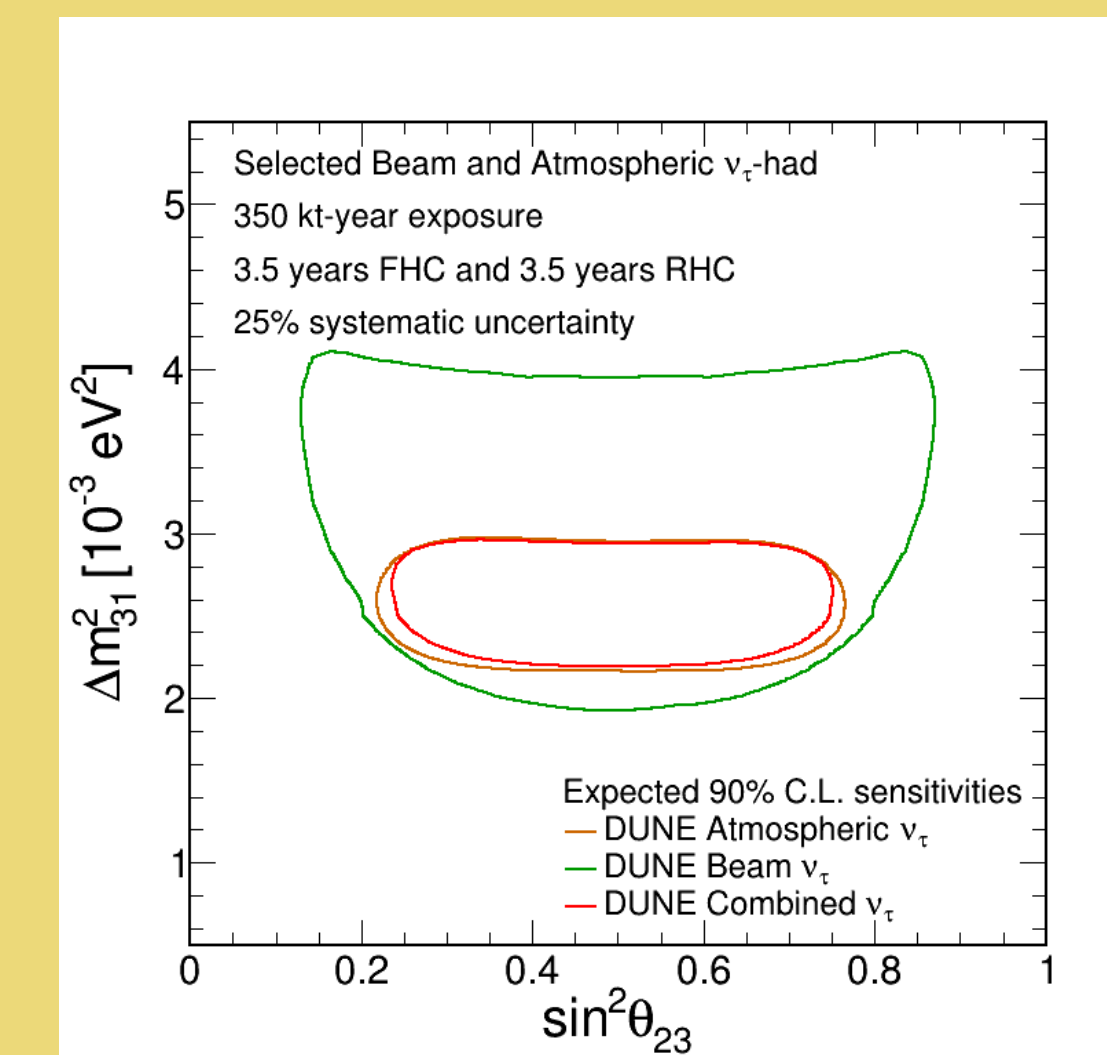
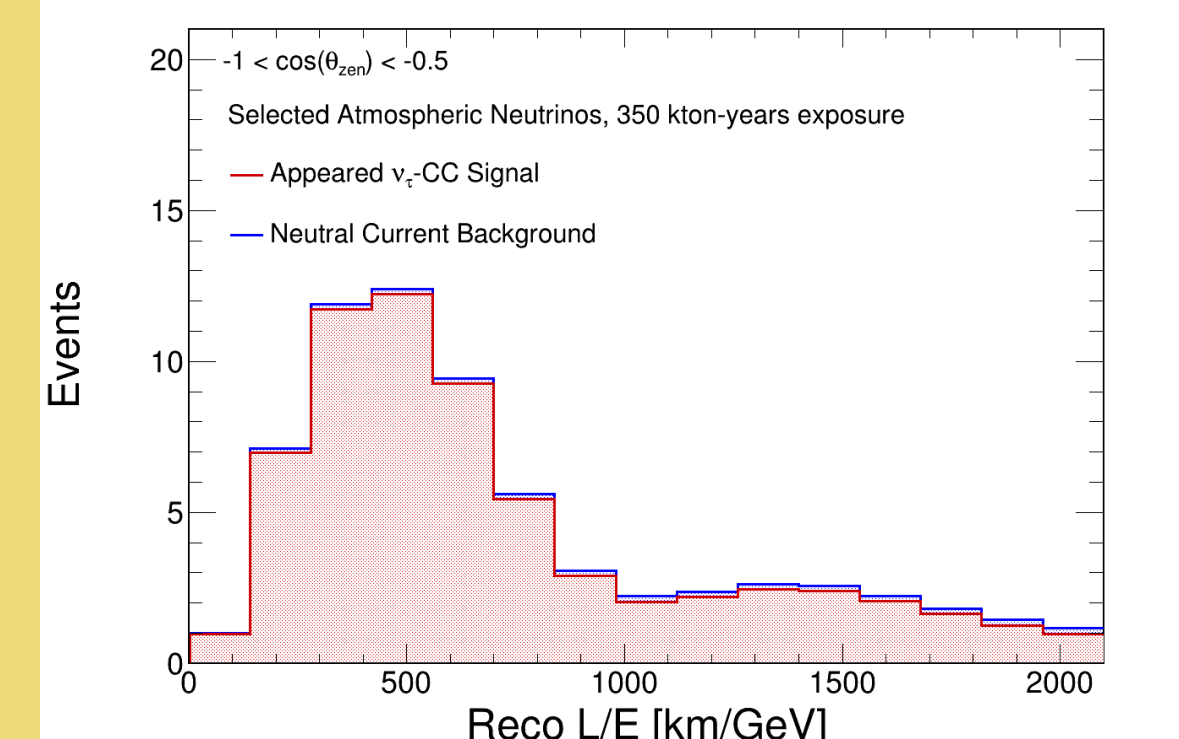
- Kinematically forbidden at typical beam energies
- Even above threshold, still suppressed
- Tau leptons have many decay modes
- Mimic ν_e CC, ν_μ CC, or NC events depending on decay



- Outgoing neutrino \rightarrow missing energy
- Worse for leptonic decay modes
- Hadronic decay modes can be complicated
- Difficult to separate hadronic systems from tau decay and nucleus

6) Atmospheric Oscillations

- Due to kinematic threshold, beam ν_τ are only detected above the atmospheric oscillation maximum
- Causes a degeneracy between Δm_{31}^2 and $\sin^2 \theta_{23}$
- Due to long baseline of atmospheric neutrinos, atmospheric maximum is above kinematic threshold
- Complements beam neutrinos



- Expect ~ 1 ν_τ / kton-year
- Lower statistics, but clear signal at atmospheric maximum
 - Breaks degeneracy
- Interplay between atmospheric and beam important
 - Atmospherics have favorable L/E
 - Beam has defined direction and Near Detector for systematic uncertainty reduction

See posters:

#199 Studies of tau neutrinos appearing at the DUNE Near Detector Complex

#418 Tau Neutrino Cross Section at DUNE
for details about the tau neutrino program at DUNE

References

- [1] S. Parke and M. Ross-Lonergan, Phys. Rev. D 93, 1103009 (2016)
- [2] J. Conrad, A. de Gouvea, S. Shalgar, J. Spitz, Phys. Rev. D 82, 093012 (2010)
- [3] C. Albright and R. Shrock, Phys. Lett. B 84, 123 (1979)
- [4] T. Kosc, PhD thesis, Université de Lyon (2021)
- [5] R. Mammen Abraham et al, arXiv:2202.05591
- [6] A. de Gouvea, K. Kelly, G. Stenico, P. Pasquini, Phys. Rev. D 100, 016004 (2019)