

# Using Machine learning to segregate cosmic muons from background events

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## Introduction

Cosmic muons are high-energy charged particles originating from cosmic rays in the upper atmosphere and reaches the Earth surface. These cosmic muons has the mean energy of  $\sim 3\text{-}4$  GeV, which make them a particularly useful probe for applications like Muon Tomography, and can also be used as a probe to calibrate particle detectors. In addition to cosmic muons, several order of magnitude more cosmic background is also available in our environment. Hence, these cosmic muons need to be segregated from background events depending upon the application specific muon trigger condition. In the current work we will present a Machine Learning based approach to segregate these cosmic muon from background gamma where the muon trigger is the coincidence between the two layers of plastic scintillator detectors separated by  $\sim 1\text{m}$ .

## Detector Setup

The detector setup consist of plastic scintillator matrix of ISMRAN detector [1]. It consist of 90 scintillator bars arranged in a matrix of  $9 \times 10$ . Each plastic scintillator bar ( $10\text{ cm} \times 10\text{ cm} \times 100\text{ cm}$ ) is coupled to the Photo Multiplier Tube (PMT) at both ends, which are used to get the timing and the integrated charge. The geometric mean of these integrated charges correspond to the energy deposited in the scintillator by the incident radiation, and is key to segregate the cosmic muon from background events in conventional techniques.

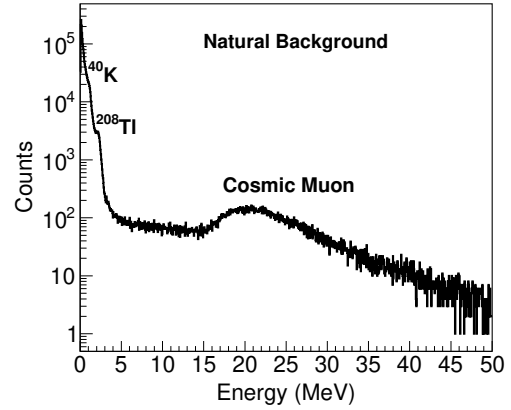


FIG. 1: Spectrum of energy calibrated background  $\gamma$  and cosmic muons

## Conventional energy deposition based cosmic muon segregation

Conventionally cosmic muons are segregated from the background based on the energy deposition threshold [2], which involves the energy calibration of these detectors. Fig. 1 shows the plot of energy deposited by cosmic particles after the application of energy calibration. The energy deposited by cosmic muons peaked at  $\sim 20\text{MeV}$ , which is also verified using GEANT4 [3] simulation. Hence, an appropriate energy threshold  $\sim 8\text{-}10$  MeV may be chosen to segregate the cosmic muons hits from the background. To further make sure that the selected events are muon events, a trigger may be generated by taking the coincidence of two layers of scintillator separated by a distance of  $\sim 1\text{m}$ . Hence an event will be considered as muon event only if crosses the energy deposition threshold and forms a muon trigger.

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## Machine learning based cosmic muon segregation

The conventional approach needs energy calibration procedure followed by muon trigger generation check for individual event to be declared as muon event. This makes the job computation intensive as background events are orders of magnitude more than muon events. Machine learning can help solving the problem where the muon segregation problem can be seen as the binary classification problem, which can be solved by building a classification model. The built model can later be used to quickly discriminate background and muon events without applying any energy calibration or trigger generation. Calibrated data and trigger is required only in the initial phase to label the individual events as muon event or background event, which is required for training the classification model.

### Data Sets

As mentioned above, the data from experiment consist of timing and integrated charge. In the present work, both the parameters are used as the features for the segregation of cosmic muon. For the learning phase of the model, the timing and integrated charge obtained from both the PMTs are used as the input features. The minimum energy deposition threshold is kept at 3 MeV, so that the muons which are just grazed through the detector should also get recorded. The event is labelled as muon event if the energy deposited is more than the threshold value and a coincidence is found between the top and the bottom layer, which are separated by  $\sim 1\text{m}$ , otherwise it is labelled as background event.

### Results

After building the labelled dataset, a regression based classification model is trained using  $3 \times 10^5$  events of labelled data. To evaluate the performance of the classification model, another labelled dataset of  $3 \times 10^5$  events is used, and muon classification performance of the model is found to be  $\sim 95\%$ . Later on the same model is applied on the input features

of raw data without applying any energy cali-

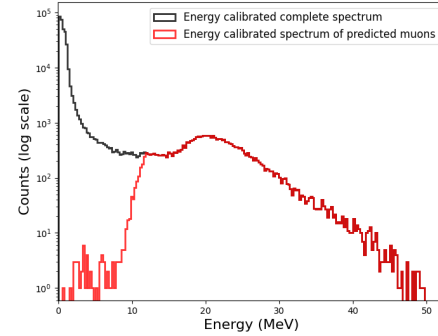


FIG. 2: Muons segregation using Machine learning classification algorithm

bration. Once the classification is over, the energy calibration is applied only on those events which are classified as muons, and the data is compared with the calibrated data of the complete spectrum, as show in Fig. 2. Here one can see in the red plot, that there are very few low energy background events which are classified as muons, but from  $\sim 10\text{-}11$  MeV onwards we are able to correctly classify them as muons, which we had previously verified by GEANT4 simulations.

### Conclusion

By harnessing the power of advanced machine learning classification algorithms, a successful segregation between cosmic muons and background events is obtained based on their distinct energy deposition and timing patterns. This has yielded the promising results with significant implications for applications like Muon Tomography.

### References

- [1] R. Dey et al., *NIMA* **1042**, **2022**
- [2] R. Sehgal et al., *JINST* **17** P02036, September **2022**
- [3] S. Agostinelli et al., *GEANT4*, *NIM-A* **506**, 250, **2003**