

M_h in the MSSM-seesaw scenario with ILC precision

S. HEINEMEYER¹, M.J. HERRERO², S. PEÑARANDA^{3,4}, A.M. RODRÍGUEZ-SÁNCHEZ²

1- Instituto de Física de Cantabria (CSIC-UC), Santander, Spain

2- Departamento de Física Teórica and Instituto de Física Teórica, UAM/CSIC
Universidad Autónoma de Madrid, Cantoblanco, Madrid, Spain

3- Departamento de Física Teórica, Universidad de Zaragoza, Zaragoza, Spain

4- Departament de Física Fonamental, Universitat de Barcelona, Spain

We review the computation of the one-loop radiative corrections from the neutrino/ sneutrino sector to the lightest Higgs boson mass, M_h , within the context of the so-called MSSM-seesaw scenario. This model introduces right handed neutrinos and their supersymmetric partners, the sneutrinos, including Majorana mass terms. We find negative and sizeable corrections to M_h , up to -5 GeV for a large Majorana scale, $10^{13} - 10^{15}$ GeV, and for the lightest neutrino mass in a range $0.1 - 1$ eV. The corrections to M_h are substantially larger than the anticipated ILC precision for large regions of the MSSM-seesaw parameter space.

1 Introduction

The current experimental data on neutrino mass differences and neutrino mixing angles clearly indicate new physics beyond the so far successful Standard Model (SM) of Particle Physics. In particular, neutrino oscillations imply that at least two generations of neutrinos must be massive. Therefore, one needs to extend the SM to incorporate neutrino mass terms.

We have explored [4] the simplest version of a Supersymmetric extension of the SM, the well known Minimal Supersymmetric Standard Model (MSSM), extended by right-handed Majorana neutrinos and where the seesaw mechanism of type I [2] is implemented to generate the small neutrino masses. For simplicity, as a first step, we focus here in the one generation case.

On the other hand, it is well known that heavy Majorana neutrinos, with a Majorana mass scale $m_M \sim 10^{13} - 10^{15}$ GeV, induce large LFV rates [3], due to their potentially large Yukawa couplings to the Higgs sector. For the same reason, radiative corrections to Higgs boson masses due to such heavy Majorana neutrinos could also be relevant. Consequently, our study has been focused on the radiative corrections to the lightest MSSM \mathcal{CP} -even h boson mass, M_h , due to the one-loop contributions from the neutrino/sneutrino sector within the MSSM-seesaw framework.

In the following we briefly review the main relevant aspects of the calculation of the mass corrections and the numerical results. Further details can be found in [4], where also an extensive list with references to previous works can be found.

2 The MSSM-seesaw model

The MSSM-seesaw model with one neutrino/sneutrino generation is described in terms of the well known MSSM superpotential plus the new relevant terms given as:

$$W = \epsilon_{ij} \left[Y_\nu \hat{H}_2^i \hat{L}^j \hat{N} - Y_l \hat{H}_1^i \hat{L}^j \hat{R} \right] + \frac{1}{2} \hat{N} m_M \hat{N}, \quad (1)$$

where $\hat{N} = (\tilde{\nu}_R^*, (\nu_R)^c)$ is the additional superfield that contains the right-handed neutrino ν_R and its scalar partner $\tilde{\nu}_R$.

There are also new relevant terms in the soft SUSY breaking potential:

$$V_{\text{soft}}^{\tilde{\nu}} = m_L^2 \tilde{\nu}_L^* \tilde{\nu}_L + m_R^2 \tilde{\nu}_R^* \tilde{\nu}_R + (Y_\nu A_\nu H_2^2 \tilde{\nu}_L \tilde{\nu}_R^* + m_M B_\nu \tilde{\nu}_R \tilde{\nu}_R + \text{h.c.}) . \quad (2)$$

After electro-weak (EW) symmetry breaking, the charged lepton and Dirac neutrino masses can be written as

$$m_l = Y_l v_1 , \quad m_D = Y_\nu v_2 , \quad (3)$$

where v_i are the vacuum expectation values (VEVs) of the neutral Higgs scalars, with $v_{1(2)} = v \cos(\sin)\beta$ and $v \simeq 174$ GeV.

The 2×2 neutrino mass matrix is given in terms of m_D and m_M by:

$$M^\nu = \begin{pmatrix} 0 & m_D \\ m_D & m_M \end{pmatrix} . \quad (4)$$

Diagonalization of M^ν leads to two mass eigenstates, which are Majorana fermions with the respective mass eigenvalues given by:

$$m_{\nu, N} = \frac{1}{2} \left(m_M \mp \sqrt{m_M^2 + 4m_D^2} \right) . \quad (5)$$

In the seesaw limit, i.e. when $\xi \equiv \frac{m_D}{m_M} \ll 1$, one finds,

$$m_\nu = -m_D \xi + \mathcal{O}(m_D \xi^3) \simeq -\frac{m_D^2}{m_M} , \quad m_N = m_M + \mathcal{O}(m_D \xi) \simeq m_M . \quad (6)$$

Regarding the sneutrino sector, the sneutrino mass matrices for the \mathcal{CP} -even, \tilde{M}_+ , and the \mathcal{CP} -odd, \tilde{M}_- , subsectors are given respectively by

$$\tilde{M}_\pm^2 = \begin{pmatrix} m_L^2 + m_D^2 + \frac{1}{2} M_Z^2 \cos 2\beta & m_D (A_\nu - \mu \cot \beta \pm m_M) \\ m_D (A_\nu - \mu \cot \beta \pm m_M) & m_R^2 + m_D^2 + m_M^2 \pm 2B_\nu m_M \end{pmatrix} . \quad (7)$$

The diagonalization of these two matrices, \tilde{M}_\pm^2 , leads to four sneutrino mass eigenstates. In the seesaw limit, where m_M is much bigger than all the other scales the corresponding sneutrino masses are given by:

$$\begin{aligned} m_{\tilde{\nu}_+, \tilde{\nu}_-}^2 &= m_L^2 + \frac{1}{2} M_Z^2 \cos 2\beta \mp 2m_D (A_\nu - \mu \cot \beta - B_\nu) \xi , \\ m_{\tilde{N}_+, \tilde{N}_-}^2 &= m_M^2 \pm 2B_\nu m_M + m_R^2 + 2m_D^2 . \end{aligned} \quad (8)$$

Finally, in the interaction Lagrangian that is relevant for the present work, there are terms already present in the MSSM: the pure gauge interactions between the left-handed neutrinos and the Z boson, those between the 'left-handed' sneutrinos and the Higgs bosons, and those between the 'left-handed' sneutrinos and the Z bosons. In addition, in this MSSM-seesaw scenario, there are interactions driven by the neutrino Yukawa couplings (or equivalently m_D since $Y_\nu = (gm_D)/(\sqrt{2}M_W \sin \beta)$), as for instance $g_{h\nu_L \nu_R} = -\frac{igm_D \cos \alpha}{2M_W \sin \beta}$, and new interactions due to the Majorana nature driven by m_M , which are not present in the case of Dirac fermions, as for instance $g_{h\tilde{\nu}_L \tilde{\nu}_R} = -\frac{igm_D m_M \cos \alpha}{2M_W \sin \beta}$. Besides, the Higgs boson sector in the MSSM-seesaw model is as in the MSSM.

3 Calculation

In the Feynman diagrammatic (FD) approach the higher-order corrected \mathcal{CP} -even Higgs boson masses in the MSSM, denoted here as M_h and M_H , are derived by finding the poles of the (h, H) -propagator matrix, which is equivalent to solving the following equation [1]:

$$\left[p^2 - m_h^2 + \hat{\Sigma}_{hh}(p^2) \right] \left[p^2 - m_H^2 + \hat{\Sigma}_{HH}(p^2) \right] - \left[\hat{\Sigma}_{hH}(p^2) \right]^2 = 0 , \quad (9)$$

where $m_{h,H}$ are the tree level masses. The one loop renormalized self-energies, $\hat{\Sigma}_{\phi\phi}(p^2)$, in (9) can be expressed in terms of the bare self-energies, $\Sigma_{\phi\phi}(p^2)$, the field renormalization constants $\delta Z_{\phi\phi}$ and the mass counter terms δm_ϕ^2 , where ϕ stands for h, H . For example, the lightest Higgs boson renormalized self energy reads:

$$\hat{\Sigma}_{hh}(p^2) = \Sigma_{hh}(p^2) + \delta Z_{hh}(p^2 - m_h^2) - \delta m_h^2. \quad (10)$$

Regarding the renormalization prescription, we have used an on-shell renormalization scheme for M_Z, M_W and M_A mass counterterms and T_h, T_H tadpole counterterms. On the other hand, we have used a modified $\overline{\text{DR}}$ scheme ($\text{m}\overline{\text{DR}}$) for the renormalization of the wave function and $\tan \beta$. The $\text{m}\overline{\text{DR}}$ scheme is very similar to the well known $\overline{\text{DR}}$ scheme but instead of subtracting the usual term proportional to $\Delta = \frac{2}{\epsilon} - \gamma_E + \log(4\pi)$ one subtracts the term proportional to $\Delta_m = \Delta - \log(m_M^2/\mu_{\overline{\text{DR}}}^2)$, hence, avoiding large logarithms of the large scale m_M . As studied in other works [5], this scheme minimizes higher order corrections when two very different scales are involved in a calculation of radiative corrections.

The full one-loop $\nu/\tilde{\nu}$ corrections to the self-energies, $\hat{\Sigma}_{hh}^{\nu/\tilde{\nu}}$, $\hat{\Sigma}_{HH}^{\nu/\tilde{\nu}}$ and $\hat{\Sigma}_{hH}^{\nu/\tilde{\nu}}$, entering (9) have been evaluated with *FeynArts* and *FormCalc* [6]. The new Feynman rules for the $\nu/\tilde{\nu}$ sector are inserted into a new model file. Since we are interested in exploring the relevance of the new radiative corrections to M_h from the neutrino/sneutrino sector, we will present here our results in terms of the mass difference with respect to the MSSM prediction. Consequently, we define,

$$\Delta m_h^{\text{m}\overline{\text{DR}}} := M_h^{\nu/\tilde{\nu}} - M_h, \quad (11)$$

where $M_h^{\nu/\tilde{\nu}}$ denotes the pole for the light Higgs mass including the $\nu/\tilde{\nu}$ corrections (i.e. in the MSSM-seesaw model), and M_h the corresponding pole in the MSSM, i.e without the $\nu/\tilde{\nu}$ corrections. Thus, for a given set of input parameters we first calculate M_h in the MSSM with the help of *FeynHiggs* [7], such that all relevant known higher-order corrections are included. Then we add the new contributions from the neutrino/sneutrino sector and eventually compute $\Delta m_h^{\text{m}\overline{\text{DR}}}$.

4 Results

We have obtained the full analytical results for the renormalized Higgs boson self-energies and their expressions in the seesaw limit. In order to understand in simple terms the analytical behavior of our full numerical results we have expanded the renormalized self-energies in powers of the seesaw parameter $\xi = m_D/m_M$:

$$\hat{\Sigma}(p^2) = \left(\hat{\Sigma}(p^2)\right)_{m_D^0} + \left(\hat{\Sigma}(p^2)\right)_{m_D^2} + \left(\hat{\Sigma}(p^2)\right)_{m_D^4} + \dots \quad (12)$$

The zeroth order of this expansion is precisely the pure gauge contribution and it does not depend on m_D or m_M . Therefore, it corresponds to the result in the MSSM. The rest of the terms of the expansion are the Yukawa contribution. The leading term of this Yukawa contribution is the $\mathcal{O}(m_D^2)$ term, because it is the only one not suppressed by the Majorana scale. In fact it goes as $Y_\nu^2 M_{\text{EW}}^2$, where M_{EW}^2 denotes generically the electroweak scales involved, concretely, p^2 , M_Z^2 and M_A^2 . In particular, the $\mathcal{O}(p^2 m_D^2)$ terms of the renormalized self-energy, which turn out to be the most relevant leading contributions, separated into the neutrino and sneutrino contributions, read:

$$\hat{\Sigma}_{hh}^{\text{m}\overline{\text{DR}}} \Big|_{m_D^2 p^2} \sim \left[\text{neutrino loop} + \text{sneutrino loop} \right] \Big|_{m_D^2 p^2} \sim \frac{g^2 p^2 m_D^2 c_\alpha^2}{64\pi^2 M_W^2 s_\beta^2} + \frac{g^2 p^2 m_D^2 c_\alpha^2}{64\pi^2 M_W^2 s_\beta^2}. \quad (13)$$

Notice that the above neutrino contributions come from the Yukawa interaction $g_{h\nu_L\nu_R}$, which is extremely suppressed in the Dirac case but can be large in the Majorana case. The sneutrino contributions come from the new couplings $g_{h\tilde{\nu}_L\tilde{\nu}_R}$, which are not present in the Dirac case. It is also interesting to remark that these terms, being $\sim p^2$, depend on the external momentum. Therefore, at large m_M , to keep just the Yukawa part is a good approximation, but to neglect the momentum dependence or to set the external momentum to zero are certainly not. In consequence, the effective potential method will not provide a realistic result

for the radiative corrections to the Higgs mass. Similarly, obtaining the leading logarithmic terms in a RGE computation, would also miss these finite terms.

The behaviour of the renormalized self-energy with all others parameters entering in the computation have been discussed in [4]. According to our detailed analysis in this paper, the most relevant parameters for our purposes are: m_M (or, equivalently, the heaviest physical Majorana neutrino mass m_N), m_ν and the soft SUSY breaking parameters $m_{\tilde{R}}$ and B_ν . In the literature it is often assumed that m_M has a very large value, $m_M \sim \mathcal{O}(10^{14-15})$ GeV, in order to get $|m_\nu| \sim 0.1 - 1$ eV with large Yukawa couplings $Y_\nu \sim \mathcal{O}(1)$. This is an interesting possibility since it can lead to important phenomenological implications due to the large size of the radiative corrections driven by these large Y_ν . We have explored, however, not only these extreme values but the full range for m_M : $\sim 10^2 - 10^{15}$ GeV.

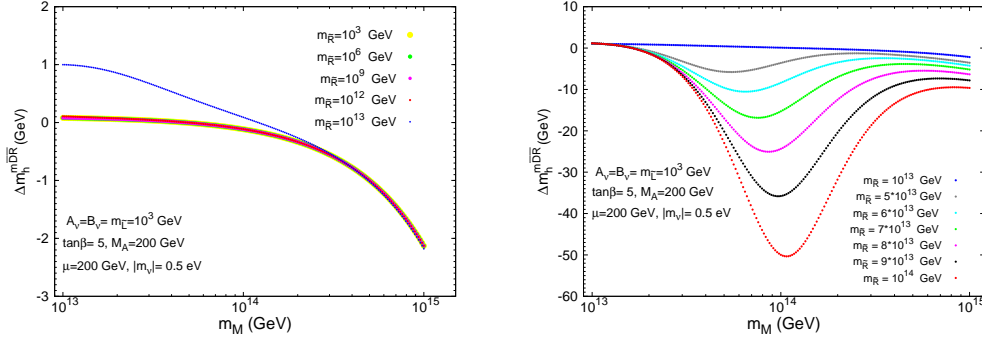


Figure 1: One-loop corrections to M_h from the neutrino/sneutrino sector as a function of m_M for: $m_{\tilde{R}} < 10^{13}$ GeV (left panel) and $10^{13} \text{ GeV} < m_{\tilde{R}} < 10^{14}$ GeV (right panel).

Fig. 1 shows the predictions for Δm_h^{mDR} as a function of m_M , for several input $m_{\tilde{R}}$ values. The Higgs mass corrections are positive and below 0.1 GeV if $m_M \lesssim 5 \times 10^{13}$ GeV and $m_{\tilde{R}} < 10^{12}$ GeV (left panel). For larger Majorana mass values, the corrections get negative and grow up to a few GeV; $\Delta m_h^{\text{mDR}} = -2.15$ GeV for $m_M = 10^{15}$ GeV. The results in the right plot show that for larger values of the soft mass, $m_{\tilde{R}} \gtrsim 10^{13}$ GeV, the Higgs mass corrections are negative and can be sizeable, a few tens of GeV, reaching their maximum values at $m_{\tilde{R}} \simeq m_M$.

The results of the Higgs mass corrections in terms of the two relevant physical Majorana neutrino masses, light $|m_\nu|$ and heavy m_N , are summarized in the left plot of Fig. 2. For values of $m_N < 3 \times 10^{13}$ GeV and $|m_\nu| < 0.1 - 0.3$ eV the corrections to M_h are positive and smaller than 0.1 GeV. In this region, the gauge contribution dominates. In fact, the wider black contour line with fixed $\Delta m_h^{\text{mDR}} = 0.09$ coincides with the prediction for the case where just the gauge part in the self-energies have been included. This means that 'the distance' of any other contour-line respect to this one represents the difference in the radiative corrections respect to the MSSM prediction. For larger values of m_N and/or $|m_\nu|$ the Yukawa part dominates, and the radiative corrections become negative and larger in absolute value, up to about -5 GeV in the right upper corner of this figure. These corrections grow in modulus proportionally to m_M and m_ν , due to the fact that the seesaw mechanism impose a relation between the three masses involved, $m_D^2 = |m_\nu| m_N$.

Finally, we present in the right plot of Fig. 2 the contour-lines for fixed Δm_h^{mDR} in the less conservative case where $m_{\tilde{R}}$ is close to m_M . These are displayed as a function of $|m_\nu|$ and the ratio $m_{\tilde{R}}/m_M$. m_M is fixed to 10^{14} GeV. For the interval studied here, we see again that the radiative corrections can be negative and as large as tens of GeV in the upper right corner of the plot. For instance, $\Delta m_h^{\text{mDR}} = -30$ GeV for $|m_\nu| = 0.6$ eV and $m_{\tilde{R}}/m_M = 0.7$.

To summarize: for some regions of the MSSM-seesaw parameter space, the corrections to M_h are of the order of several GeV. For all soft SUSY-breaking parameters at the TeV scale we find correction of up to -5 GeV to M_h . These corrections are substantially larger than the anticipated ILC precision of about

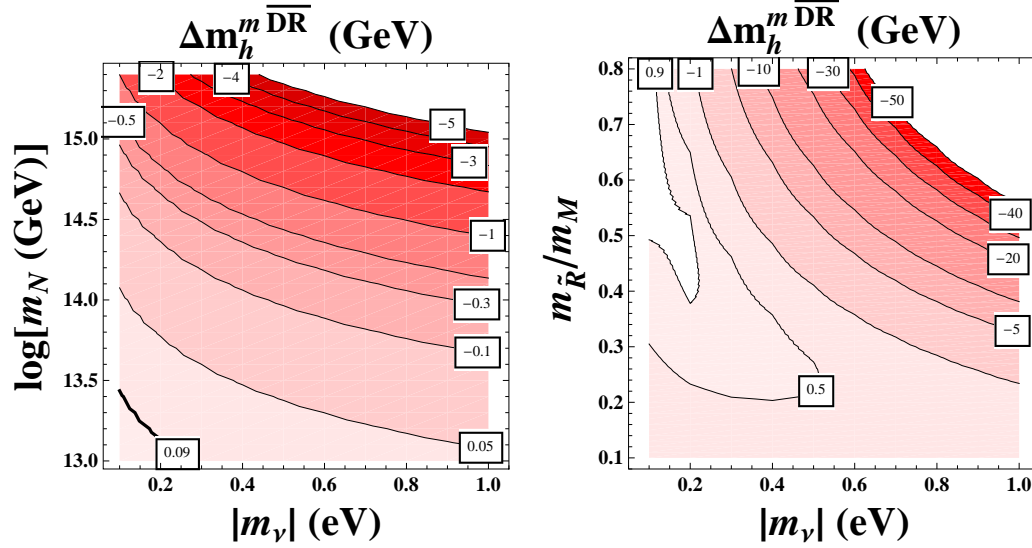


Figure 2: Contour-lines for the Higgs mass corrections from the $\nu/\bar{\nu}$ sector as a function of: $|m_\nu|$ and m_N (left panel) and the ratio $m_{\tilde{R}}/m_M$ and $|m_\nu|$ (right panel). The other parameters are: $A_\nu = B_\nu = m_{\tilde{L}} = m_{\tilde{R}} = 10^3$ GeV, $\tan\beta = 5$, $M_A = 200$ GeV, $\mu = 200$ GeV.

50 MeV (and also larger than the anticipated LHC precision of ~ 200 MeV). Consequently, they should be included in any phenomenological analysis of the Higgs sector in the MSSM-seesaw.

References

- [1] M. Frank, T. Hahn, S. Heinemeyer, W. Hollik, H. Rzehak and G. Weiglein, *JHEP* **0702** (2007) 047 [arXiv:0611326 [hep-ph]].
- [2] P. Minkowski, *Phys. Lett. B* **67** (1977) 421;
- [3] F. Borzumati and A. Masiero, *Phys. Rev. Lett.* **57** (1986) 961;
M. Raidal *et al.*, *Eur. Phys. J. C* **57**, 13 (2008) [arXiv:0801.1826 [hep-ph]].
- [4] S. Heinemeyer, M. J. Herrero, S. Penaranda and A. M. Rodriguez-Sanchez, *JHEP* **1105** (2011) 063 [arXiv:1007.5512 [hep-ph]].
- [5] J. Collins, F. Wilczek and A. Zee, *Phys. Rev. D* **18** (1978) 242.
- [6] J. Küblbeck, M. Böhm and A. Denner, *Comput. Phys. Commun.* **60** (1990) 165;
T. Hahn, *Comput. Phys. Commun.* **140** (2001) 418 [arXiv:hep-ph/0012260];
T. Hahn and C. Schappacher, *Comput. Phys. Commun.* **143** (2002) 54 [arXiv:hep-ph/0105349];
T. Hahn and M. Pérez-Victoria, *Comput. Phys. Commun.* **118** (1999) 153 [arXiv:hep-ph/9807565].
- [7] S. Heinemeyer, W. Hollik and G. Weiglein, *Comput. Phys. Commun.* **124** (2000) 76 [arXiv:hep-ph/9812320];
T. Hahn, S. Heinemeyer, W. Hollik, H. Rzehak and G. Weiglein, *Comput. Phys. Commun.* **180** (2009) 1426;
see: www.feynhiggs.de.