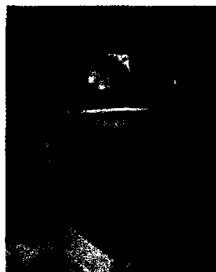


LESS-DIMENSIONS AND THE ORIGIN OF DARK MATTER

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We extend the concept of matter parity $P_M = (-1)^{3(B-L)}$ to non-supersymmetric theories and argue that P_M is the unique explanation to the existence of Dark Matter of the Universe. The argument is general but we motivate it using constraints on GUT particle content from lower-dimensional field theories. The non-supersymmetric Dark Matter must be contained in scalar **16** representation(s) of $SO(10)$, thus the unique low energy Dark Matter candidates are P_M -odd complex scalar singlet(s) S and inert scalar doublet(s) H_2 . We have calculated the thermal relic Dark Matter abundance of the model and shown that its minimal form may be testable at LHC via the SM Higgs boson decays $H_1 \rightarrow DM\ DM$. The PAMELA anomaly can be explained with the decays $DM \rightarrow \nu l W$ induced via seesaw-like operator which is additionally suppressed by Planck scale. Because the SM fermions are odd under matter parity too, Dark Matter sector is just our scalar relative.

1 Introduction

While the existence of Dark Matter (DM) of the Universe is now established beyond doubt¹, its origin, nature and properties remain obscured. In most models beyond the standard model (SM), such as the minimal supersymmetric SM, additional discrete Z_2 symmetry is imposed by hand to ensure the stability of the lightest Z_2 -odd particle. There is no known general principle for the origin of DM which could discriminate between DM models.

In this Letter we propose that there actually might exist such a common physics principle for the theories of DM. It follows from the underlying unified symmetry group for all matter fields in grand unified theories (GUT's) and does not require supersymmetry. One can classify all matter fields in Nature under the discrete remnant of the matter symmetry group which is nothing but the matter parity P_M . Thus the existence of DM might be a general property of Nature rather than an accidental outcome of some particular model. As a general result, there

is no “dark world” decoupled from us, rather we are part of it as the SM fermions are also odd under the matter parity P_M .

We argue that, assuming all matter fields to respect $SO(10)$ ², the gauged GUT group $SU(5)$ ³ is complemented by an additional discrete symmetry Z_n . For the simplest case, $n = 2$, the GUT symmetry is broken to $SU(5) \times P_M$ and all the fermion and scalar fields of the GUT theory, including the SM particles plus the right-handed neutrinos N_i , carry well defined discrete quantum numbers uniquely determined by their original representation of $SO(10)$. Therefore non-supersymmetric DM candidates can come only from **16** scalar representations of $SO(10)$, and the unique low energy DM fields are new $SU(2)_L \times U(1)_Y$ P_M -odd scalar doublet(s) H_2 ⁴ and singlet(s) S ^{5,6}.

While our argument is general, we motivate it with an example new physics scenario called less-dimensions. In the less-dimensional scenario new physics effects arise from lower-dimensional quantum field theories (QFT) as opposed to the extra dimensions in which new space dimensions are added. The consistency of field theory in three dimensions implies constraints on the number of fermions and gauge bosons of the theory. The 3-dimensional constraints apply also in 4-dimensions⁷ if one space dimension is very small and compactified to a circle, and the number of fermion generations is odd. During inflation the compactified dimension can be expanded as much as needed in order to achieve the present flat and homogeneous Universe. However, particle physics “remembers” the initial conditions. It is interesting that the WMAP observations indeed point to a preferred direction in space^{8,1}.

We formulate and study the minimal matter parity induced phenomenological DM model which contains one inert doublet H_2 and one complex singlet S . We show that the observed DM thermal freeze-out abundance can be achieved for wide range of model parameters. We also show that the PAMELA⁹ and ATIC¹⁰ anomalies in $e^+/(e^- + e^+)$ and $e^- + e^+$ cosmic ray fluxes can be explained by DM decays to the SM leptons via Planck scale suppressed P_M -violating seesaw-like operator of the form $m/(\Lambda_N M_P) LL H_1 H_2$, where m/M_P is P_M -violating heavy neutrino mixing. In this model the SM Higgs boson H_1 is the portal¹¹ to the DM. We show that for well motivated model parameter values the DM abundance predicts the decay $H_1 \rightarrow DM DM$, which allows to test the model at LHC¹².

2 Less-dimensions

To simplify the presentation we first discuss our example scenario and then present the general argument. If the topology of our 4-dimensional space-time is actually $\mathcal{M}^3 \times S^1$. *i.e.*, the usual Minkowski space-time with one spatial coordinate compactified to a circle, topological anomalies occur in non-Abelian gauge theories with *odd* number of massless fermion generations in a direct analogy with the corresponding theories in three dimensions⁷. Therefore the consistency of QFT in four dimensions must follow from the consistency of 3-dimensional QFT. In three dimensions there occur topological Chern-Simons terms in total non-Abelian gauge action as well as in the gravitational action which must be quantized¹³. At one loop level, corrections to the gravitational Chern-Simons term require¹⁴

$$\frac{1}{16}N_F - \frac{1}{8}N_G = 0, \quad (1)$$

where $N_{F,G}$ is the number of fermions and gauge bosons, respectively. Eq. (1) implies constraints on the number of fermions as well as on the gauge group of the theory.

At early Universe at very high energies the $\mathcal{M}^3 \times S^1$ space-time topology leads to CPT violation and generates a topological mass for the photon. However, today, after inflation, the

space is almost flat and homogeneous. The Chern-Simons-like photon mass is estimated to be⁷

$$m_\gamma \sim 10^{-35} \text{ eV} \left(\frac{\alpha}{1/137} \right) \left(\frac{1.5 \cdot 10^{10} \text{ yr}}{R} \right), \quad (2)$$

where R is the radius of the compact spatial coordinate. For the observable Universe, R_{obs} , m_γ is a factor $\mathcal{O}(10^2)$ smaller than the present experimental bound. Since inflation can generate $R \gg R_{\text{obs}}$, the topology of the Universe may remain unknown. Nevertheless, the WMAP results may support this scenario¹.

3 Matter parity as the origin of DM

An immediate consequence of Eq. (1) is that chiral fermionic matter must come in multiples of sixteen. This is in a perfect agreement with experimental data as there exist 15 SM fermions plus right-handed N for the seesaw mechanism¹⁵, and fermions of every generation naturally form one $SO(10)$ multiplet $\mathbf{16}_i$, $i = 1, 2, 3$. As a result, Eq. (1) implies that there must be 24 gauge bosons. This is the dimension of adjoint representation of $SU(5)$ and suggests that $SU(5)$ is the gauge group of GUT. In that case, if all matter fields, fermions and scalars, respect $SO(10)$, the group theoretic branching rule,

$$SO(10) \rightarrow SU(5) \times U(1)_X \rightarrow SU(5) \times Z_2, \quad (3)$$

implies that every $SU(5)$ GUT matter multiplet carries an additional uniquely defined quantum number under the (global or gauged) $U(1)_X$ symmetry. The $U(1)_X$ symmetry is further broken to its subgroup Z_n by order parameter carrying n charges of X ¹⁶. The simplest case Z_2 , which also allows for the seesaw mechanism induced by heavy neutrinos N_i ¹⁵, yields the new parity P_X with the field transformation $\Phi \rightarrow \pm \Phi$. Therefore the actual GUT group is $SU(5) \times P_X$.

Of course, the group theory in Eq. (3) is general and does not necessarily require less-dimensions nor global $U(1)_X$ because the Z_2 can also be gauged^{16,17}. We do not speculate on details of GUT model building here and adopt the breaking chain (3).

Under Pati-Salam charges $B - L$ and T_{3R} the X -charge is decomposed as

$$X = 3(B - L) + 4T_{3R}, \quad (4)$$

while the orthogonal combination, the SM hypercharge Y , is gauged in $SU(5)$. Because X depends on $4T_{3R}$ which is always an even integer for $T_{3R} = 1/2, 1, \dots$, the Z_2 X -parity of a multiplet is determined by $3(B - L) \bmod 2$. Therefore one can write

$$P_X = P_M = (-1)^{3(B-L)}, \quad (5)$$

and identify P_X with the well known matter parity¹⁸, which is equivalent to R -parity in supersymmetry. While $U(1)_X$, $X = 5(B - L) - 2Y$, has been used to consider and to forbid proton decay operators¹⁹, so far the parity (5) has been associated only with SUSY phenomenology.

Due to Eq. (3) a definite matter parity P_M is the general intrinsic property of every matter multiplet. The decomposition of $\mathbf{16}$ of $SO(10)$ under (3) is $\mathbf{16} = \mathbf{1}^{16}(5) + \bar{\mathbf{5}}^{16}(-3) + \mathbf{10}^{16}(1)$, where the $U(1)_X$ quantum numbers of the $SU(5)$ fields are given in brackets. This implies that under the matter parity all the fields $\mathbf{10}^{16}$, $\bar{\mathbf{5}}^{16}$, $\mathbf{1}^{16}$ are odd. At the same time, all other fields coming from small $SO(10)$ representations, $\mathbf{10}$, $\mathbf{45}$, $\mathbf{54}$, $\mathbf{120}$ and $\mathbf{126}$, are predicted to be even under P_M . Thus the SM fermions belonging to $\mathbf{16}_i$ are all P_M -odd while the SM Higgs boson doublet is P_M -even because it is embedded into $\mathbf{5}^{10}$ and/or $\bar{\mathbf{5}}^{10}$, and $\mathbf{10} = \mathbf{5}^{10}(-2) + \bar{\mathbf{5}}^{10}(2)$. Although $B - L$ is broken in nature by heavy neutrino Majorana masses, $(-1)^{3(B-L)}$ is respected by interactions of *all* matter fields.

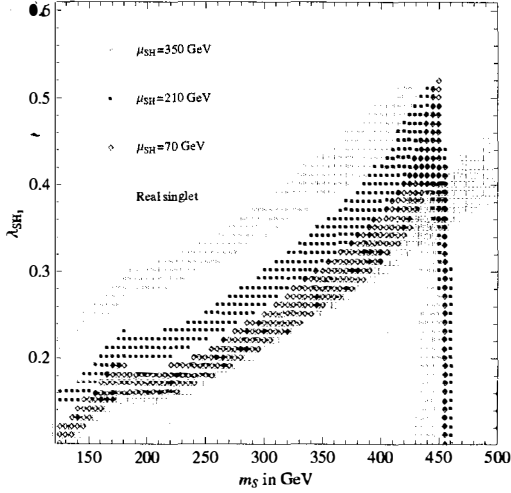


Figure 1: Allowed 3σ regions for predominantly singlet DM in (m_S, λ_{SH_1}) plane for $b_S = 5$ GeV, $m_{H_0} = 450$ GeV.

As there is no DM candidate in the SM, we have to extend the particle content of the model by adding new $SO(10)$ multiplets. The choice is *unique*, only **16** contains P_M -odd particles. Adding new fermion **16** is equivalent to adding a new generation, and this does not give DM. Thus we have only one possibility, the scalar(s) **16** of $SO(10)$. Because DM must be electrically neutral, **16** contains only two DM candidates. Under $SU(2)_L \times U(1)_Y$ those are the complex singlet $S = \mathbf{1}^{16}$ and the inert doublet $H_2 \in \mathbf{\bar{5}}^{16}$.

4 DM predictions of the minimal model

GUT symmetry groups are known to be very useful for classification of particle quantum numbers, and this is sufficient for predicting the DM candidates. Unfortunately GUTs fail, at least in their minimal form, to predict correctly coupling constants between matter fields. Therefore we cannot trust GUT model building for predicting details of DM phenomenology. Instead we study *phenomenological low-energy* Lagrangian for the SM Higgs H_1 and the P_M -odd scalars S and H_2 ,

$$\begin{aligned}
V = & -\mu_1^2 H_1^\dagger H_1 + \lambda_1 (H_1^\dagger H_1)^2 + \mu_S^2 S^\dagger S + \lambda_S (S^\dagger S)^2 \\
& + \lambda_{SH_1} (S^\dagger S) (H_1^\dagger H_1) + \mu_2^2 H_2^\dagger H_2 + \lambda_2 (H_2^\dagger H_2)^2 \\
& + \lambda_3 (H_1^\dagger H_1) (H_2^\dagger H_2) + \lambda_4 (H_1^\dagger H_2) (H_2^\dagger H_1) \\
& + \frac{\lambda_5}{2} [(H_1^\dagger H_2)^2 + (H_2^\dagger H_1)^2] + \frac{b_S^2}{2} [S^2 + (S^\dagger)^2] \\
& + \lambda_{SH_2} (S^\dagger S) (H_2^\dagger H_2) + \frac{\mu_{SH}}{2} [S^\dagger H_1^\dagger H_2 + S H_2^\dagger H_1],
\end{aligned} \tag{6}$$

which respects $H_1 \rightarrow H_1$ and $S \rightarrow -S, H_2 \rightarrow -H_2$. The doublet terms alone form the inert doublet model⁴. To ensure $\langle S \rangle = 0$, we allow only the soft mass terms b_S, μ_{SH} and the λ_5 term to break the internal $U(1)$ of the odd scalars⁶. Thus the singlet terms in (6) alone form the model A2 of⁶. The two models mix via λ_{SH}, μ_{SH} terms. Notice that mass-degenerate scalars are strongly constrained as DM candidates by direct searches for DM. The λ_5, b_S^2 and μ_{SH} terms in Eq. (6) are crucial for lifting the mass degeneracies.

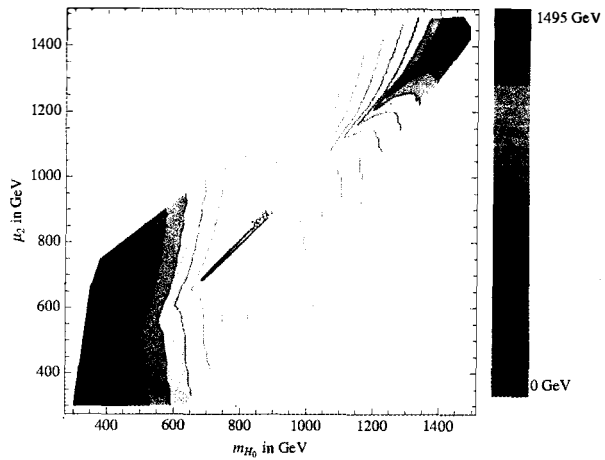


Figure 2: Allowed (m_{H_0}, μ_2) parameter space for $\mu_{SH} = 0$ and different values of m_S represented by color code.

In the following we assume that DM is a thermal relic and calculate its abundance using MicrOMEGAs package²⁰. The DM interactions (6) were calculated using FeynRules package²¹. To present numerical examples we fix the doublet parameters following Ref.²² as $m_{A_0} - m_{H_0} = 10$ GeV, $m_{H^+} - m_{H_0} = 50$ GeV and treat m_{H_0} and μ_2 as free parameters. For predominantly singlet DM we present in Fig. 1 the allowed 3σ regions in the $m_S^2 := \mu_S^2 + \lambda_{SH_1} v^2/2 - b_S^2$ and λ_{SH_1} plane for $b_S = 5$ GeV, $m_{H_0} = 450$ GeV and the values of μ_{SH} as indicated in the figure. For comparison we also plot the corresponding prediction of the real scalar model (light green band). For those parameters the observed DM abundance can be obtained for $m_S < m_{H_0}$. Due to the mixing parameter μ_{SH} , large region in the (m_S, λ_{SH_1}) plane becomes viable.

To study DM dependence on doublet parameters we present in Fig. 2 the (m_{H_0}, μ_2) parameter space for which the observed DM abundance can be obtained. Values of the singlet mass are presented by the colour code and we take $\mu_{SH} = 0$, $b_S = 5$ GeV. Without singlet S , in the inert doublet model²², the allowed parameter space is the narrow region on the diagonal of Fig. 2 starting at $m_{H_0} \approx 670$ GeV. In our model much larger parameter space becomes available.

5 PAMELA, ATIC and FERMI data

PAMELA satellite has observed a steep rise of $e^+/(e^- + e^+)$ cosmic ray flux with energy and no excess in \bar{p}/p ratio⁹. ATIC experiment claims a peak in $e^- + e^+$ cosmic ray flux around 700 GeV¹⁰, a claim that will be checked by FERMI satellite soon. To explain the cosmic e^+ excess with annihilating DM requires enhancement of the annihilation cross section by a factor 10^{3-4} compared to what is predicted for a thermal relic. Non-observation of photons associated with annihilation²³ and the absence of hadronic annihilation modes²⁴ constrains this scenario very strongly. However, the PAMELA anomaly can also be explained with decaying thermal relic DM with lifetime 10^{26} s²⁵, 3-body decays in our case.

In our scenario the global Z_2 matter parity can be broken by Planck scale effects¹⁶. If there exists, at Planck scale, a $SO(10)$ fermion singlet N' , its mixing with the $SU(5)$ P_M -odd singlet neutrinos N via a mass term $m_N N N'$ breaks P_M explicitly but softly. The exchange of N now induces also a seesaw-like¹⁵ operator

$$\frac{\lambda_N}{M_N} \frac{m}{M_P} LLH_1 H_2 \rightarrow 10^{-30} \text{GeV}^{-1} \nu l^- W^+ H_2^0, \quad (7)$$

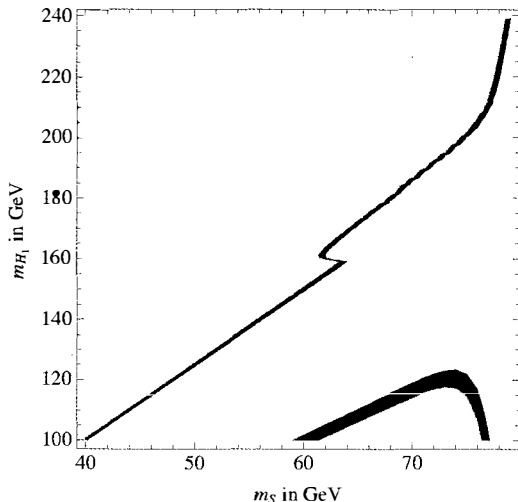


Figure 3: Allowed 3σ regions in the singlet DM and SM Higgs boson mass plane for $\mu_S = 0$ and $b_S = 5$ GeV.

where we have taken $\lambda_N \sim 1$, $M_N \sim 10^{14}$ GeV and $m \sim v \sim 100$ GeV. Such a small effective Yukawa coupling explains the long DM lifetime 10^{26} s.

5 LHC phenomenology

In our scenario the DM couples to the SM only via the Higgs boson couplings Eq. (6). Therefore, discovering ~ 1 TeV DM particles at LHC is very challenging. However, if DM is relatively light the SM Higgs decays $H_1 \rightarrow DM DM$ become kinematically allowed and the SM Higgs branching ratios are strongly affected. Such a scenario has been studied by LHC experiments¹² and can be used to discover light scalars.

In our model such a scenario is realized for $\mu_S = 0$, small $b_S \ll v$ and heavy doublet. In this case the DM is predominantly split singlet and, in addition, the DM abundance relates the DM mass $m_S^2 \approx \lambda_{SH_1} v^2/2 - b_S^2$ to the SM Higgs boson mass m_{H_1} , as seen in Fig. 3. For $m_{H_1} = 120$ GeV, $b_S = 5$ GeV we predict $m_S = 48$ GeV with the Higgs branching ratios $BR(H_1 \rightarrow b\bar{b} + c\bar{c} + \tau\bar{\tau}) = 14.2\%$, $BR(H_1 \rightarrow DM DM) = 42.4\%$ and $BR(H_1 \rightarrow S_2 S_2) = 42.4\%$. The second heaviest singlet S_2 with the mass $m_{S_2}^2 \approx \lambda_{SH_1} v^2/2 + b_S^2$ decays via the SM Higgs exchange to $S_2 \rightarrow DM \mu \mu$ or $S_2 \rightarrow DM c \bar{c}$ with almost equal branching ratios. Thus the SM Higgs boson decay modes are very strongly modified. This makes the H_1 discovery more difficult at LHC but, on the other hand, allows the scenario to be tested via the Higgs portal¹¹.

6 Conclusions

We have extended the concept of Z_2 matter parity, $P_M = (-1)^{3(B-L)}$, to non-supersymmetric GUTs and argued that P_M is the unique origin of DM of the Universe. Assuming that $SO(10)$ is the matter symmetry group, the matter parity of all $SU(5)$ GUT matter multiplets is determined by their $U(1)_X$ charge under Eq. (3). We have motivated this scenario with the constraint Eq. (1) from lower-dimensional effective field theories but our argument is general. Consequently, the non-supersymmetric DM must be contained in the scalar representation **16** of $SO(10)$. This implies that the theory of DM becomes completely predictive and the only possible low energy DM candidates are the P_M -odd scalar singlet(s) S and doublet(s) H_2 . We have calculated the

DM abundances in the minimal DM model and shown that it has a chance to be tested at LHC via Higgs portal. Planck-suppressed P_M breaking effects may occur in the heavy neutrino sector leading to decays $DM \rightarrow \nu l W$ which can explain the PAMELA and FERMI anomalies.

Our main conclusion is that there is nothing unusual in the DM which is just scalar relative of the SM fermionic matter. Although $B - L$ is broken in Nature by heavy neutrino Majorana masses, $(-1)^{3(B-L)}$ is respected by interactions of all matter fields implying stable scalar DM.

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